

ASTRONOMICAL HANDBOOK FOR SOUTHERN AFRICA

1976

This booklet is intended both as an introduction to observational astronomy for the interested layman - even if his interest is only a passing one - and as a handbook for the established amateur or professional astronomer.

Front Cover: The multiple photographic telescope array belonging to Mr. C. Papadopoulos.

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NOTE

Unless stated otherwise, all times are SOUTH AFRICAN STANDARD TIME. In order to avoid confusion between a.m. and p.m., the 24-hour clock is used (e.g. 1800 hrs. is 6 p.m. and 2100 hrs. is 9 p.m.). Emphasis is given to phenomena visible in the evening sky - between sunset and midnight.

This handbook is produced for the Astronomical Society of Southern Africa. Most of the data it contains have been adapted for Southern Africa from the "Astronomical Ephememeris for 1976" issued jointly by the Nautical Almanac Offices of the Royal Greenwich Observatory, Herstmonceux and the U.S. Naval Observatory, Washington D.C. Additional information has also been supplied direct from the Herstmonceux Office and from the Hydrographer, South African Navy.

Thanks are due to Dr A.P. Fairall for his invaluable assistance and advice and to Messrs G and C Larmuth for the preparation of much of the data, and to the Printing Department of the University of Cape Town.

All correspondence concerning this booklet should be addressed to the Handbook Editor, Astronomical Society of Southern Africa, 8 Glebe Road, Rondebosch 7700, Cape. Further copies can also be obtained from the same address. Enclose a postal order or cheque (in favour of the Astronomical Society of Southern Africa) for R1 per copy.

Although every care has been taken in the compilation of the Handbook, it is distributed and sold on the explicit that neither the Astronomical Society of Southern Africa nor any of its members accepts any responsibility for errors.

Dit is met spyt dat ons u meedeel dat as gevolg van beperkte fondse en produksiefasiliteite dit nie moontlik was om hierdie handboek in Afrikaans te laat druk nie.

> R.F. HURLY EDITOR

ASTRONOMY IN SOUTHERN AFRICA

As one of the few parts of the Earth having both access to the rich southern skies and a suitable climate, Southern Africa holds a favoured position in astronomy. Consequently it has seen the establishment of a number of professional observatories engaged in research while many individuals have become enthusiastic amateur astronomers. Planetaria and visiting nights at observatories convey to the general public much of what goes on in this field.

OBSERVATORIES

Boyden Observatory, situated at Mazelspoort, just outside Bloemfontein, is operated by an international consortium representing Belgian, Irish and South African interests. Thus many astronomers come from overseas to make use of its observing facilities which include the 1,5 m Rockefelter Reflector and the 0,9 m ADH Baker Schmidt. Its site offers good oberving conditions, without being remote from a large centre.

On the other hand, the observatory sites in the hearts of Cape Town and Johannesburg have become unsuitable with the rapid expansion of those cities. The South African Astronomical Observatory a joint venture between the South African Council for Scientific and Industrial Research and the British Science Council-has merged their facilities and moved the larger instruments to a new site near Sutherland in the Karroo. Observing continues at Cape Town which is also the Headquarters of the S. A. A. O. The S. A. A. O. outstation at Hartebeespoort continues to share the site with the Leiden Observatory Southern Station which has a 0,9 m. "light collector" reflector.

The 1,9m Radcliffe reflector formerly near Pretoria and the largest telescope in Southern Africa has now been erected in Sutherland.

In the field of radio astronomy, the 25 m dish of the Deep Space Tracking Station near Krugersdorp is used for research work when not required for tracking spacecraft, while the Rhodes University Radio Observatory just outside Grahamstown, has a number of arrays for receiving radio emission from the planet Jupiter.

In addition to the professional observatories listed above, South Africa and Rhodesia have numerous private observatories, built and operated by amateur astronomers.

OBSERVATORIES OPEN TO THE PUBLIC

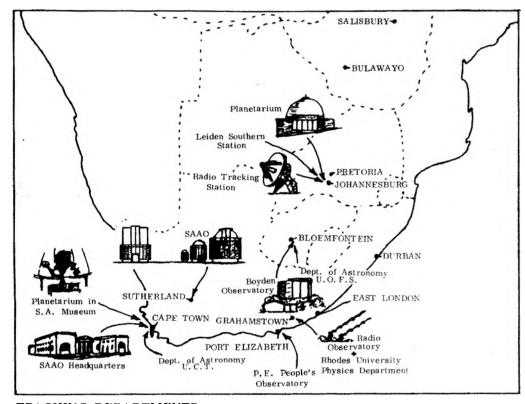
Visiting nights at Boyden Observatory are held twice per month usually around the time of first quarter. Intending visitors should contact the Information Office in Hoffman Square, Bloemfontein for tickets (gratis). Numbers are restricted to twenty persons on each visiting night.

SAAO Headquarters, Observatory, Cape are open to visitors on the second Saturday each month at 8.00 p.m. No tickets are necessary but parties of more than ten persons should contact the observatory in advance.

PLANETARIA

The major planetarium in South Africa is that situated on the grounds of the University of the Witwatersrand (entrance in Yale Road - alongside MI). It is equipped with a highly complex Zeiss projector and seats over 400 persons.

A smaller planetarium, with a Spitz projector and seating approximately 70, is located within the South African Museum, Cape Town. Shows are given each Saturday at 3.00 p.m., each Sunday at 3.30 p.m., and at 11.00 a.m. and 3.30 p.m. on public and school holidays. Further information can be obtained by phoning the museum at 41-2668.



TEACHING DEPARTMENTS

Both the University of the Orange Free State and the University of Cape Town have departments of astronomy - concerned with optical astronomy - while the Physics department of Rhodes University specialises in radio astronomy. The UOFS department is incorporated into Boyden Observatory and is headed by the director, Prof. A.H. Jarrett. Prof. Brian Warner occupies the chair of Astronomy at UCT. His department makes use of the SAAO observing facilities at Sutherland. The Physics Department at Rhodes has its own radio observatory outside Grahamstown.

THE ASTRONOMICAL SOCIETY OF SOUTHERN AFRICA

The Astronomical Society of Southern Africa is a body consisting of both amateur and professional astronomers. Membership is open to all interested persons, regardless of knowledge or experience. In addition to this handbook, the Society issues "The Monthly Notes of the Astronomical Society of Southern Africa" (MNASSA). Members also receive copies of "Sky and Telescope", an excellent and very popular monthly magazine published in the United States. It provides up to date information on both professional and amateur activities, together with news of space research and other related subjects. The Society's annual subscription is R8.00 and there is an entrance fee of R2.50. Information can be obtained from the Honorary Secretary, Astronomical Society of Southern Africa, c/o The South African Astronomical Observatory, P.O. Box 9, Observatory 7935, Cape.

LOCAL CENTRES OF THE SOCIETY

Autonomous local Centres of the Society hold regular meetings in Cape Town, Bloemfontein, Durban, Johannesburg, Pietermaritzburg and Pretoria. Visitors are very welcome at meetings and may, if they wish, join a Centre, without becoming a full member (i.e. receiving publications for R8 subscription) of the Society.

CAPE CENTRE (Cape Town) - Meetings on 2nd Wednesday of the month (except Jan. and Dec.) at the South African Astronomical Observatory at 8.00 p.m. The Centre possesses a small observatory housing the twelve inch Ron Atkins Telescope. There is also an active occultation section. Secretarial address: 4 Chalfont Road, Newlands 7700. Information on meetings also available from telephone (day time) 69-8531 ext. 256 (evenings) 61-5300.

TRANSVAAL CENTRE (Johannesburg) - Alternate lecture and observing meetings are held each month. There is a very enthusiastic lunar occultation group which observes grazing occultations within a radius of approximately 250 km of Johannesburg. Mirror making classes are held at suitable intervals and the centre has its own aluminising plant. Secretarial address: P.O. Box 5595, Johannesburg 2001.

NATAL CENTRE (Durban) - Monthly meetings are held at the Teacher's Centre. Occultation and telescope making sessions. The Centre has a portable twelve inch reflecting telescope. Secretarial address: Box 2704, Durban 4000. Telephone (evenings) 72-3187.

NATAL MIDLANDS CENTRE (Pietermaritzburg) - Meetings usually on the first Thursday of the month either in a private home. Secretarial address: 17 Yalta Rd., Pietermaritzburg 3201. Telephone 54038

FREE STATE CENTRE (Bloemfontein) - The Centre has its own observatory at Brandkop. Meetings on 2nd Thursday of the month at the homes of members. For further information, contact Mr. G.J. Muller, 35 Wilcocks Road, Bloemfontein 9301. Telephone (evenings) 7-3442 or Mr. J. Rhodes, Telephone 7-1981 (day time). Associate members should endeavour to have communications in the hands of the secretary a week before the monthly meetings.

PRETORIA CENTRE - The Centre has its own observatory containing a twelve inch reflecting telescope. For information contact Mr. K.J. Sterling, 5 Hekla Road, Valhalla 0137 - Phone 713272.

OBSERVING SECTIONS OF THE SOCIETY

These sections exist to coordinate and encourage constructive observing programmes. Mention of the type of observations and equipment involved are made in the appropriate parts of this handbook together with the names and addresses of the directors.

Comets and Meteors	see page 28
Grazing Occultations	see page 40
Nova Search Section	see page 33
Ordinary Occultations	see page 35
Variable stars	see page 34

DIARY 1976

-	13h				
-	F-13	Earth at perihelion	14 ^d	12^{h}	Spica 00.3 S of Moon
7	7	Mercury at greatest elongation E	15	8	Uranus 1° N of Moon
		(19^{0})	27	22	Jupiter in conjunction with the Sun
8	14	Venus 7º N of Antares	28	4	Mercury greatest elongation E 210
12	6	Venus 0 ⁰ .4 N of Neptune	29	12	Annular eclipse of the sun visible
13	4	Ceres 0°, 3 N of Moon (mag 7.6)			North Africa, Greece, Turkey,
14	5	Mars 50 N of Moon			Tibet
17	15	Saturn 5° N of Moon			
20	13	Saturn at opposition			
20	12	Sun overhead at Bulawayo	M.	AY	
23	9	Spica occulted			0
24	9	Uranus 20N of Moon	5	6	Mars 50S of Pollux
26	23	Neptune (mag 7,8) occulted	5	16	Mars 70 N of Moon
28	10	Venus 20 S of Moon	5	22	Saturn 60 N of Moon
30	12	Sun overhead at Salisbury	11	16	Venus 00.2 S of Jupiter
			11	22	Spica 0°, 3 S of Moon Mars 1°, 3 N of Saturn
E.E.	BRU	ADV	12	4	Mars 17, 3 N of Saturn
	DICO	VIV I	27	6	Jupiter 00.8 S of Moon
6	5	Jupiter 4°S of Moon			
10	18	Mars 5 ⁰ N of Moon			
13	21	Saturn 50N of Moon	311	NE	
16	17	Mercury greatest elongation W 26 ⁰			0.4 .0
17	0	Juno occulted	2	8	Saturn 60 N of Moon
19	15	Spica occulted	3	4	Mars 7º N of Moon
20	15	Uranus 10 N of Moon	8 9	7	Spica 0 ⁰ , 4 S of Moon Uranus 1 ⁰ N of Moon
23	6	Neptune 00, 9 S at Moon	15	11	Mercury greatest elongation W 23 ^c
27	16	Venus 6° S of Moon	21	- 8	Winter solstice
28	2	Mercury 7° S of Moon	22	19	Mercury 3 ⁰ N of Aldebaran
			24	1	Jupiter 0°.1 S of Moon
M	ARCE	1	26	0	Mercury 1 ^o N of Moon
	*****		29	20	Saturn 6 ⁰ N of Moon
4	22	Jupiter 3 ^o S of Moon Mars 6 N of Moon	29	20	Strain o Wat Moon
9	21	Mars 6 N of Moon			
12	5	Saturn 50 N of Moon	JU	LY	
18	1	Spica 0.3 S of Moon			
18	22	Uranus 1 ⁰ N of Moon	1	16	Mars 6° N of Moon
20	14	Equinox	3	6	Earth at aphelion
29	2	Venus 60 S of Moon	5	13	Spica 0°. 6 S of Moon
			5	20	Mars 0°.7 N of Regulus
AP	PRIL		6	6	Uranus 10 N of Moon
	10		8	23	Neptune 1° S of Moon
1	16	Jupiter 2 ⁰ S of Moon Mars 7 ⁰ N of Moon	21	19	Jupiter 0°.5 N of Moon
7	5 14	Mars 7 ⁻ N of Moon Saturn 6 ⁰ N of Moon	24	16	Mercury 0°.4 N of Venus
- 0			29	16	Saturn in conjunction with Sun
8 12	20	Mercury 10, 9 N of Jupiter	30	04	Mars 5 N of Moon

