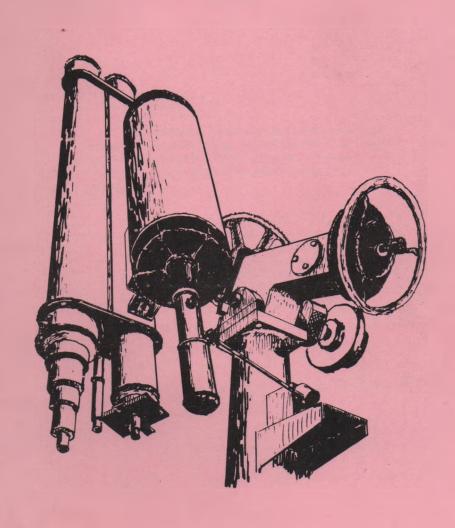
# ASTRONOMICAL HANDBOOK FOR SOUTHERN AFRICA

1979

published by the Astronomical Society of Southern Africa



# ASTRONOMICAL HANDBOOK FOR SOUTHERN AFRICA

## 1979

This booklet is intended both as an introduction to observational astronomy for the interested layman — even if his interest is only a passing one — and as a handbook for the established amateur or professional astronomer.

Front Cover: Transvaal Centre Telescopes 30cm cassegrain and 18cm and 15cm refractors.

® the Astronomical Society of Southern Africa, Cape Town. 1978

## CONTENTS

ASTRONOMY IN SOUTHERN AFRICA	
DIARY 4	
THE SUN	
THE MOON 8	
THE PLANETS	
THE MOONS OF JUPITER AND SATURN 23	
COMETS AND METEORS	
THE STARS 30	
ORDINARY OCCULTATIONS	
GRAZING OCCULTATIONS	
TIME SYSTEMS AND TELESCOPE SETTING 43	
ASSA OFFICE BEARERS	
JULIAN DATE inside back cover	

## NOTE

Unless stated otherwise, all times are SOUTH AFRICAN STANDARD TIME. In order to avoid confusion between a.m. and p.m., the 24-hour clock is used (e.g. 1800 hrs. is 6 p.m. and 2100 hrs. is 9 p.m.). Emphasis is given to phenomena visible in the evening sky - between sunset and midnight.

This handbook is produced for the Astronomical Society of Southern Africa. Most of the data it contains have been adapted for Southern Africa from the "Astronomical Ephememeria for 1979" issued jointly by the Nautical Almanac Offices of the Royal Greenwich Observatory, Herstmonceux and the U.S. Naval Observatory, Washington D.C. Additional information has also been supplied direct from the Herstmonceux Office and from the Hydrographer, South African Navy.

Thanks are due to Messrs G and C Larmuth for the preparation of much of the data.

All correspondence concerning this booklet should be addressed to the Handbook Editor, Astronomical Society of S. A 8 Glebe Road, Rondebosch 7700, Cape. Further copies can also be obtained from the same address. Enclose a postal order or cheque (in favour of the Astronomical-Society of Southern Africa) for R1 per copy.

Although every care has been taken in the compilation of the Handbook, it is distributed and sold on the explicit understanding that neither the Astronomical Society of Southern Africa nor any of its members accepts any responsibility for errors.

Dit is met spyt dat ons u meedeel dat as gevolg van beperkte fondse en produksiefasiliteite dit nie moontlik is om hierdie handboek in Afrikaans te laat druk nie.

R.F. HURLY EDITOR

# ASTRONOMY IN SOUTHERN AFRICA

As one of the few parts of the Earth having both access to the rich southern skies and a suitable climate, Southern Africa holds a favoured position in astronomy. Consequently it has seen the establishment of a number of professional observatories engaged in research while many individuals have become enthusiastic amateur astronomers. Planetaria and visiting nights at observatories convey to the general public much of what goes on in this field.

## **OBSERVATORIES**

Boyden observatory, situated at Mazelspoort, 25 km from Bloemfontein is operated by the department of Astronomy of the University of The Orange Free State. Observing facilities include the 1,5m Rockefeller Reflector. The site offers good observing conditions, without being remote from a large centre. Enquiries as to use of the observatory should be addressed to the Director.