AUGUST

OCTOBER

1^{d}	18 ^h	Spica 0°.8 S of Moon	7^{d}	18 ^h	Mercury greatest elongation W 180
3	8	Mercury 0.º 7N of Regulus Venus 1.1 N of Regulus	12	3	Jupiter 10 N of Moon
7	18	Venus 1.1 N of Regulus	18	7	Saturn 6 ⁰ N of Moon
18	11	Jupiter 1 ⁰ N of Moon	19	0	Mars 0 ⁰ .4 S of Uranus
24	1	Saturn 6 ^O N of Moon	23	7	Total eclipse of the sun, seen in
26	12	Mercury greatest elongation E 270			Southern Africa as a partial eclipse
27	2	Venus 50 N of Moon	25	15	Venus 40 S of Moon
27	17	Mars 40 N of Moon	28	3	Venus 30 N of Antares
29	1	Spica 10 S of Moon	31	8	Venus 3 ⁰ S of Neptune
29	19	Uranus 0.6 N of Moon			

SEPTEMBER

NOVEMBER

100		··	7	1	Penumbrial eclipse of the Moon
6	6	Mercury 5 ^o S of Venus Venus 0.4 N of Mars	8	3	Jupiter 1 ⁰ N of Moon
11	0	Venus 0.4 N of Mars	14	17	Saturn 60 N of Moon
14	21	Jupiter 1 ⁰ N of Moon	18	10	Jupiter at opposition
20	3	Venus 3 ⁰ N of Spica	19	7	Spica 1 ⁰ S of Moon
20	17	Saturn 60 N of Moon	20	8	Uranus 0°, 05 S of Moon, Occultation
25	7	Mars 2 ⁰ N of Moon	24	15	Venus 7 ⁰ S of Moon
25	20	Venus 0.7 N of Moon	25	3	Mars in conjunction with the Sun
27	21	Mars 3 ⁰ N of Spica	25	17	Mercury 30 S of Neptune
30	24	Venus 0.0 5 S of Uranus			

DECEMBER

5d	2^{h}	Jupiter 0 ⁰ .8 North of Moon.
		Occultation
5	19	Neptune in Conjunction with the sun.
11	23	Saturn 60 N of Moon
16	16	Spica 1 S of Moon
20	12	Mercury greatest elongation East 20°
22	17	Mercury 6°S of Moon
24	17	Venus 7 ⁰ S of Moon

THE SUN 1975

BASIC DATA

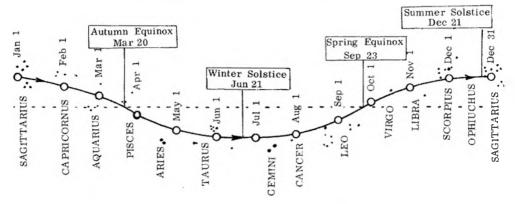
Diameter: 1 392 000 km (109 times Earth diameter) Mass: 1,99 x 10³⁰ kg (330 000 times Earth Mass)

Surface Temperature: Approx. 6000°C

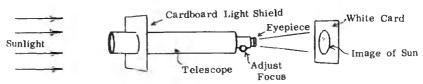
Temperature at centre: Approx. 10 million C

The Sun is our nearest star. It is composed chiefly of hydrogen and is in a gaseous state throughout. So hot and dense is its interior that nuclear reactions occur there - thus producing the energy that is eventually radiated from its surface. At times its surface is disturbed by sunspots (which may persist for some weeks) and flares (short lived).

The Earth's orbit around the Sun is not quite circular. In 1976 we will be closest to the Sun on January 4 (perihelion - approx. distance 147 million km.) and furthest from the Sun on July 3 (aphelion - approx. 152 million km.) During the year, the Sun appears to us to make a complete circuit of the sky (i.e. relative to the starry background) as indicated in the diagram.



Permanent damage to the eye can be caused by looking directly at the Sun. The diagram below shows how a small telescope (or half a binocular) may be used to project an image of the solar disk onto a piece of white card. It may also be advisable to stop down the telescope aperture so that the eyepiece is not damaged by the intense light passing through it. Tiny black sunspots are generally visible on the otherwise white solar disk - if monitored over a period of a week or so, the rotation of the Sun should be apparent.



The Sun 1976

TIMES OF SUNRISE AND SUNSET

		CAPE	TOW	/N		DUR	BAN		В	LOEMF	тио	EIN	JO	DHANN	ESBU	IRG	;	SALISE	URY	
	SU	NRISE	SUI	NSET	SU	NRISE	SU	NSET	SU	NRISE	803	NSET	SUI	NRISE		NSET	SUN	IR ISE	SUN	SET
Jan 1	os ^h	38 ^m	20 ^h	01 ^m	0.1	58 ^m	19 ^h	01 ^m	05 ^h	21 m	19 ^h	18 ^m	05 ^h	18 ^m	19 ^h	04 ^m	osh	24 ^m	18 ^h	35 ^m
11	05	46	20	02	05	06	19	02	05	29	19	18	05	25	19	05	05	29	18	37
21	05	55	19	59	05	14	19	00	05	37	19	17	05	33	19	04	05	37	18	38
Feb 1	06	07	19	52	05	24	18	55	05	46	19	13	05	42	19	00	05	42	18	36
11	06	17	19	44	05	32	18	48	05	54	19	06	05	49	18	55	05	47	18	32
21	06	26	19	33	05	41	18	39	06	02	18	57	05	56	18	47	.05	52	18	27
Mar 1	06	33	19	23	05	46	18	30	06	80	18	48	06	00	18	39	05	55	18	21
11	06	41	19	11	05	53	18	19	06	13	18	38	06	06	18	29	05	57	18	15
21	06	49	18	58	05	59	18	80	06	18	18	27	06	11	18	19	06	00	18	06
Apr 1	06	58	18	41	06	06	17	53	06	25	18	13	06	17	18	06	06	02	17	57
11	07	04	18	30	06	11	17	43	06	30	18	03	06	21	17	56	06	04	17	50
21	07	13	18	17	06	17	17	31	06	35	17	52	06	25	17	47	06	07	17	43
May 1	07	20	18	05	06	24	17	22	06	42	17	44	06	31	17	38	06	10	17	37
11	07	28	17	57	06	31	17	14	06	49	17	36	06	37	17	31	06	13	17	32
21	07	34	17	50	96	36	17	08	06	54	17	30	06	41	17	26	06	16	17	29
Jun 1	07	43	17	45	06	43	17	04	07	01	17	27	06	47	17	23	06	20	17	28
11	07	48	17	44	06	48	17	03	07	05	17	26	06	52	17	22	06	23	17	27
21 Jul 1	07	51	17	44	06	51	17	04	07	08	17	27	06 06	55 57	17	24	06	26	17	29
Jui 1	07	53	17	48	06	53	17	07	07	10	17	30	06	55	17	27 30	06	27	17	32
21	07	51	17	52	06	51	17	11	07	80	17	34		53	17		06	27	17	35
Aug 1	07	47 39	17	58 06	06	48	17	16	07	05	17	39	06 06	48	17	35	06	26	17	40
7.ug 1	07 07	30	18 18	13	06	42 34	17	22 29	07	00	17	45	06	41	17 17	41 46	06	23 18	17	42
21	07	19	18	20	06		17		06	53	17	51	06	32	17	50	06		17 17	46
Sep 1	07	06	18	27	06 06	24 12	17 17	35 40	06	42	17	55	06	21	17	54	06 06	11 04	17	48 49
лер 1 11	06	52	18	34	06	00	17	46	06 06	31 19	18	01	06	11	17	59	05	55	17	51
21	06	38	18	41	05	48	17	51	06	07	18 18	06	05	59	18	03	05	46	17	52
Oct 1	06	25	18	48	05	37	17	57	05	57	18	10 16	05	50	18	08	05	39	17	54
11	06	12	18	55	05	25	18	03	05	45	18	22	05	39	18	12	05	30	17	57
21	05	58	19	04	05	12	18	09	05	33	18	27	05	27	18	17	05	23	17	59
Nov 1	05	16	19	13	05	02	18	17	05	24	18	35	05	19	18	24	05	16	18	03
11	05	36	19	23	04	55	18	26	05	17	18	44	05	13	18	32	05	14	18	08
21	05	31	19	33	04	49	18	34	05	12	18	52	05	08	18	39	05	11	18	13
Dec 1	05	29	19	43	04	48	18	42	05	11	19	00	05	07	18	46	05	12	18	19
11	05	28	19	50	04	48	18	50	05	11	19	07	05	08	18	53	05	14	18	25
21	05	32	19	57	04	52	18	57	05	15	19	14	05	12	19	00	05	18	18	31
																	- 5			

SOLAR ECLIPSES

Annular eclipse of the Sun, April 29. This eclipse is not visible from Southern Africa.

Total eclipse of the Sun, October 23. From Southern Africa this eclipse will be seen as a partial eclipse in the early morning. Predictions are as follows:

		Cape T	own	;	Johannes	burg		Salisbu	<u>ry</u>	
Eclipse begins (P.A.)	5 h	37 ^m 05	(341°)	5 ^h	14 ^m 73	(325 ⁰)	4^{h}	58 ^m 98	(311 ⁰)	
Maximum eclipse (magnitude)	6	08,17	(19%)	5	57,49	(41%)	5	47,77	(63%)	
Eclipse ends (P. A.)	6	40,59	(56°)	6	43,28	(74°)	6	40,87	(88°)	

The position angle (P.A.) of the point of contact is measured eastwards from the north point of the Sun. The magnitude is the percentage of the Sun's Diameter obscured.

THE MOON 1976

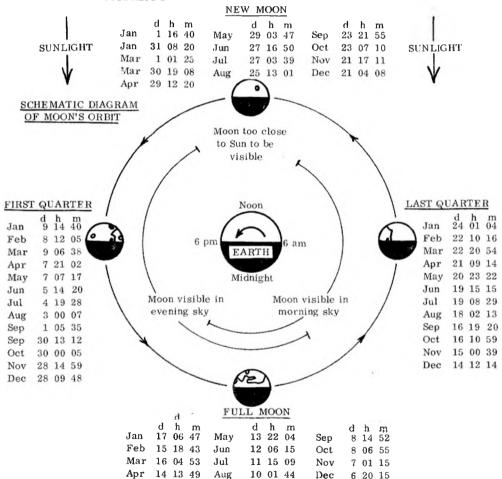
BASIC DATA

Diameter: 3 480 km (0,27 of Earth) Mass: 7,35 x 10^{22} kg (1/81 of Earth)

Surface Gravity: 0,16 of Earth

Average distance from Earth: 384 000 km

PHASES AND VISIBILITY



THE MOON'S ORBIT

Dates of Apogee, when the Moon is furthest from the Earth (approx. 407 000 km) and of Perigee, when the Moon is closest to the Earth (approx. 357 000 km) are given below.

		Apogee					Perigee		
Jan	8	May 25	Oct	10	Jan	20	Jun 9	Sep	25
Feb	5	Jun 21	Nov	6	Feb	17	Jul 7	Oct	23
Mar	4	Jul 19	Dec	3	Mar	16	Aug 1	Nov	21
Mar	31	Aug 16	Dec	31	Apr	14	Aug 28	Dec	19
Apr	27	Sep 13			May	12			

As a result of its motion around the Earth, the Moon appears to make a complete circuit of the heavens in just under a month. It occasionally passes in front of bright stars (details given in Occultation section - page 43) and close to visible planets (details given in Diary pages 4 and 5).

TIMES OF MOONRISE AND MOONSET

Times for Bloemfontein, Cape Town, Durban, Johannesburg and Port Elizabeth can be obtained from the tables on pages 11 to 16.

LUNAR ECLIPSES

Partial eclipse of the Moon, May 13. Details of this eclipse are as follows:

4	-	-1	1.	and a
Moon enters penumbra	May	13"	19"	46 ^m ,6
Moon enter umbra		13	21	15, 7 (at 1680 East of North point)
Middle of eclipse		13		54, 3 (12% of Moon's diameter obscured)
Moon leaves umbra		13	22	32, 9 (at 150° West of North point)
Moon leaves penumbra		14	00	02,0

Penumbral eclipse of the Moon, November 6-7. Details are as follows:

Moon enters penumbra	November 6 ^d	22 ^h	45, 6 (at 34° East of North point)
Middle of eclipse			01,1 (86% of Moon's diameter obscured)
Moon leaves penumbra	7	03	16,5 (at 57° West of North point)

THE SURFACE OF THE MOON

In common with the inner planets of our solar system, the Moon's surface suffered bombardment by numerous minor bodies during the period 4,5 to 3,0 billion years ago. This has produced the heavily cratered topography now visible. Some particularly large impacts caused large circular depressions, which were flooded by molten lava from the Moon's interior. These are the maria basins which appear smoother and darker then the rest of the surface (the latin words mare and maria come from older times when they were mistaken for seas). The maria surfaces, being younger have fewer large craters, but the entire surface is peppered with tiny craters produced by tiny bodies which have also served to plough up the ground thus forming the regolith - a layer of loose material a metre or so deep.