On the other hand, the observatory sites in the hearts of Cape Town and Johannesburg have become unsuitable with the rapid expansion of those cities. The South African Astronomical Observatory - a joint venture between the South African Council for Scientific and Industrial Research and the British Science Research Council, has merged their facilities and moved the larger instruments to a new site near Sutherland in the Karroo, Observing continues at Cape Town which is also the Headquarters of the S. A. A. O. The S. A. A. O. outstation at Hartebeespoort continues to share the site with the Leiden Observatory Southern Station which has a 45cm reflector and a 25cm photographic refractor.

The Radio Astronomy Observatory at Hartebeesthoek, near Krugersdorp is used by the National Institute for Telecommunications Research for observations on extragalactic radio sources while the Rhodes University Radio Astronomy Group uses the same facilities to study Ionized Hydrogen regions associated with star formation and supernova remnants within the galaxy at a wavelength of 2,3 GHz. The Rhodes group results are reduced in Grahamstown.

In addition to the professional observatories listed above, South Africa and Rhodesia have numerous private observatories, built and operated by amateur astronomers.

## OBSERVATORIES OPEN TO THE PUBLIC

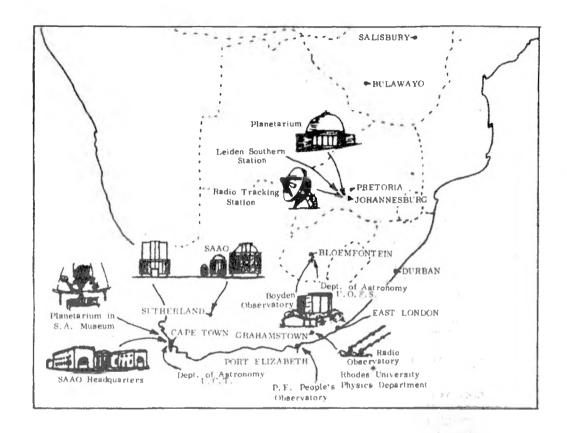
Visiting nights at Boyden Observatory are generally held around the time of first quarter. Enquiries should be made direct to the Observatory. Numbers are restricted to twenty persons on each occasion.

SAAO Headquarters, Observatory, Cape are open to visitors on the second Saturday each month at 8.00 p.m. No tickets are necessary but parties of more than ten persons should contact the observatory in advance.

## **PLANETARIA**

The major planetarium in South Africa is that situated on the grounds of the University of the Witwatersrand (entrance in Yale Road - alongside M1). It is equipped with a highly complex Zeiss projector and seats over 400 persons.

A smaller planetarium, with a Spitz projector and seating approximately 70, is located within the South African Museum, Cape Town. Shows are given each Saturday at 3.00 p.m., each Sunday at 3.30 p.m., and at 11.00 a.m. and 3.30 p.m. on public holidays. Further information can be obtained by phoning the museum at 41-2668.



## TEACHING DEPARTMENTS

Both the University of the Orange Free State and the University of Cape Town have departments of astronomy - concerned with optical astronomy - while the Physics department of Rhodes University specialises in radio astronomy. The UOFS department is incorporated into Boyden Observatory and is headed by the director, Prof. A.H. Jarrett, Prof. Briad Warner occupies the chair of Astronomy at UCT. His department makes use of the SAAO observing facilities at Sutherland. The Physics Department at Rhodes has its own radio observatory outside Grahamstown.

## THE ASTRONOMICAL SOCIETY OF SOUTHERN AFRICA

The Astronomical Society of Southern Africa is a body consisting of both amateur and professional astronomers. Membership is open to all interested persons, regardless of knowledge or experience. In addition to this handbook, the Society issues "The Monthly Notes of the Astronomical Society of Southern Africa" (MNASSA). Members also receive copies of "Sky and Telescope", an excellent and very popular monthly magazine published in the United States. It provides up to date information on both professional and amateur activities, together with news of space research and other related subjects. The Society's annual subscription is R15.00 ind there is an entrance fee of R2.50. Information can be obtained from the Honorary Secretary, Astronomical Society of Southern Africa, c/o The South African Astronomical Observatory, P.O. Box 9, Observatory 7935, Cape.