LIBRATIONS



Jan 5 Feb 1/28 Mar 26 Apr 22 May 19 Jun 16 Jul 13 Aug 9 Sep 5 Oct 2/30 Nov 26 Dec 23



Jan 19 Feb 15 Mar 13 Apr 10 May 7 Jun 3/30 Jul 27 Aug 24 Sep 20 Oct 17 Nov 13 Dec 11 Dates of Maximum Exposure of Indicated Limbs Jan 15 Feb 12 Mar 11 Apr 8 May 6 Jun 2/29 Jul 26 Aug 22 Sep 19 Oct 17 Nov 15 Dec 13

Jan 2/29 Feb 24 Mar 23 Apr 20 May 19 Jun 16 Jul 14 Aug 10 Sep 5 Oct 2/30 Nov 27 Dec 25



JOHANNESBURG - TIMES OF

		Jan				Leb	0			2	4	1
-	RISE 05h 04	SE O4m		ST 56m	R 106	RISE 6h 40m	SE 19h	$^{ m T}$	RISE 06h 19	- 8	SE 18h	T 45
N	90	03		42	0.7		20	10	07		19	18
co	07	00	20	24	90	25	20	44	08		1.9	52
4	07	56	2,	02	60		21	00	08		20	27
S	80	20	21	80	10	20	21	2	60		21	04
9	60	43	22	12	10	58	22	27	10	34	21	4
c -	10	63	22	44	11	50	23	90	11	56	22	28
00	11	25	23	18	12	42	23	47	12	18	23	15
6	12	15	23	52	13	35			13	10		
10	13	0.7			14	29	00	34	14	000	00	07
11	14	00	00	30	15	21	01	25	14	49	01	03
12	14	53	0.1	10	16	13	02	19	15	36	02	02
13	15	80	01	55	17	01	03	19		21	03	04
4	16	42	02	44	17	49	10	21	17	90	04	07
15	17	35	03	38	18	33	02	25		49	05	13
16	18	26	04	37	19	17	90	30	18	33	90	19
17	19	14	05	38	20	00	07	36	19	20	0.2	26
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MOONRISE AND MOONSET

CAPE TOWN - TIMES OF MOONRISE AND MOONSET FOR PORT ELIZABETH SUBTRACT 28 MINUTES

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DURBAN – TIMES OF MOONRISE AND MOONSET FOR BLOEMFONTEIN ADD 19 MINUTES

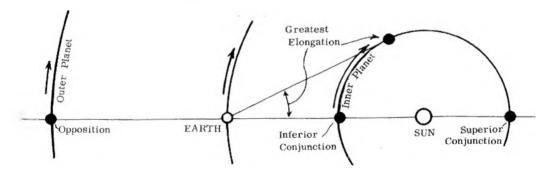
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THE PLANETS 1976

		BA:	SIC DATA			
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	10 ⁶ km	years	(Earth = 1)	10 ³ km		to Orbit
Mercury	58	0,24	0,056	4,98	59d	?
Venus	108	0,62	0,817	12,4	244 h m	23027'
Earth	150	1,00	1,000	12,8	23 ^h 56 ^m	23 27'
Mars	228	1.88	0,108	6,76	24 37	23 59
Jupiter	778	11,9	318,0	142,7	09 51	03 04
Saturn	1426	29,5	95,2	120,8	10 14	26 44
Uranus	2868	84,0	14,6	47,1	10 49	97 53
Neptune	4494	164,8	17,3	44,6	14 ?	28 48
Pluto	5896	247,6	0,9?	?	6d?	?

GENERAL

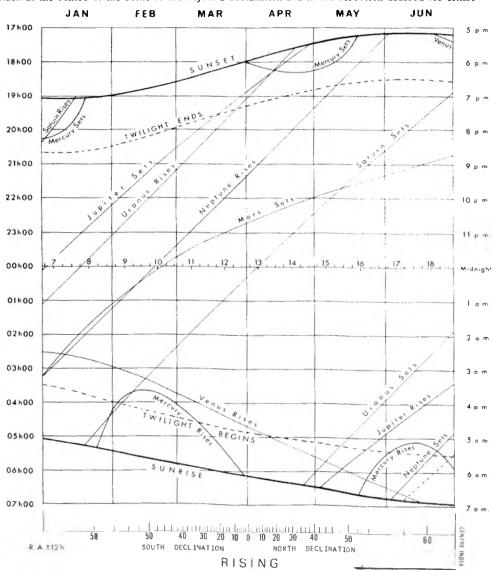
Apart from Uranus, Neptune and Pluto, the planets of our solar system are amongst the brightest objects in the night sky. Unlike the distant stars, their relative positions do not remain fixed, but continually change as, like the Earth, they orbit around the Sun. Their apparent movements against the starry background are complicated as they result from a combination of their own motion and the Earth's motion. Their brightnesses also vary considerably, as both their distances from the Earth and the visible portions of their sunlit hemispheres change. Since the period of a planet increases with increasing distance from the Sun, so we find that the inner planets - Mercury and Venus - appear to "overtake" the Earth in their orbits, while the Earth in turn "overtakes" the outer planets - Mars, Jupiter and Saturn. The terms given in astronomy to the various Sun-Earth-Planet configurations are illustrated in the accompanying diagram. Dates of such configurations occuring in 1976 are listed chronologically in the Diary (pages 4 and 5) and are also mentioned in the text below.



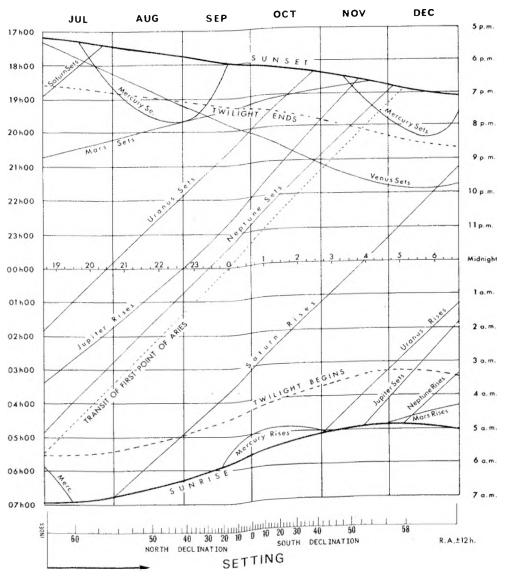
TIMES OF RISING AND SETTING

The times of rising and setting given by this diagram are accurate for position 30° East, 30° South, and approximately correct for other places in Southern Africa. Strictly speaking, corrections for latitude and longitude should be applied, but the latitude correction is, in general, sufficiently small to be ignored and in no case will exceed 15 minutes. Longitude corrections are given on page

The scales at the bottom of these pages are for finding rising or setting times of any object of which the right ascension and declination are known. Set dividers or a strip of paper from the index at the centre of the scale to the object's declination and in the direction desired for either



rising or setting. Measure this same distance and direction along the midnight line, beginning at the objects right ascension indicated by the numerals. (Should this end point fall outside the chart, 12 hours should be added to or subtracted from the right ascension. Reset the dividers using the end of the scale instead of the centre index, and measure in the opposite direction to that first used.) Through the point established draw a line parallel to the First Point of Aries transit line (indicated by the dashed line on the chart).



OBSERVING THE PLANETS

To the naked eye, planets appear as virtually pinpoint sources of light. However, their disks can be readily resolved with the aid of a small telescope. Even so, their angular diameters are of the order of 10 seconds of are - roughly 1/200 of the Moon's angular diameter - so it is not always possible to distinguish details on their disks. The disks of Mercury and Venus are only seen fully illuminated when they are furthest from us - as they draw closer, their disks grow larger but the phase changes to a crescent as we see more of their dark hemispheres. In contrast, the disks of the outer planets are always seen fully or near fully illuminated.

MERCURY

The innermost planet, Mercury, revolves around the Sun faster than any of the other planets. Being close to the Sun it can only be seen just after sunset or just before sunrise, when it is near greatest elongation (greatest angle between Mercury and Sun as seen from the Earth). It can never be seen near inferior conjunction (passing between Earth and Sun) or near superior conjunction (passing round the far side of Sun). The angular diameter of Mercury's disc rarely exceeds 10 seconds of arc so it is difficult to make out any features on the disc, but phases (like those of the Moon) might just be visible. Close up photography by the Mariner 10 spacecraft has shown that the planet's surface is incredibly similar to that of the Moon.

Greatest Elongation East	Jan 7 (19 ⁰)	Apr 28 (210)	Aug 26 (27°)	Dec 20 (20°)
Inferior Conjunction	Jan 23	May 20	Sept 22	
Greatest Elongation West	Feb 16 (26 ⁰)	Jun 15 (23°)	Oet 7 (180)	
Superior Conjunction	Apr 1	July 15	Nov 7	

VENUS

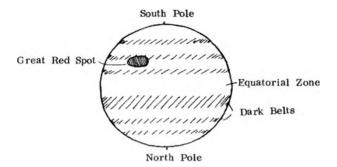
During the early months of the year, Venus will appear in the morning sky (above the eastern horizon). On June 18, it passes around the far side of the Sun and thereafter appears in the evening sky. As its distance from the Earth decreases it grows brighter and for the latter part of the year will be prominent in the western sky. If observed through a telescope over this period, its angular diameter will be seen to increase only slightly. Virtually no details can be seen on the disc of the planet because of the dense cloud covering.

MARS

Mars can be seen only in the evening sky until November when it passes behind the Sun. Its greatest angular diameter is only about 15 seconds of arc (about 1/100 of the Moon's diameter) at the beginning of the year and gradually decreases to 3 seconds of arc when it passes behind the Sun. Through a telescope it may be possible to glimpse the white polar caps and darker markings contrasting against the general "desert" orange colour of the disc. Regrettably Earth based telescopes cannot show the craters, conical volcanos and gigantic canyons that have been found on this planet. Mars has two tiny moons but a very large telescope is required to see these.

JUPITER

Jupiter is a prominent object in the evening sky in January and February, and from October onwards. It is at its brightest (magnitude ~2,4) at opposition on November 18. Because of its large angular size (45 seconds of arc at opposition), Jupiter makes an excellent object for a small telescope. It is often possible to see features on the disc; dark and light cloud bands, running parallel to the equator, and spots, in particular the famous Great Red Spot. These are indicated in the diagram below. The Great Red Spot is not always visible because of the rotation of the planet. Also clearly visible are four of Jupiter's thirteen moons. An entire section of this Handbook is devoted to the movements of these satellites and the phenomena associated with them (see page 23)



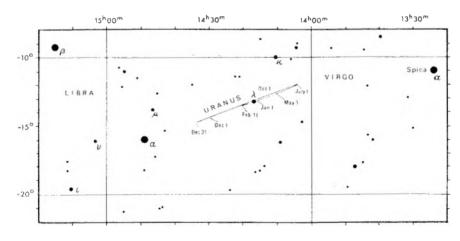
SATURN

Saturn will be clearly seen in the evening sky for the first six months of 1976 and again at the end of the year in the late evening. Unfortunately it will never be very high above the horizon because of its northerly declination - as it moves between the constellations of Cancer and Gemini. It is at greatest brightness (magnitude -0,1) at opposition on January 20. The diagram below shows its appearance through a small telescope (the scale is the same as for the Jupiter diagram) - including the spectacular ring system.



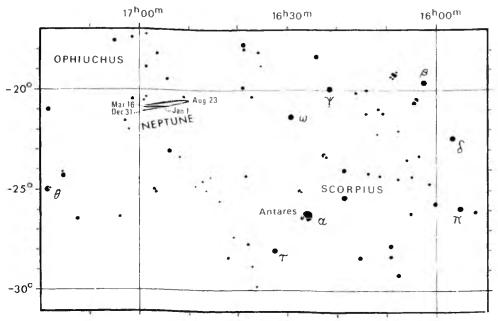
URANUS

Uranus at opposition on April 25 and conjunction on October 30, moves from the constellation of Virgo to the constellation of Libra early in the year. It is on the borderline of naked eye visibility (magnitude 5,7 at opposition) and may be found with the aid of binoculars and reference to the accompanying finding chart, which shows all stars in the region down to the same faintness. With a small telescope, its disc (angular diameter 4 seconds of are) may just be distinguished.



NEPTUNE

Neptune lies in the constellation of Ophiuchus close to the conspicuous constellation of Scorpius. It is far too faint to be seen with the naked eye - magnitude 7,7 at opposition on June 3 - but may be located using the finding chart below (which shows all stars down to magnitude 7,7) and a small telescope. Its angular diameter is only 2,5 seconds of arc, but its non-stellar appearance should aid in identification.