## LOCAL CENTRES OF THE SOCIETY

Autonomous local Centres of the Society hold regular meetings in Cape Town, Bloemtontein, Durban, Pietermaritzburg, Pretoria and Salisbury. Visitors are very welcome at meetings and may, if they wish, join a Centre, without becoming a full member (i.e. receiving publications for R15 subscription of the Society.

CAPE CENTRE (Cape Town) - Meetings on 2nd Wednesday of the month (except Jan. and Dec.) at the South African Astronomical Observatory at 8.00 p.m. The Centre possesses a small observatory housing the 30cm Ron Atkins Telescope. There is also an active occultation section. Secretarial address: c/o S.A.A.O. Box 9 Observatory 7935. Information on meetings also available from telephone (day time) 69-8531 ext. 256, 41-3471 ext. 286 evenings 65-6976.

TRANSVAAL CENTRE (Johannesburg) - Alternate lecture and observing meetings are held on the 2nd Wednesday of each month. There is a very enthusiastic lunar occultation group which observes grazing occultations within a radius of approximately 250 km of Johannesburg. Mirror making classes are held at suitable intervals and the centre has its own aluminising plant. Secretarial address: Box 1595, Edenvale 1610. Telephone 609 4686.

NATAL CENTRE (Durban) - Monthly meetings are held every third Wednesday at 7.45 p.m. at St. Paul's Church Hall, Church Street, Durban (near G.P.O.) and the Committee meets in private homes on the Monday evening after the general meeting. Secretarial address: Box 2704, Durban 4000. Telephone 323411 Office, 723187 Home.

NATAL MIDLANDS CENTRE (Pietermaritzburg) - Meetings are held monthly at various venues. Secretarial address 17 Yalta Rd., Pietermaritzburg 3201. Telephone 5-4038.

FREE STATE CENTRE (Bloemfontein) - Meetings are held every third Wednesday of the month. For further information contact Mr F.C. Neser telephone-22-1142, 108 Japie Neser St, Uitsig Bloemfontein.

PRETORIA CENTRE - Meetings are held on the 4th Wednesday of each month (except December) at 8 p.m., the venue alternating between the University of South Africa and the Christian Brothers' College, Silverton Road. The Centre's observatory containing a 32cm reflector is situated on the latter site. For information contact the Secretary, Mr W.R. Windisch, 932 Crots Street, Rietfontein, Pretoria 0084.

SALISBURY CENTRE - The Centre holds fairly frequent meetings, usually at eight o'clock in the evening at which talks on various subjects are given and/or films shown. In addition, social "star-gazing" sessions are arranged at intervals, at which telescopes are set up by those members who possess them and made available for observing by all members present.

The address of the Salisbury Centre is P.O. Box UA. 428, Union Avenue Salisbury and the Hon. Secretary (to whom communications should be addressed) is Mr. W.L. Stedman telephone number 5-0921.

## **OBSERVING SECTIONS OF THE SOCIETY**

These sections exist to coordinate and encourage constructive observing programmes. Mention of the type of observations and equipment involved are made in the appropriate parts of this handbook together with the names and addresses of the directors.

Comets and Meteors	see page 28
Grazing Occultations	see page 38
Nova Search Section	see page 33
Ordinary Occultations	see page 35
Variable stars	see page 34

# **DIARY 1979**

## **JANUARY**

ld	21 <sup>h</sup>	Juno 0°.2 S. of Moon
5	00	Earth at Perihelion
9	20	Aldebaran 00.5 S. of Moon
14	13	Jupiter 40 N. of Moon
15	20	Venus 80 N. of Antares
18	18	Saturn 20 N. of Moon
18	80	Venus greatest elong. W. (47°)
20	14	Mars in conjunction with Sun
22	23	Uranus 40 S. of Moon
24	17	Jupiter at opposition
25	00	Venus 20 S. of Moon
25	03	Neptune 40 S. of Moon
26	20	Venus 10.9 N. of Neptune
30	02	Juno 00.4 S. of Moon

## APRIL

2	00	Mercury 30 N. of Mars
5	20	Jupiter 50 N. of Moon
6	03	Mercury stationary
9	03	Saturn 30 N. of Moon
14	17	Uranus 40 S. of Moon
17	01	Neptune 4° S. of Moon
21	15	Mercury greatest elong. W. (270)
24	05	Venus O .3 S. of Moon
24	15	Mercury 10 S. of Moon
25	01	Mars 20 N. of Moon
29	03	Aldebaran 00.4 S. of Moon

 $1^{\alpha}$   $18^{\alpha}$  Aldebaran 0°.3 S. of Moon Occ $^{n}$ .