PLUTO

During the year Pluto moves in the constellation of Virgo between the coordinates RA 12^h58^m Dec + 12 24' and RA 13^h17^m Dec - 10 09'. Since it is very faint, magnitude 14, it can only be found using a large telescope and specially prepared finding charts.

THE MOONS OF JUPITER AND SATURN 1976

JUPITER'S MOONS

One of the most popular sights for an observer with a small telescope is Jupiter and its moons. Four of Jupiter's thirteen moons are large enough to be seen very easily – they would even be visible to the naked eye were it not for the glare of the mother planet. In order of increasing distance from Jupiter, the four moons are Io, which orbits once around Jupiter in less than 2 days; Europa, $3\frac{1}{2}$ days; Ganymede, 7 days; and Callisto which takes 17 days for a full circuit. All the orbits lie in Jupiter's equatorial plane and the system is seen almost edge on. As the moons circle Jupiter, they appear to oscillate from side to side alternatively passing in front of and behind the planet. Their configurations change from night to night and are shown in the diagrams on pages 34 and 35.

EVENTS RELATED TO THE MOONS PASSING IN FRONT OF AND BEHIND JUPITER

The table below lists all events occurring between the end of twilight and just after midnight when the planet is above the horizon in Southern Africa.

Explanation of table:

Date and predicted times are given - these are for mid-phenomena and are not instantaneous.

The moons concerned are I - Io III - Ganymede II - Europa IV - Callisto

Phenomena - the abbreviations used are:

Ec. - Eclipse: the satellite passes through the shadow of Jupiter D - Disappearance
Oc. - Occultation: the satellite is obscured by the disk of Jupiter R - Reappearance
Tr. - Transit: the satellite crosses the disk of Jupiter I - Ingress
Sh. - Shadow transit: the shadow of the satellite transits the disk E - Egress

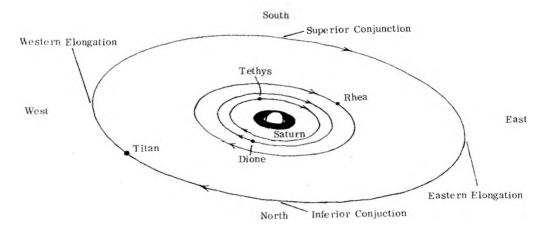
	0	20h361	11 111	Tr	F	Jan	14	20 ^h 18	o ^m I	Tr	E	Feb 17	20 ^h 11 ^m	H	Sh E
Jan	4							21 34	1 I	Sh	E				_
		23 32	Ш	Sh			20	20 32	lI S	Tr	Y				
	4	20 17	II	Oc	D			21 42				Sept 7	23 55	П	Sh I
		23 01	H	Oc	R					Ec		8	00 48	I	Tr I
		23 05	II	Ec	n			22 47	I	Oc	D	9	00 12	ı	Oc R
	_						21	20 00) I	Tr	I	-		:	
	5	22 58	I	Sh				21 19	1	Sh	I	16	23 12	1	Tr E
		23 49	I	Tr	E			22 12	_				23 37	II	Ec R
	6	20 29	11	Sh	E				_	Tr			23 38	II	Oc D
		22 26	I	Ec			22	20 22	II	Ee	\mathbf{R}	17	00 20	HI	Sh I
	_		_					20 45	I	Ec	R			711	
	9	21 59	III	Tr	I		27	20 21	Ш	Oc	n	23	00 30	1	Ec D
1	l 1	22 57	II	Oc	D			-					22 55	I	Tr I
	13	20 02	Ш	Ec	R		28	21 59	I	Sh	1		23 41	II	Ec D
		20 30	11	Tr	E	Feb	5	20 29	II (œ	D		23 52	I	Sh E
		20 33	II	Sh	I			21 13	3 I	Oc	D	24	22 20	1	Oc R
		20 50	I	Oc	n		6	20 40	1 T	Tr	F	25	23 20	II	Tr E
			_				•					_			
		23 05	II	Sh	E			21 51	. 1	Sh	E	27	23 12	Ш	Oc D

The Moons of Jupiter and Saturn 1976

Sep 30	$_{23}$ h $_{38}$ m	I	Sh I	Nov 4	00 ^h 09 ^m	II	Tr E	Dec	$20^{\rm h}45^{\rm m}$	III	Tr E
1	00 43	I	Tr I	8	22 04	I	Sh I		22 37	Ш	Sh E
2	00 08	Ţ	Oc R	8	22 20	I	Tr I	5	20 39	H	Sh [
	23 16	П	Tr I	9	00 13	I	Sh E		22 15	11	Tr E
	23 39	П	Sh E		00 28	Ī	Tr E		23 11	П	Sh E
4	22 25	Ш	Ec D		21 24	Ш	Oc R	8	23 40	I	Tr I
8	22 47	I	Ec D		21 47	I	Oc R	9	00 11	I	Sh I
9	23 03	I	Tr E	10	23 29	H	Sh I		20 59	I	Oc D
	23 42	II	Sh I		23 55	П	Tr I		23 43	I	Ec R
11	22 22	П	Oc R	12	20 30	П	Oc. R	10	20 15	I	Tr E
15	21 38	III	Tr E	15	23 58	ı	Sh I		20 50	Ţ	Sh E
16	21 54	I	Sh 1	16	00 03	I	Tr I	12	0 0 25	111	Sh 1
	22 42	I	Tr I		21 18	I	Ec D		22 - 02	11	Tr I
17	00 02	I	Sh E		22 29	111	Ec D		23 16	П	Sh I
	22 08	I	Oc R		23 31	1	Oc R	13	00 32	Ħ	Tr E
22	22 31	III	Sh E	17	20 37	1	Tr E	16	22 - 45	I	$O_C - D$
	23 15	Ш	Tr I		20 37	I	Sh E	17	20 35	Ţ	Sh I
23	23 47	I	Sh I	19	20 17	II	Oc D		22 - 01	[Tr F
24	00 26	I	Tr I		22 51	П	Ec R		22 - 45	ī	Sh E
	21 05	I	Ec D	23	$23 \ 05$	I	Oc. D	18	20 07	1	Ec. R
	23 53	I	Oc. R	24	20 12	I	Tr I	20	00 20	II .	Tr 4
25	21 00	E	Tr E		20 22	I	Sh I	21	22 - 35	11	Ec. R
	23 16	ΙΙ	Ec D		22 20	1	Tr E	22	20 - 49	111	Ee R
27	21 53	11	Tr E		22 31	I	Sh E	21	21 - 39	Ī	Tr E
31	23 00	I	Ec D	25	19 52	I	Ec R		22 - 30	Ī	Sh 1
Nov 1	20 10	I	Sh I	26	22 30	II	Oc D		20 48	1	T≀ E
	20 36	I	Tr I	28	20 35	Ħ	Sh E	25	22 - 02	į.	Ec. R
	22 19	I	Sh E	Dec 1	21 56	ſ	Tr - I	41.8	20.46	11	Oc. D
	22 44	I	Tr E		22 16	ĺ	Sb 1	29	20 - 53	111	Oc. R
2	20 03	I	Oc R	2	00 04	Ţ	Tr E		22 - 36	Ш	Ec. D
3	20 52	II	Sh I		00 26	I	Sh E	30	20 20	11	Sh E
	21 41	П	Tr I		21 47	1	Ec R	31	23 27	1	Tr I
	23 26	II	Sh E	4	20 24	Ш	Sh I	32	00 - 26	I	Sh 1

The Moons of Jupiter and Saturn 1976

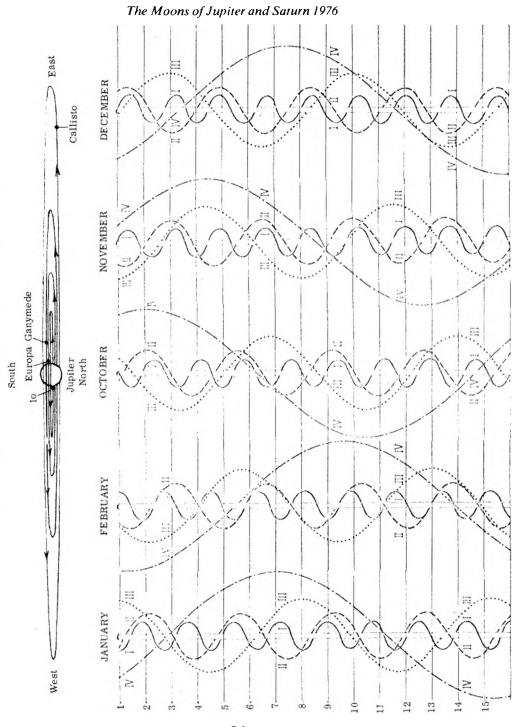
SATURN'S MOONS

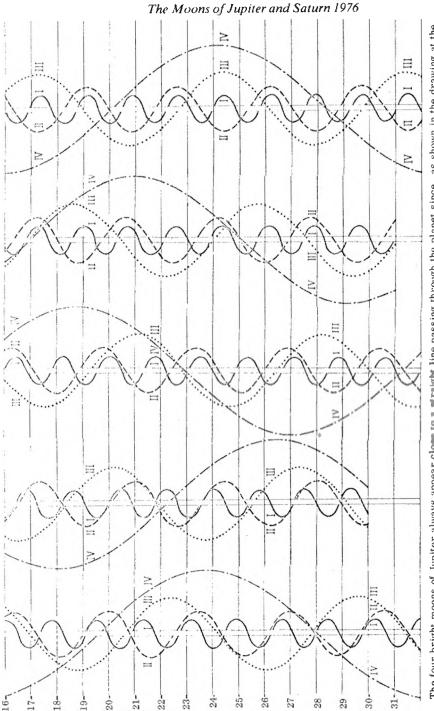


Saturn's moons are considerably fainter than the four Galilean moons of Jupiter. The diagram shows the orbits of four of Saturn's ten moons. The easiest to find is Titan (magnitude 5.5), according to the diagram and the information in the table below (which covers the period when Saturn is clearly visible in the evening sky).

TITAN 1976

Eastern Elongation	Inferior Conjunction	Western Elongation	Superior Conjunction
Jan 1	Jan 5	Jan 8	Jan 12
Jan 17	Jan 20	Jan 24	Jan 28
Feb 2	Feb 5	Feb 9	Feb 13
Feb 17	Feb 21	Feb 25	Feb 29
Mar 4	Mar 8	Mar 12	Mar 16
Mar 20	Mar 24	Mar 28	Apr 1
Apr 5	Apr 9	Apr 13	Apr 17
Apr 21	Apr 25	Apr 29	May 3
May 7	May 11	May 15	May 19
May 23	May 27	May 31	June 4
June 8	June 12		
5-2	· · ·		
Nov 15	Nov 19	Nov 23	Nov 27
Dec 1	Dec 5	Dec 9	Dec 13
Dec 17	Dec 21	Dec 25	Dec 29





top, their orbits are seen nearby edge on. The main part of the diagram then shows how their positions along such a straight line change The four bright moons of Jupiter always appear close to a straight line passing through the planet since, as shown in the drawing at the during the five months when Jupiter is prominent in the evening sky. For each month, time increases downward; the disk of Jupiter is --presenting 2 a.m. (0 hrs. Universal time), are shown for every day of the month. The wavy lines show how the Moons appear to escillate from one side of the planet to the other. stretched to make the central column and horizontal

COMETS AND METEORS

COMETS

Comets are celestial bodies moving around the sun, mostly in very elongated orbits. The typical comet consists of a nucleus surrounded by a hazy aura of gas and dust called the coma, and in many cases there is a tail stretching away from the sun. Faint comets, several of which are discovered each year, usually appear only as fuzzy patches without nucleus or tail.

While they are believed to be true members of the solar system, comets differ radically from the planets in that their orbits, besides being highly eccentric, are inclined at all angles to the plane of the ecliptic, and their motion may be direct (like that of the planets) or retrograde. Compared with that of a planet, a comet's mass is almost negligible; nearly all this mass is concentrated in the nucleus, which is believed to be not one solid piece but composed of many separate particles of various sizes.

Comets are the most mysterious and capricious of solar system objects and the nature of the physical changes which they exhibit is still not fully understood.

Observers with quite modest equipment, say a refracting telescope of not less than 7.5 cm, can do useful work by following known comets and reporting on their appearance. The ability to make accurate brightness estimates is especially useful and well worth cultivating. Sweeping the sky for new comets, though requiring considerable patience and perseverance, is also well within the scope of the equipment mentioned.

Many of the fainter comets are undoubtedly escaping detection, particularly in the Southern skies which are not being as intensively searched as the Northern. There is a need for more amateurs to undertake this work.

Interested persons are asked to contact the Director of the Comet and Meteor Section at 90 Malan Street, Riviera, Pretoria 0002,

METEORS

Meteors or "shooting stars" result from small bodies entering the Earth's upper atmosphere, and are generally seen in greater abundance after midnight (due to the direction of the Earth's motion) than in the early evening.

There are two categories of meteors — the sporadic ones and the showers. A meteor shower comes from a certain direction in space (the Radiant) and is thought to be associated with the remains of a comet. When the Earth passes close to the comet's original orbit, such a shower can be expected. A list of these predicted showers is given in the table opposite.

The term "shower" is perhaps misleading (as the table indicates); the most prolific of these showers normally yields an average of less than one meteor per minute. On rare occasions however, as in the case of the Leonids, there is a phenomenal rise in the number of meteors observed.

There is always the possibility of new showers occurring, and any large-scale meteor activity observed on dates other than those mentioned should be reported without delay.

Reports by a reliable observer of the number of moteors seen coming from a particular radiant in a given period are always useful, but the best work is done by organised teams making a full sky coverage.

"Fireballs" are meteors of a luminosity equalling or exceeding that of the brightest planets. Accurate reports of their path among the stars, or their altitude and azimuth, at specific times, are of great value, particularly if made by observers at different places along the trajectory.

PREDICTED METEOR SHOWERS 1976

CTF
20 -48 Mar 16
16 -27 Mar 25 ?
18 08 +32 Apr 22 15
22 24 00 May 5 18
18 00 -30 Jun 14 ?
18 32 +35 Jun 16 8
17 20 -20 Jun 13(?) 8
21 00 -15 Jul 25
22 36 {-17 Jul 27 35
22 40 -30 Jul 30
36
22 04 - 6 Aug 22 32 -15 Aug
06 24 +15 Oct 20
03 44 +14 Nov 8
10 08 +22 Nov IT
01 00 -55 Dec
37 28 +32 Dec 14
09 56 -51 Dec 10

THE STARS

CONSTELLATIONS

Apart from our Sun all the stars that we see are so incredibly distant that, despite their high speed velocities, their apparent positions change by only minute amounts each year. Consequently the patterns that they form appear unchanged. The Greeks and other ancient civilisations identified these patterns, or constellations, with various mythological characters and creatures, and most of the names they gave are still used today.

In all there are 88 constellations, roughly one half of which would be above the horizon at any one time. Some contain distinctive patterns of bright stars and are relatively easy to find; others are difficult to locate, even with suitable maps. The Southern Cross and Centaurus, Orion and Taurus, Scorpius and Sagittarius, are featured later in this section. Detailed information on other constellations is beyond the scope of this handbook and interested observers are advised to obtain a suitable star atlas.

STAR NAMES

Within each constellation, the brightest star is generally labelled α (Alpha), the next β (Beta) and so on through the Greek alphabet. Most of the brightest stars also have their own names - usually of arabic origin. For example α Canis Majoris, otherwise known as Sirius, is the brightest star in the constellation Canis Major.

STELLAR MAGNITUDES

The apparent brightness of a star - which depends both on its true luminosity and its distance - is indicated by its magnitude. Equal intervals of magnitude represent equal ratios in light intensity. A star of magnitude 1,0 (typical of the brightest stars in the night sky) would be exactly one hundred times more luminous than a star of magnitude 6,0 (about the limit of visibility to the naked eye). The maps in this section show stars down to magnitude 4,5.

STELLAR DISTANCES

Distances are often expressed in units of light years - the distance light would travel in a year (equal to 9.5 x 10^{12} km).

DOUBLE STARS

It now appears that single stars such as our Sun are the exception, the majority of stars being double or multiple - two or more suns in orbit around one another.

STAR CLUSTERS

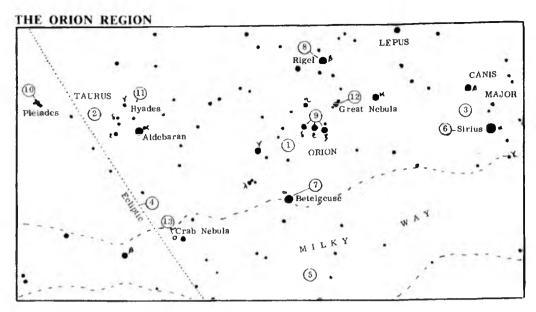
These are of two completely different sorts. Galactic clusters, having of the order of 100 stars, are found close to the plane of the Milky Way. The ones we can see are relatively nearby. Globular clusters are much larger and far more distant. They contain of the order of 100 000 stars each and are seen above and below the Milky Way on that side of the sky towards the centre of our galaxy. So great is their distance that small telescopes fall to resolve individual stars - instead they appear as fuzzy balls.

NEBULAE

Possibly one third of the matter in our region of the Galaxy is in the form of gas and dust (the remainder being contained in stars, plus a tiny amount in planets). Condensations of this material are called nebulae, some of which are illuminated by nearby stars while others are dark. They are usually referred to by their numbers in Messier's catalogue (M) or the New General Catalogue (NGC).

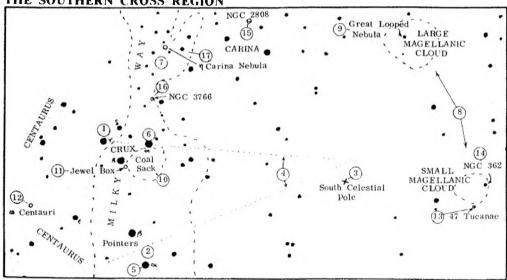
THREE POPULAR REGIONS

The dominating constellation of the summer skies is Orion, that of the winter skies is Scorpius, while the Southern Cross is conspicuous for most of the year. The regions around these constellations are also rich in interesting objects - visible either to the naked eye, or with the aid of binoculars or a small telescope - and are featured in the maps and text below. It may be necessary to rotate the maps to match the orientation of the constellations in the sky.



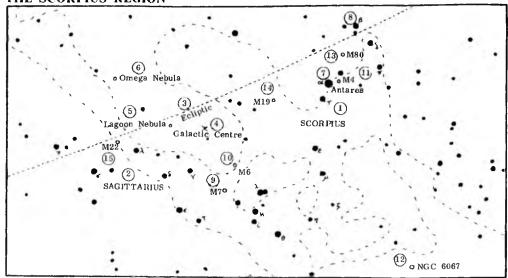
- 1 The constellation of Orion. The figure of the legendary hunter of Greek mythology is unfortunately upside down when seen from Southern Africa. The faint stars by λ represent the head, α and Y the shoulders, $\delta \epsilon n$ the belt, and β and K the legs. Orion forms part of the "great hunting scene" in which he faces the onslaught of 2 Taurus, the bull. Only the forepart of the bull is depicted and, like Orion, it is upside down. α and ϵ are the eyes, Y the nose. Orion is accompanied by 3 Canis major, the large dog, and the small dog (off map) while Lepus, the hare, crouches at his feet.
- A section of the Ecliptic a line encircling the entire sky and representing the plane of the Earth's orbit. As the Earth revolves around the Sun, the Sun appears to move along the ecliptic through the constellations of the Zodiac, of which Taurus is one.
- (5) A portion of the Milky Way (looking out towards the edge of our Galaxy).
- 6 Sirius the brightest star in the night sky. It is somewhat brighter than our Sun and relatively close by at a distance of 9 light years. It is a double star but the companion is a white dwarf (only slightly larger than the Earth, and with a mass comparable to our Sun) and is only visible through a large telescope.
- (7) Betelgeuse most famous of the red giant stars. Its diameter is of the order of the size of the Earth's orbit and its luminosity is nearly 10 000 times that of our Sun. Its red colour should be obvious to the eye. It is 520 light years distant.
- (8) Rigel, despite being physically smaller than Betelgeuse, is more luminous (higher surface temperature bluish colour) and more distant.
- (9) The stars in Orion's belt are distant hot blue stars.
- (10) The Pleiades or Seven Sisters form the best known nearby star cluster. Six or seven stars are visible to the naked eye, binoculars or a small telescope show more.
- (11) The Hyades is another nearby galactic cluster, but Aldebaran is not a member (it lies closer to us).
- (2) The Great Nebula in Orion, just visible to the naked eye, shows up as a fan shaped mass of luminous gas through binoculars or a telescope. A telescope will also show a tiny "Trapezium" of four stars in the centre.
- (3) The Crab Nebula, the remnant of a supernova recorded by the Chinese in 1054, requires a moderate sized telescope for observation. In its heart is located the extraodinary pulsar which emits a double flash of light 30 times every second. The current belief is that it is a rapidly rotating neutron star a star with the mass of our sun but with a diameter of only 10 km.

THE SOUTHERN CROSS REGION



- (1) Crux, the Southern Cross, is one of the most compact patterns of bright stars to be found in the sky. It lies on the border of that region of the sky which never sets as seen from Southern Africa.
- (2) The two "Pointer" stars lie close to the Cross. (A similar pattern to the Southern Cross called the False Cross lies just outside and above the map, but has no accompanying pointer stars).
- 3 The South Celestial Pole: This is one of two opposite points in space towards which the Earth's axis of rotation is directed. As the Earth rotates so the sky appears to pivot about this point. It always lies above the south point on the horizon, elevated by an angle equal to the observer's southern latitude. (The north celestial pole lies below the northern horizon and can never be seen from the Earth's southern hemisphere)
- 4 The intersection of a line extended through the major axis of the Cross and the perpendicular bisector to the Pointers indicates the approximate position of the South Celestial Pole.
- 6) a Centauri has the distinction of being the closest star to our solar system at a distance of approximately 40 million million km or 4,3 light years. A small telescope readily shows that it is a double star the two components take 80 years to revolve about one another. A much fainter third star also belongs to the system.
- 6) a Crucis can also be resolved as a double star by a small telescope (separation 5 sec of arc).
- 7) The region indicated is one of the brightest sections of the entire Milky Way.
- (8) The Large and Small Magellanic Clouds are the nearest of the external galaxies (see also next section). They can be seen with the naked eye provided the sky is reasonably dark.
- The Great Looped Nebula possibly the remnant of a supernova explosion in the Large Magellanic Cloud. (Naked eye or binoculars).
- (10) The "Coal Sack" a dark mass of gas and dust obscuring a part of the Milky Way. (Naked eye or binoculars).
- (1) Herschel's "Jewel Box" a galactic cluster containing stars of different colours. (Small telescope or binoculars).
- (12) ω Centauri and (13) 47 Tucanae are perhaps the best known globular clusters. Binoculars will show their fuzzy appearance. (14) NGC 362 and (15) NGC 2808 are fainter globular clusters.
- 16 NGC 3760 a fine galactic cluster. (Binoculars or small telescope).
- (17) The m Carinae nebula site of a slow supernova that brightened to magnitude -0.8 in 1843 and is now of magnitude 6.4.

THE SCORPIUS REGION



- (1) The constellation of Scorpius. The creature is depicted with α in the centre of the body and β and π the claws. The distinctive tail $\epsilon = \zeta = \theta$ curls round to the sting λ .
- (2) Sagittarius the figure of the centaur archer is very difficult to make out.
- (3) A section of the Ecliptic. Like Taurus, Scorpius and Sagittarius are constellations of the Zodiac.
- 4 The direction of the centre of our Galaxy the Milky Way is that part of our Galaxy visible to us. Unfortunately the central nucleus is obscured by foreground gaseous and dusty matter both dark and luminous hence the irregular shape of the Milky Way in this region. Luminous nebulae include (5) the Lagoon nebula and (6) the Omega nebula. These are best seen with the aid of binoculars.
- (7) Antares a distant red giant, several hundred times the diameter of our Sun is so named because its red colour rivals that of the planet Mars.
- (B) β Scorpii can be resolved as a double star (separation 16 sec of arc) with a small telescope. In fact
 the brighter component is in itself a triple star, and the fainter component a double star!

This region includes a number of galactic clusters including (9) M7, (1) M8, (1) M4 and (12) NGC 6067. (Use binoculars or a small telescope).

Further from the plane of the Milky Way are some globular clusters: 13 M80 14 M19 and 15 M22.

NOVA SEARCHING

On rare occasions a star may undergo a nova outburst, its light increasing tremendously. The result is that a "new" star appears where previously no star was visible to the naked eye, or even with a small telescope. The light of the nova may fluctuate for a time, then gradually fades over a period of days, weeks or months.

Even observers having no telescopes can perform a useful task in keeping a watch for such novae in an allocated area of the sky. A good knowledge of the constellations is a recommendation, since part of the procedure is to scan the whole sky for bright novae before the more detailed search in the allocated area is begun. However, anyone can be given training in star recognition.

Interested persons should contact the Director of the Nova Search Section, Mr. J.C. Bennett, 90 Malan Street, Riviera, Pretoria 0002.