## **FEBRUARY**

6	02	Aldebaran Oo.3 S. of Moon Occn.
9	08	Mercury in superior conjunction
10	12	Jupiter 4º N. of Moon
13	21	Saturn 30 N. of Moon
19	07	Uranus 4° S. of Moon
21	13	Neptune 40 S. of Moon
23	17	Venus 30 S. of Moon
24	15	Uranus stationary
26	19	Total eclipse of Sun
		(not visible in S.A.)
27	14	Juno 00.5 S. of Moon
27	20	Mercury O°.6 N. of Moon

## MAY

3	08	Jupiter 40 N. of Moon
5	10	Mercury 2º S. of Mars
6	09	Saturn 30 N. of Moon
10	06	Saturn stationary
10	08	Uranus at opposition
11	23	Uranus 40 S. of Moon
14	07	Neptune 40 S. of Moon
20	08	Venue 10.1 S. of Mars
23	21	Mars 30 N. of Moon
24	00	Juno 10 S. of Moon
30	Ol	Mercury in superior conjunction
31	00	Juniter 40 N. of Moon

## MARCH

1	20	Saturn at opposition
5	09	Aldebaran Oo.2 S. of Moon Occn.
8	03	Mercury greatest elong. E. (180)
9	14	Jupiter 50 N. of Moon
12	23	Saturn 30 N. of Moon
13	23	Total eclipse of Moon
		(visible from S.A.)
14	17	Mercury stationary
18	13	Uranus 40 S. of Moon
20	20	Neptune 40 S. of Moon
21	07	Equinox
23	12	Neptune stationary
24	16	Mercury in inferior conjunction
25	11	Venus 20 S. of Moon
26	03	Jupiter stationary ·
27	04	Mars 0°.7 S. of Moon

JUN	E	
2	18	Saturn 20 N. of Moon
8	06	Uranus 40 S. of Moon
10	14	Neptune 40 S. of Moon
10	17	Neptune at opposition
20	05	Venus 50 N. of Aldebaran
21	18	Mars 50 N. of Moon
22	02	Solstice
22	14	Aldebaran 00.4 S. of Moon Occh.
23	00	Mercury 50 S. of Pollux
23	02	Venus 40 N. of Moon
26	20	Mercury 50 N. of Moon
27	16	Jupiter 30 N. of Moon
30	03	Saturn 20 N. of Moon

## JULY

**AUGUST** 

- ad ooh Earth at aphelion
- 4 00 Mercury greatest elong. E. (26°)

- 5 13 Uranus 5° S. of Moon
  7 22 Neptune 4° S. of Moon
  10 10 Mars 5° N. of Aldebaran
  17 04 Mercury stationary
- 20 Ol Aldebaran 00.3 S. of Moon Occn.
- 20 14 Mars 50 N. of Moon
- 26 17 Uranus stationary
- 27 16 Saturn 2º N. of Moon

1 22 Uranus 5° S. of Moon

10 14 Mercury stationary

18 10 Mars 50 N. of Moon

21 O4 Mercury 2º N. of Moon

29 O5 Uranus 50 S. of Moon

31 15 Neptune 40 S. of Moon

22 19 Annular Eclipse of Sun

31 19 Mercury in inferior conjunction

13 11 Jupiter in conjunction with Sun

16 Ol Aldebaran OO.2 S. of Moon Occn.

19 06 Mercury greatest elong. W. (190)

(not visible from S.A.)