VARIABLE STAR OBSERVING

The General Catalogue of Variable Stars lists some 20 000 stars. Professional observatories cannot possibly monitor all of these and the observation of variable stars is a field therefore in which amateurs can make real contributions to astronomical knowledge.

Of the 20 000 stars, at least 2 000 are suitable for monitoring by Southern Hemisphere observers but less than 200 are in fact observed from South Africa and a still smaller number receive adequate attention.

The Variable Star Section of the ASSA exists for the purpose of encouraging observers and of acting as a medium of communication. The Section disseminates incoming information amongst observers and will forward (on request) the observations of individuals to various variable star bodies. These include the American Association of Variable Star Observers and the Variable Star Sections of the British Astronomical Association and Royal Astronomical Society of New Zealand. These bodies combine the South African observers' light estimates with those from other parts of the world. The resulting light curves and tables are sent to a large number of professional observatories where astronomers are interested in investigating certain of the stars more fully.

In addition to the international work, the VSS of the ASSA supplies information direct to certain South African astronomers. It is in a position to warn observatories of sudden changes in certain "VIP" stars before the overseas bodies can do so.

Some of these VIP stars are:

111661	RS Cen	145971	S Aps
123753	UW Cen	174406	Rs Oph
124728	Ех Нуа	181824	Gu Sgr
131953	Dy Hya	183423	V348 Sg1
132554	Bv Cen	191033	Ry Sgr
135870b	Z Aps		

Most of the above stars could be observed every night and would still not be over-observed.

Many long period variable stars vary through a large range of brightness. This means that even approximate estimates by beginners can generate quite acceptable light curves. Some of these stars are relatively easy to locate and identify and this makes them very suitable for new observers. Examples of this class of stars are:

```
134236 RT Cen
025050 R Hor
                         161122 a and b R and S Sco
034625
       U Eri
                         165030a Rr Seo
035124 T Eri
040725
       W Eri
                         174162 W Pav
043262 R Ret
                         180222a VX Sgr
                         180363 R Pav
054331
       S Col
091868
       Rw Car
                         191019
                                R Sgr
094953
       Z Vel
                         214247 R Gru
100661
       S Car
```

Several variable stars are situated near the South Celestial Pole and can be observed throughout the year during the evening hours. Three such stars are:

```
055686 R Oct 131283 U Oct 172486 S Oct
```

Others will need early morning observing sessions for fuller coverage.

Certain stars have such fast variations that they need to be monitored continuously, preferably by several observers working independently. Amateurs were asked to monitor 013418 UV Ceti in this way during the 1975 Appollo- Soyuz space mission. It is possible that the Variable Star Section will again be asked to participate in exercises of this nature.

A few variables are so bright that they can be monitored throughout their cycles without optical aid. 071044 L 2 Puppis is such a star. Nortons's Star Atlas names several more.

From the above it can be seen that there are variable stars to suit amateurs with different amounts of experience, ability, optical aid and spare time. Amateurs who want to start observing stars in any of the above categories should contact the Director, Mr. M.D. Overbeek, 60 Edward Drive, Glendower, Edenvale, 1610. New observers will be given charts of a few easy objects and (if possible) a certain amount of instruction at the eyepiece. When writing, prospective observers should give brief details of their equipment.

ORDINARY OCCULTATIONS

An occultation occurs when the moon passes in front of a star. The disappearance or reappearance of the star is an instantaneous event, or almost so, and is therefore easily timed. Time signals are broadcast continuously through the 24 hours from station ZUO (see page 54) and, radio propagation permitting, occultation timings can be made, either by eye and ear or by stop watch.

The essential record of a occultation observation consists of

- (i) Date and observed time of the event (to a tenth of a second if possible). Universal Time is preferable.
- (ii) The Z.C. number of the star.
- (iii) Whether disappearance or reappearance.
- (iv) The observer's estimate of the accuracy of the observation.
- (v) Whether the time given in (i) has been corrected for personal error, and if so by how much.

Also required are the size and type of telescope used, the method of timing used, and the position of the telescope.

The telescope position must be to an accuracy of 1" or better in latitude and longitude and 30 metres or better in altitude. An observer can usually read his position off a 1/50 000 map published by the Director General of Trig. Survey, Mowbray, or his equivalent in other countries. Positions determined astronomically are NOT acceptable, no matter how carefully determined.

Timings of occultations are very valuable in determining the moon's shape and motion and constitute a field in which the amateur astronomer, often with only a small, unsophisticated telescope, can make an extremely useful contribution. Interested persons are urged to contact the Director of the Society's Occultation Section, Mr. A.G. F. Morrisby (c/o Dept. of Surveyor General, P. O. Box 8099, Causeway, Salisbury, Rhodesia).

Predictions of occultations of stars brighter than magnitude 7,5, supplied by H. M. Nautical Almanac Office, are reproduced on the next two pages.

Explanation of Table:

- Z.C. the number of the star in the "Catalogue of 3539 Zodiacal Stars for the Equinox 1950.0" by James Robertson (U.S. Naval Observatory, 1939). A short index of the brighter stars is given on page 47.
- Mag. the visual magnitude of the star
- P = Phase. 1 = Disappearance 2 = Reappearance
- El. of
 - Moon the elongation of the Moon. 0° = New Moon, 90° = First Quarter, 180° = Full Moon, 270° = Last Quarter.
- U.T. The predicted time in Universal Time which is exactly two hours behind South African Standard Time. For example 0h55, UT is 2h55, a.m. SAST.
- a,b the approximate time of an occultation at a place $\Delta\lambda$ degrees west and $\Delta\phi$ degrees north of the city concerned is
 - Predicted time + $a.\Delta\lambda$ + $b.\Delta\phi$
 - where a and b are in minutes of time.
- P the position angle on the Moon's limb, measured eastward from the north point.
- N = no occultation A Moon at very low altitude
- S = sunlight interferes G grazing occultation

Occultation Predictions

				ι.		PE TO					annes					LISBU		
Date	Z.C.	Mag.	P. 0		18.50	00, s	. 33. 9	000	E.	28.0	00, S	26.2	200	E.	o 31.1	00, S	. 17.	700
	110.		FA		T.T.	a	ъ	P	U	т.	а	ь	P	U	т.	a	b	P
Jan. 1	1 416 1 429	5.4 6.9	1 1: 1 1: 1 1:		45.2 17.1 31.5	S -2.1	-0.2 +0.6	97	h 19 20	54.4 43.0	m -0.5 S -1.8		0 103 78	h 16 21	02.7	m A -1.7 -1.8		
12				26	21.2	S	T2.J	4)	17	22.4	-2.3	+0.6	65	17		-2.4	+2.0	44
12 13 14 14 15	700 847 847	5.7 3.0 3.0	1 15	38 50 20 50 21	20.7	S N -2.4 -2.4 S	-0.9 +0.9	113 255	20	41.5	N -2.7	+0.1 +0.1 +0.2	93	20 22	57.8 55.8 14.0	-3.1 -3.1 -1.5 -0.8	+1.3	136 67 310
19 19 21 22 22	1519 1649 1886	6.3	2 25	16 21 31 0 57 22	48.1 51.9 19.0	-1.4	-2.0	319	21	48.2 37.6	-'.4 K		308	21	32.7	-1.1 N N	-2.9	335
27 27 Feb. 3	1925 2456 3453	1.2 1.2 6.2 4.9 6.4	2 26 2 31 1 3	51 8 11 39		-0.6 N N	+1.6	259	8	49.6	-0.3 N N	+0.4	283	8 2	48.9 23.4	-0.6 -0.4 -2.0 -0.9	-0.9 +1.9	313 224
20 2. 2. Mar. 6	2118 2117 2118	2.9 5.3 2.9	2 25	54 23 54 0	16.9 41.4 38.5 45.1 18.3	-0.4 -1.4 -1.5	-2.1 -1.1 -0.9	275 269	0	39.3	-1.2 -1.4	-1 -2 -1.9 +2.8	3.0	0	16.9	N -2.3 0.0 -0.5 N	-4.3	348
10 12 13 23 28	1271 3 1410 3 2680	5.8	1 13 1 15	55 50 20 76 2	48.1 43.3		-3.2			44.7	-3.0 N -'.8	-1.3	66 130 335	20	44.5	N -2.6 N -0.5	-0.3	100
Apr. 5	971 1364 1489	7.3 6.5 6.8	1 7 1 13 1 13	7 21	37.9 45.3 26.0			86 49 158	17 19		-2.2 N A N	+1.5	7* 38	18	30.2	G N N A -1.6	+1.6	229
15 15 17 21 22	2118 2436	2.9 6.3 5.0	2 22	00 28 58		N N N N -2.0	-3.8	306		47.6		+0.1 +1.4		18 19	29.5 45.9	-0.3 -1.5 N	-0.7	215 274
6	913 915 1190 1320 1429		1 7	18 18 16 73 18 36 19 98 18	24.8 37.2 48.6	-0.7 +0.1	-0.9 -2.6		18	43.6	-1.0	+2.1 +0.4 -0.4			56.1		+1.8	68

Occultation Predictions

			P.)		CAPE TOWN			JOHANNESBURG						
Date		Mag.		Ē.	18.500, S. 33.9	00	E.	28.000, S. 26.2	200	E.	31.100, S. 17.700			
May 8	No. 1546 1779	7.2 6.7	Moor 0 1 110 1 138	n 17	m m m 01.5 -1.7 -1.6 14.4 -2.1 -0.1	0 129 70		m m m 13.7 -2.7 -0.3	o 95	h	m m m o G N			
16 18 18	2686 2876	5.2 2 5.4 2		19 1	42.9 +0.1 -1.4 56.7 -1.7 +3.2 17.4 -1.6 +2.8	287 209	2	24.7 +0.6 -3.3 34.4 -2.0 +2.2	229	2 !	N 55.2 – 2.5 +1.1 253 S			
		5.8 2 6.4 2 5.7 2 5.7 6.4	2 285 2 30'7 1 122	4 22		2 44 46	1	46.3 -2.0 -2.3 S N	290		39.0 -2.1 +2.5 218 N S N 25.3 -0.3 +0.2 102			
17 19 30 July 4 6	3455 1364 1843	6.4 2	2 253 2 265 1 37 1 91	0	N 36.9 -1.3 -1.6 S 42.3 -0.9 -0.1	122	0 16 20	30.9	304 52 93	21 (22.5 -1.6 +2.0 219 N N 05.6 -0.3 +2.4 56 42.5 +0.3 +4.1 34			
13	2232 2259 3133 3272 3287	7.2 6.7 6.7 5.8 25.8 25.8 25.8	132 208 220	0 19 20		56 242 273	0 19 20	32.3 -1.1 -3.9	309		N N 01.7 -0.4 -3.6 314 N 20.1 -2.2 +1.7 241			
30	1564 1685 1792	5.8 2 6.6 4.5 7.1 7.2	33 46 59	17	A N 15.2 -1.3 +1.2 N G	86	17	21.7 -0.6 -1.0 25.9 -0.2 -1.0 N 26.0 05.7 -0.8 -0.9	174	17	06.0			
4 6 6 6	2345 2497 2649 2658 2680	6.9 6.6 6.6 5.8	125 138 139	18		70 83	16 18	G 32.9 27.0 -2.5 +1.9	19 50		47.1 -3.6 +2.8 50 N N N 55.1 -0.8 +0.5 94			
7 16 16 17 19	2686 299 403 517 668	6.3 5.8 6.4	1 141 2 248 2 257 2 268 1 280	0 2	09.9 -0.7 +1.6 52.2 -2.6 +0.3 N A N			26.6 -0.2 +1.6 22.3 -3.0 +0.5 A A N	62 266	22 :	42.0 +0.3 +2.1 39 S 22.4 -0.1 +1.2 221 56.8 -0.9 -1.1 281 11.6 139			
19 20 27 27 28		3.6 6.4 0.5 0.5 5.7	1 25	10 11	N - N - 28.5 -1.2 -2.6 43.2 -2.6 -0.3	146 270	11	N A 28.9 -2.2 -1.4 58.5 -2.1 -1.6 58.8	308	1 (10) 11 -	43.7 191 07.1 -0.3 +1.0 229 31.6 -3.5 +0.3 81 42.8 -1.2 -3.7 342 44.5 -1.0 -1.0 130			
30	2160	6.6 6.8 6.8	1 69	19	29.5 -1.7 -3.1 N 53.1 -0.6 +1.1 N 16.2 -1.9 +0.5	89	20 17	34.3 -1.7 -0.6 N 06.6 -0.1 +1.5 12.8 40.1 -1.4 +1.5	69	17 9	39.2 -1.6 +0.5 91 59.5 -1.7 -1.9 141 A 02.4 -2.5 -0.9 115 02.9 -0.5 +3.5 38			
-	2316 2463 2640 2647 2789	6.9 6.1	1 83 1 96 1 111 1 111 1 122	0	19.7 -1.5 +0.1 N 23.1 -0.6 +0.3 11.9 0.0 +1.0 44.6 -1.6 +1.8	115 90	20	A A		20	47.7 -0.6 +1.5 65 19.1 -1.4 -0.2 109 A A 38.9 -0.1 +3.5 23			

Occultation Predictions

OCCULTATION PREDICTIONS

		CAPE TOWN						JOHANNESBURG SALISBURY							
v	ate			E.	P. of	E				E	0 .28.000, S. 26		Ε.		0 7.700
		No.			Мо	on U	.Т. а	ъ	P	IJ	.T. a b	P	U	.T. a b	P
	4 5 13 16 16	2913 3045 374 760 888	5.0 6.0 6.1 6.5 6.0	1	146		N	Ю.9	239		55.7-2.8-1.3 N 58.5-2.7+1.3 S 59.3-1.3-1.8	247	17	m m m 02.0-2.8+0.6 53.8-2.9-1.3 15.6-3.1+0.5 8	266
Oct.	17 27 29 29	895 2259 2573 2578 2880	6.7 7.3 6.4	2 1 1 1	272 52 79 80 103		N N			21	34.9-1.7-0.2 N 02.4-0.3+0.5 04.6	103	18 19	37.5-2.3-0.7 48.6-1.2-2.5 33.2-1.3-0.4 A 05.5-3.2 0.0	151
	4 5 18 25 27	3272 3287 1364 2361 2686	5.8 6.5	1 2 1 1	141 289	18 2 18	09.0-1.7- N 35.0-1.2- 11.0-0.7- N	-0.3	246	18	37.2-1.9+2.1 N S A N	44		07.1-1.1+4.0 06.6-1.0-0.2 S Å	
Nov.	30 30 31 1 2	3111 3125 3229 3371 3477	6.8 6.9 5.6 6.4 6.6	1 1 1 1	110		% 46.3-0.2- S 19.5-0.6-		7 0 52	20 17	13.6		20 18 16	^6.36+0.2 A 07.4-2.4+2.3 A 28.8-1.1-2.8	100 47 24
	8 9 9	668 668 793 798 940	3.6 3.6 6.2 6.4 5.7	1 2 2 2 2	210	21 22 21 23 22	20.8 . 06.6 . 55.5 . 15.7-1.8-	:	199 327	22	27.0-3.1-2.3 39.0-2.1+2.4 N 45.0-2.8+0.9	215	23	03.1-3.1+1.4 N	
Dec.	11 13 29 29	947 1197 3444 3455 143	5.2 6.0 6.5 6.4 6.8	1	245 102 103	20 22	N N S 39.2-0.6- 26.1-1.5-		13		N N 25.1-2.7+0.9 G 44.4-1.0+1.1		0 17	16.5 08.3-2.5+0.8 45.0-2.4+1.8 N 58.0-0.8+1.5	213 244 57
	1 3 5 5 25	145 374 4005 4005 3272	6.1 -2.4 -2.4	1 1 1 2 1	147 160 160	22 19 0 1	27.6-2.1- 43.3-1.5-	+0.5		-	56.7-2.5+1.4 13.8-1.4+3.2 A S			N 21.7-2.4+2.8 N N 26.5-0.6+2.9	

INDEX OF OCCULTED STARS BRIGHTER THAN FIFTH MAGNITUDE

Z.C.	Name	Mag.	Sp
628	Omega Tauri	4,8	A3
668	Epsilon Tauri	3,6	K0
847	Zeta Tauri	3,0	ВЗр
915	Chi 2 Orionis	4,7	B2p
995	Mu Geminorum	4,1	B 5
1685	Upsilon Leonis	4,5	K0
1925	Alpha Virginis (Spica)	1,2	B2
2118	Alpha Librae	2,9	A3
2361	Chi Ophiuchi	4,8	ВЗр
2913	61 Sagi Harii	5,0	A0
3453	Kappa Piscium	4,9	A2p
"4001"	Mercury		
"4005"	Jupiter		

"STOP PRESS" Information on Comet West has been received as this handbook goes to press. It is included on this page because of space available.

Predictions:

Date	R.A.	Dec.	Magnitude
Jan 3	21 ^h 11 ^m ,8	-34° 50°	10,0
Jan 23	21 52,7	-31 09	8,2
Feb 12	22 48,9	-23 17	4,5
Feb 22	23 11,1	-12 50	0,5
Mar 3	22 04,0	+ 4 09	1,6
Mar 13	21 19,4	+ 9 36	4,6
Mar 23	21 00,4	+11 59	6,3
Apr 12	20 36,4	+15 11	8,4
May 2	20 06,1	+17 30	9,6

(Information obtained from I.A.U. Circular 2871)

GRAZING OCCULTATIONS

When a star moves tangentially to the limb of the Moon, and is occulted for a very short period only - a few minutes, or even seconds - a grazing occultation is said to occur. Because the limb, as seen from the Earth, is in fact the outline of numerous mountains and valleys, there may be several disappearances and reappearances, which are not only fascinating to observe, but which may be accurately timed to yield valuable data on the relative positions of star and Moon, in both right ascension and declination, as well as on the shape of the Moon. Some of these data cannot readily be obtained in any other way.

The maps on the following pages have been prepared by H.M. Nautical Almanac Office to show the tracks of stars brighter than 7.5 magnitude which will graze the limb of the Moon when it is at a favourable elongation from the Sun and at least 10° above the observer's horizon (2° in the case of bright stars). Each track starts in the West at some arbitrary time given in the key and ends beyong the area of interest, except where the letters "A", "B", or "S" are given. "A" denotes that the Moon is at a low altitude, "B" that the bright limb interferes, and "S" that sunlight interferes. The tick marks along the tracks denote 5 minute intervals of time which, when added to the time at the beginning of the track, give the approximate time of the graze at places along the tracks.

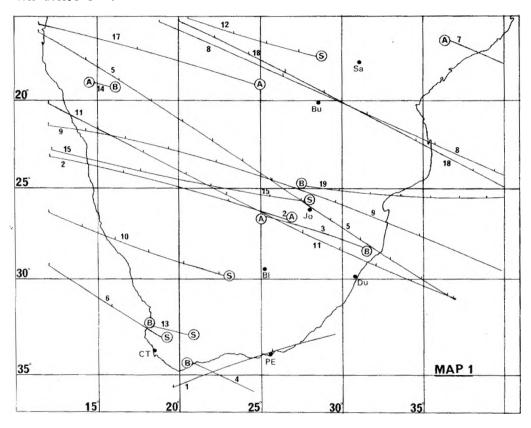
The tracks as shown on the maps are approximate only. Since the observer's location is very critical, successful observations call for very accurate predictions. With the aid of the IBM computer of the CSIR at Pretoria such predictions are at present prepared at 6-monthly intervals for a number of centres in South Africa, Rhodesia and Malawi. By plotting the predicted graze track on a reliable survey map (e.g. the South African 1:50 000 series) it is usually possible to select a convenient site from where the graze may be observed. Ideally a team of observers would be stationed at intervals along a line running at right angles to the graze track - say, along a main road - each with his own telescope and timing equipment. Each observer will see a different sequence of events, the combined results forming an accurate picture of the limb of the Moon.

The equipment needed is similar to that used for ordinary (or 'total') occultations, but must, of course, be portable. A 75 mm refractor is ideal for average events, but better instruments with a larger aperture have often shown their superiority under difficult conditions. Timing is best carried out with a portable tape recorder and radio receiver tuned to ZUO or other time signal station.

It will be seen from the maps that many grazing occultations occur in regions which are rather far removed from the main cities, and which cannot easily be reached by teams of observers from one of the ASSA centres. It is worth remembering, however, that a team of many observers, while ideal, is by no means essential; that a single good observer is worth more than many unsuccessful ones, and that one good observation is worth infinitely more than no observations at all.

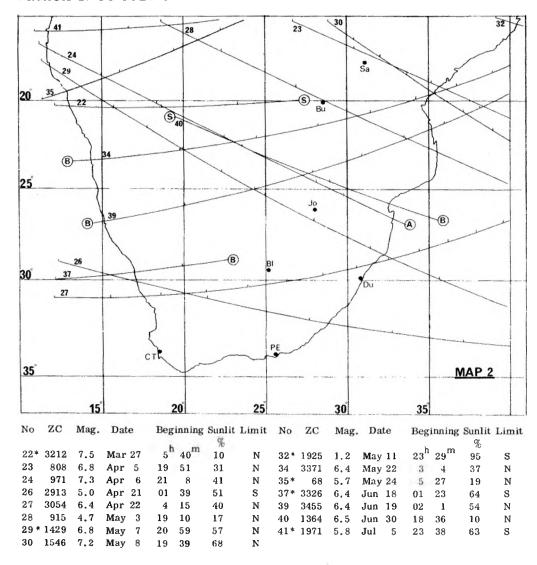
Observers in other parts of southern Africa - especially the more distant regions - who may be interested, are therefore invited to contact the coordinator for grazing occultations: Mr. J. Hers, 48, Central Road, Linden Extension, Randburg 2001, so that they may be informed of all favourable grazes occurring within their neighbourhood.

JANUARY 1 TO MARCH 23

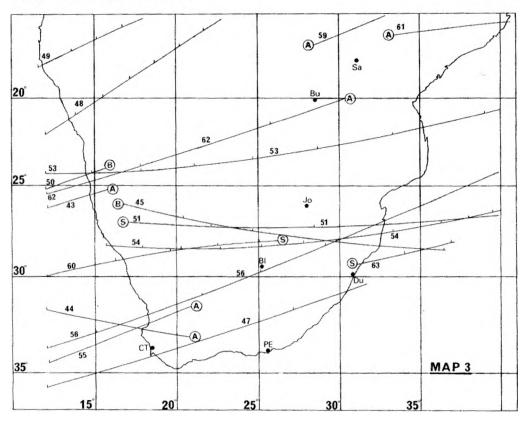


No	ZC	Mag.	Date	Begi	inning	Sunlit	Limit	No	ZC	Mag.	Date	Beg	inning	Sunlit	Limit
				h	m	%						ь	m	%	
1	3260	6.4	Jan 5	20"	32 m	17	N	10	2132	7.1	Feb 21	5 ^h	23^{m}	62	S
2	433	5.6	Jan 12	0	7	72	N	11	2264	7.4	Feb 22	3	13	51	S
3	1886	5.7	Jan 22	23	48	61	N	12	2425	5.9	Feb 23	5	23	40	S
4	1887	6.4	Jan 23	0	37	61	S	13	2571	6.9	Feb 24	5	51	29	N
5	190 0	7.2	Jan 23	3	16	59	S	14	2871*	7.1	Feb 26	4	40	13	S
6	1908	7.2	Jan 23	5	12	59	S	15	2876	5.4	Feb 26	5	32	12	N
7	2446	7.2	Jan 27	2	30	18	S	17	455	6.1	Mar 6	21	8	28	N
8	2456	6.2	Jan 27	3	58	17	S	18	1025	7.4	Mar 10	21	19	68	N
9	628*	4.8	Feb 9	21	49	64	N	19	2680	5.8	Mar 23	4	16	45	N

MARCH 27 TO JULY 5

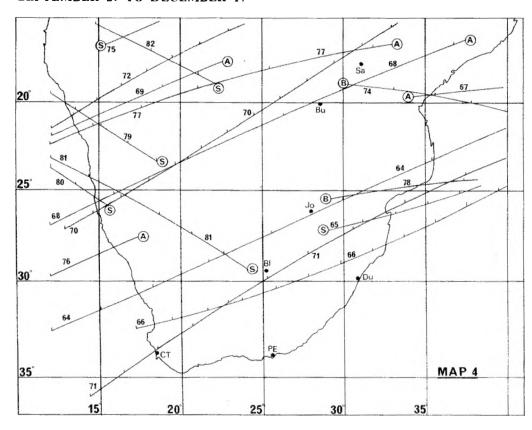


JULY 7 TO SEPTEMBER 28



No	ZC	Mag.	Date	Beg	inning	Sunlit	Limit	No	zc	Mag.	Date	Beg	inning	Sunlit	Limit
				,	. m	%						h		%	
43*	2118	2.9	Jul 7	2 '	1 59 ^m	75	S	54	2301	6.8	Aug 31	19 h	7 ^m	44	S
44*	1564	6.6	Jul 29	19	44	8	S	55	2331	6.4	Sep 1	0	2	45	S
45	1792	7.1	Jul 31	19	28	25	S	56	2463	6.9	Sep 1	22	12	56	S
47	2211	7.2	Aug 4	0	12	60	S	59	888	6.0	Sep 17	01	24	48	N
48	299	6.3	Aug 16	4	1	69	N	60	1176	7.4	Sep 19	5	18	27	S
49	423	6.4	Aug 17	6	4	59	N	61*	1410	5.3	Sep 21	4	3	11	S
50	971	7.3	Aug 21	6	26	21	N	62	2259	6.7	Sep 27	20	49	19	S
51	1886	5.7	Aug 28	19	1	13	S	63	2408	6.9	Sep 28	18	18	29	S
53	2160	6.6	Aug 30	19	55	33	S				_				

SEPTEMBER 29 TO DECEMBER 17



Nô	zc	Mag.	Date	Begin	ing Sunlit	Limit	No	ZC	Mag.	Date	Beg	inning	Sunlit	Limit
				ls.	%								%	
64	2573	7.3	Sep 29	21 ^h 2	5 ^m 41	S	74*	1429	6.8	Nov 15	2 ⁿ	25 ^m	50	S
65	2876	5.4	Oct 1	18 1	5 62	S	75	2596	7.3	Nov 23	19	39	6	S
66	2880	5.1	Oct 1	19	4 62	S	76	2789	7.3	Nov 24	21	28	13	S
67	1468	4.9	Oct 19	2 2	5 24	N	77*	3326	6.4	Nov 28	22	52	52	S
68	2686	5.2	Oct 27	21 13	3 26	S	78	3455	6.4	Nov 29	23	26	62	N
69	2690	7.0	Oct 27	22	7 26	S	79*	1410	5.3	Dec 12	5	21	74	S
70	3104	6.5	Oct 30	19 5	0 57	S	80	1519	6.5	Dec 13	5	26	64	S
71	3111	6.8	Oct 30	21 5	5 58	S	81	1744	6.5	Dec 15	4	22	43	S
72	3494	4.6	Nov 2	23 3	3 85	S	82	2000	7.1	Dec 17	5	13	21	S

EXPLANATION OF COLUMN HEADINGS IN TABLES

No. - the number of the track on the map. An asterisk denotes that the same is double - notes are given below.