25 12 Venus in superior conjunction

30 11 Mercury O°.7 N. of Jupiter

## OCTOBER (cont.)

- 18d 07h Saturn 00.7 N. of Moon Occh.
- 22 O2 Mercury 30 S. of Uranus
- 22 10 Venus 5º S. of Moon
- 22 22 Uranus 5º S. of Moon
- 23 00 Mercury 80 S. of Moon
- 25 O5 Neptune 40 S. of Moon
- 27 18 Venus O<sup>O</sup>.2 S. of Uranus
  - 29 18 Mercury greatest elong, E. (240)

# **NOVEMBER**

- 6 08 Aldebaran 00.4 S. of Moon Occn.
- 8 22 Mercury 2º S. of Venus
- 9 20 Mercury stationary
- 11 16 Venus 40 N. of Antares
- 12 12 Mars 30 N. of Moon
- 12 17 Regulus 10 N. of Moon
- 13 09 Jupiter 00.8 N. of Moon Occn.
- 14 20 Saturn Oo.3 N. of Moon
- 17 19 Mars 10.6 N. of Regulus
- 20 06 Mercury in inferior conjunction
- Venus 2º S. of Neptune 20 07
- 21 13 Neptune 4º S. of Moon
  - 21 16 Venus 60 S. of Moon 25 04 Mercury 10.7 N. of Uranus
  - 29 12 Mercury stationary

## **SEPTEMBER**

- 6 11 Total eclipse of moon (not visible in S.A.)
- 10 16 Saturn in conjunction with Sun
- 12 13 Aldebaran O<sup>O</sup>.2 S. of Moon 13 O7 Mercury in superior conjunction 15 Ol Mars 6<sup>O</sup> S. of Pollux
- 16 O5 Mars 50 N. of Moon
- 17 O3 Vesta stationary
- 19 00 Jupiter 20 N. of Moon
- 23 19 Equinox
- 25 13 Uranus 5º S. of Moon
- 26 15 Jupiter 00.3 N. of Regulus
- 27 22 Neptune 4º S. of Moon

## **DECEMBER**

- 3 18 Aldebaran OO.5 S. of Moon
- 5 02 Mercury 20 N. of Uranus
- 7 18 Mercury greatest elong. W. (210)
- 10 Ol Regulus O.8 N. of Moon Occh.
- 10 20 Mars 20 N. of Moon
- 10 22 Jupiter 00.4 N. of Moon
- 12 O7 Saturn OO.Ol S. of Moon
- 12 22 Neptune in conjunction with Sun
- 13 19 Mars 10.7 N. of Jupiter
- 16 20 Uranus 50 S. of Moon
- 17 22 Mercury 40 S. of Moon
- 19 00 Mercury 60 N. of Antares
- 21 19 Venus 50 S. of Moon
- 22 13 Solstice
- 27 O7 Jupiter stationary 27 O9 Mercury 10.4 S. of Neptune
- 31 O2 Aldebaran O0.4 S. of Moon

## **OCTOBER**

- 14 Mercury 10.9 N. of Spica
- 5 09 Venus 3º N. of Spica
- 9 22 Aldebaran OO.3 S. of Moon Occn.
- 14 22 Mars 40 N. of Moon 16 17 Jupiter 10 N. of Moon

# **THE SUN 1979**

## **BASIC DATA**

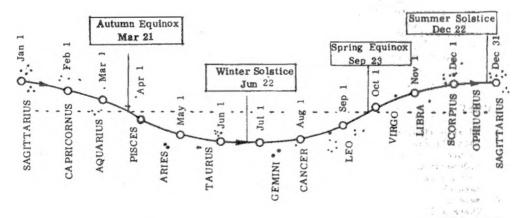
Diameter: 1 392 000 km (109 times Earc

Diameter: 1 392 000 km (109 times Earth diameter) Mass: 1,99 x 10<sup>30</sup> kg (330 000 times Earth Mass)

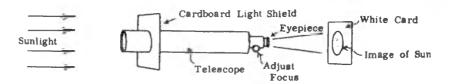
Surface Temperature: Approx. 6000°C Temperature at centre: Approx. 10 million°C

The Sun is our nearest star. It is composed chiefly of hydrogen and is in a gaseous state throughout. So hot and dense is its interior that nuclear reactions occur there - thus producing the energy that is eventually radiated from its surface. At times its surface is disturbed by sunspots (which may persist for some weeks) and flares (short lived).

The Earth's orbit around the Sun is not quite circular. In 1978 we will be closest to the Sun on January 2 (perihelion - approx. distance 147 million km.) and furthest from the Sun on July 5 (aphelion - approx. 152 million km.) During the year, the Sun appears to us to make a complete circuit of the sky (i.e. relative to the starry background) as indicated in the diagram.