Z.C. - the number of the star in the Zodiacal Catalogue.

Date

Beginning an arbitrary time of the beginning of the track in the west.

Sunlit - the percentage of the Moon's disk lit by the Sun

Limit - N = northern limit (a complete occultation takes place south of the track)

S = southern limit (complete occultation north of the track)

NOTES ON DOUBLE STARS

Track No.	ZC	
9	628	is possibly a close double star. The components are magnitude 5.0 and
		7.0 with an estimated separation of 0'.05.
14	2871	is the brighter component of the double star Aitken 12728. The companion
		is 7.6 magnitude; separation 10" in p.a. 2360,
22	3212	is the brighter component of the double star Aitken 15459. The companion
		is 10th magnitude; separation 0'!9 in p.a. 47°.
27, 74	1429	is possibly a close double star. The components are estimated to be
,		magnitude 7.6 with a separation of 0.02 in p.a. 340.
32	1925	is a spectroscopic binary.
35	58	is the brighter component of the double star Aitken 449. The companion
		is 9.7 magnitude; separation 27" in p.a. 830.
37, 77	3326	is a close double star: the components are magnitude 6.7 and 7.8 with a
,		separation of 0''.06 in p.a. 254°.
‡ I	1971	is the brightest component of the triple system Aitken 9018. One compo-
		nent is 10th magnitude; separation 1"1 in p. a. 3040. (The other compo-
		nent is fainter than 10th magnitude).
43	2118	is a close double star. The components are magnitude 3.4 and 3.8 with
		a separation of 0.01 in p.a. 1100.
4.1	1564	is the mean of a double star Aitken 7896. The components are 6,7 and
		8.8 magnitude; separation 0"5 in p.a. 180°.
61, 79	1.410	is the brighter component of the double star Aitken 7416. The companion
•		is 8.3 magnitude; separation 38" in p.a. 74°.

TIME SYSTEMS AND TELESCOPE SETTING

This section is intended to serve established amateurs and professional astronomers - i.e. those having some knowledge of time and coordinate systems. Space in the booklet does not permit full explanation, which in any case would appear complicated to the layman.

TIME SIGNALS FROM RADIO STATION ZUO

Radio signals of mean solar time are generated by the Precise Physical Measurements Division of the National Physical Research Laboratory in Pretoria and broadcast by the Post Office transmitting station at Olifantsfontein

Carrier Frequency	Radiated Power	Time of Transmission
2,5 MHz	4 Kw	2000 - 0600 SAST
5 MHz	4 Kw	Continuous
100 MHz	80 w	Continuous

The signals consist of one pulse per second, each pulse consisting of 5 cycles of $1000~{\rm Hz}~{\rm tone}$. The first pulse in every minute is lengthened to 500 milliseconds. Morse code announcements are made during the minute preceding every fifth minute. They consist of the call sign ZUO (repeated 3 times) and the Universal Time (formally known as Greenwich Mean Time) at the next minute. (A special coding indicating UT1 minus UTC is also indicated in the first 15 seconds of the minute by slightly lengthened second pulses)

SOUTH AFRICAN STANDARD TIME

South African Standard Time (as in everyday use) is mean solar time for the 300 East moridian (which runs east of Johannesburg and just west of Durban) and is exactly 2 hours ahead of Universal Time.

TIME OF SUN'S TRANSIT OVER THE 300 MERIDIAN

The table below gives the SAST when the Sun transits the $30^{\rm O}$ meridian - and a sundial on that meridian reads noon.

Jan	1	12 ^h	03 ^m	13 ^S	May	11	11^{h}	56 ^m	20 ^S	Sep	18	11^h	54 ^m	04^8
1	11	12	07	40		21	11	56	32		28	11	50	36
2	21	12	11	08		31	11	57	38	Oct	8	11	47	31
3	31	12	13	22	Jun	10	11	59	23		18	11	45	08
Feb 1	0	12	14	16		20	12	01	32		28	11	43	49
2	05	12	13	51		30	12	03	36	Nov	7	11	43	43
Mar	2	12	12	09	Jul	10	12	05	18		17	11	45	01
1	2	12	09	46		20	12	06	17		27	11	47	41
2	22	12	06	52		30	12	06	22	Dec	7	11	51	33
Apr	1	12	03	51	Aug	9	12	05	26		17	11	56	12
1	.1	12	01	01		19	12	03	31		27	12	01	11
2	21	11	58	40		29	12	00	50		31	12	03	07
May	1	11	57	03	Sep	8	11	57	36					

SIDEREAL TIME ON THE 30° MERIDIAN

			0 hra ST		t 21 hrs AST			At 0		At 2 SAS	1 hrs T			At 0 h	ra	At 21 SAST	
Jan	1	6 ^h	39 ^m	3^{h}	42 m	May	11	15 ^h	15 ^m	12 ^h	19 ^m	Sep	18	23 h	48 ^m	20 ^h	51 ^m
	11	7	18	4	22		21	15	55	12	58		28	0	27	21	30
	21	7	58	5	01		31	16	34	13	38	Oct	8	1	07	22	10
	31	8	37	5	41	Jun	10	17	13	14	17		18	1	46	22	50
Feb	10	9	17	6	20		20	17	53	14	56		28	2	26	23	29
	20	9	5 6	6	59		30	18	32	15	36	Nov	7	3	05	0	09
Mar	2	10	39	7	43	Jul	10	19	12	16	15		17	3	44	0	48
	12	11	19	8	22		20	19	51	16	54		27	4	24	1	27
	22	11	58	9	02		30	20	31	17	34	Dec	7	5	03	2	07
Apr	1	12	38	9	41	Aug	9	21	10	18	14		17	5	43	2	46
	11	13	17	10	20		19	21	50	18	53		27	6	22	3	25
	21	13	56	11	00		29	22	29	19	32		31	6	38	3	41
May	1	14	36	11	39	Sep	8	23	08	20	12						

CORRECTION FOR PLACES NOT ON THE 30° MERIDIAN

Approximate longitude corrections from the 30° East Meridian are provided below.

To find time of Sun's transit over local meridian, apply the longitude corrections to the data in the table above.

To find the sidereal times at SAST 0 hrs and SAST 21 hrs apply the corrections with the sign reversed to the data in the table.

Bloemfontein	+15 ^m	East London			+18 ^m
Bulawayo	+ 6 111	Grahamstown	+14 ^m	Pretoria	+ 7 ^m
Cape Town	+46 ^m	Johanne sburg	+ 8 ^m	Salisbury	- 4 ^m
Durban	- 4 ^m	Kimberlev	+21 ^m	Windhoek	+52 ^m

TELESCOPE SETTING

When a telescope equipped with setting circles is aimed on the meridian, its R.A. circle should read the sidereal time. Thus one can calculate the sidereal time and then set the circle, but is is usually simpler to aim the telescope at one of the well known stars given below and then to adjust the R.A. circle.

A LIST OF BRIGHT STARS FOR CHECKING TELESCOPE CIRCLES

Star	R.A.	Dec.	Mag.	Sp.	Star	R. A.	Dec.	Mag.	Sp.
Ache rnar	1h36m8	-57°21'	0,5	В5	Procyon	7h38,0	+5° 17'	0,4	F5
Aldebaran	4 34,5	+16 28	0,9	K5	Regulus	10 07,1	+12 06	1,4	B7
Rigel	5 13,3	- 8 14	0,1	B8	Spica	13 23,8	-11 02	0,9	Bl
Betelgeuse	5 53,8	+ 7 24	0,4	M2	Arcturus	14 14,5	+19 19	-0,1	K2
Canopus	6 23,4	-52 41	0,7	F0	Antares	16 27,9	-26 23	0,9	M1
Sirius	6 44,0	-16 41	-1,4	A1	Altair	19 49,5	+8 48	0,8	A7

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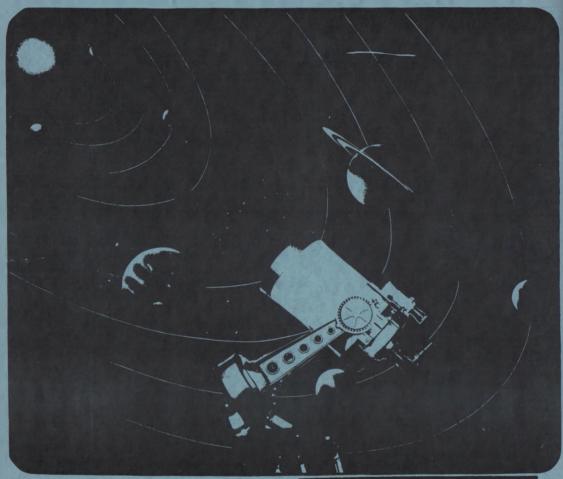
JULIAN DATE AT 1400 HOURS

	Jan	Feb	Mar	Apr	May	Jun	Jul	Aug	Sep	Oct	Nov	Dec
	244 2	2442	2442	2442	2442	2442	2442	2442	2443	2443	2443	2443
1	779	810	839	870	900	931	961	992	023	053	084	114
2	780	811	840	871	901	932	962	993	024	054	085	115
3	781	812	841	872	902	933	963	994	025	055	086	116
-4	782	813	842	873	903	934	964	995	026	056	087	117
5	783	814	843	874	904	935	965	996	027	057	088	118
6	784	815	844	875	905	936	966	997	028	058	089	119
7	785	816	845	876	906	937	967	998	029	059	090	120
8	786	817	846	877	907	938	968	999	030	060	091	121
9	787	818	847	878	908	939	969	-000	031	061	092	122
10	788	819	848	879	909	940	970	*001	032	062	093	123
11	789	820	849	880	910	941	971	*002	933	063	094	124
12	790	821	850	881	911	942	972	*003	034	064	095	125
13	791	822	851	882	912	943	973	*004	035	065	096	126
14	792	823	852	883	913	944	974	*005	036	066	097	127
15	793	824	853	884	914	945	975	*006	037	067	098	128
							• • •	000	001	001	030	120
16	794	825	854	885	915	946	976	*007	038	068	099	129
17	795	826	855	886	916	917	977	*008	039	069	100	130
18	796	827	856	887	917	948	978	*009	040	070	101	131
19	797	828	857	888	918	949	979	*010	041	071	102	132
20	798	829	858	889	919	950	980	*011	042	072	103	133
										0.12	2011	100
21	799	830	859	890	920	951	981	*012	043	073	104	134
22	800	831	860	891	921	952	982	*013	044	074	105	135
23	801	832	861	892	922	953	983	*014	045	075	106	136
24	802	833	862	893	923	954	984	*015	046	076	107	137
25	803	834	863	894	924	955	985	*016	047	077	108	138
									0.11	V	100	100
26	804	835	864	895	925	956	986	*017	048	078	109	139
27	805	836	865	896	926	957	987	*018	049	079	110	140
28	806	837	866	897	927	958	988	*019	050	080	111	141
29	807	838	867	898	928	959	989	*020	051	081	112	142
30	808		868	899	929	960	990	*021	052	082	113	143
											~ 417	470
31	809		869		930		991	*022		083		144
								V		000		134

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