Permanent damage to the eye can be caused by looking directly at the Sun. The diagram below shows how a small telescope (or half a binocular) may be used to project an image of the solar disk onto a piece of white card. It may also be advisable to stop down the telescope aperture so that the eyepiece is not damaged by the intense light passing through it. Tiny black sunapots are generally visible on the otherwise white solar disk - if monitored over a period of a week or so, the rotation of the Sun should be apparent.



## TIMES OF SUNRISE AND SUNSET

		CAPE	TOW	N		DUR	BAN		В	LOEME	ONT	EIN	JC	HANN	ESBU	RG	:			
	SI	NRISE	SU	SET	SU	MRISE	SU	NSET	SU	VRISE	SUNSET		SUNRISE		SUN	SET	SUN	RISE	SUNSET	
Jan 1	os <sup>h</sup>	34 <sup>m</sup>	20 <sup>h</sup>	n1 m	-94 <sup>h</sup>	ã8 <sup>™</sup>	19 <sup>h</sup>	01m	05 <sup>h</sup>	21 <sup>m</sup>	19 <sup>h</sup>	18 <sup>m</sup>	os <sup>h</sup>	18 <sup>m</sup>	19 <sup>h</sup>	04 <sup>m</sup>	05 <sup>h</sup>	24 <sup>m</sup>	18 <sup>h</sup>	
11	05	46	20	02	05	06	13	02	05	29	19	18	05	25	19	05	05	29	16	37
21	05	55	19	59	05	14	19	00	05	37	19	17	05	33	19	04	05	37	18	38
Feb 1	06	07	19	52	0.5	24	18	55	05	46	19	13	05	42	19	00	05	42	18	36
11	06	17	19	44	05	32	18	48	0.5	54	19	06	05	49	18	55	05	47	18	32
21	06	26	19	33	05	41	18	39	06	02	18	57	05	56	18	47	.05	52	18	27
Mar 1	06	33	19	23	05	16	18	30	06	08	18	48	06	00	18	39	05	55	18	21
11	06	41	19	11	05	53	18	19	06	13	18	38	06	06	18	29	05	57	18	15
21	06	49	18	Sá	0.5	59	15	On	06	16	18	27	06	11	18	19	06	00	18	06
Apr 1	06	58	18	41	06	96	1.7	571	06	25	18	13	06	17	18	06	06	02	17	57
11	07	94	16	30	06	11	1.	43	90	30	18	03	06	21	17	56	06	04	17	50
21	97	15	l e	17	06	17	1.7	31	06	35	17	52	06	25	17	4.7	06	07	17	43
May 1	07	20	18	05	06	24	17	22	06	42	17	44	06	31	17	38	06	10	17	37
11	07	24	17	57	06	31	17	14	06	49	17	36	06	37	17	31	06	13	17	32
21	07	34	17	50	06	36	17	06	06	54	17	30	06	41	17	26	06	16	17	29
Jun 3	07	43	17	45	06	43	17	04	07	01	17	27	06	47	17	23	0-6	20	17	28
11	07	46	17	44	06	48	17	03	07	05	17	26	06	52	17	22	06	23	17	27
21	07	51	17	44	90	51	17	04	07	08	17	27	06	55	17	24	06	26	17	29
Jul 1	07	53	1 4	48	06	53	17	07	07	10	17	30	06	57	17	27	06	27	17	32
11	07	51	1.7	52	06	51	17	11	07	08	17	34	06	55	17	30	06	27	17	35
21	07	47	17	58	06	48	17	16	0?	05	17	39	06	53	17	35	06	26	17	40
Aug 1	07	39	18	06	06	42	17	22	07	00	17	45	06	48	17	41	06	23	17	42
11	07	30	18	13	06	34	17	29	06	53	17	51	06	41	17	46	06	18	17	46
21	07	19	18	20	06	24	17	35	06	42	17	55	06	32	17	50	06	11	17	48
Sep 1	07	06	18	27	06	12	17	40	06	31	38	01	06	21	17	54	06	04	17	49
- 11	06	52	18	34	06	00	17	16	06	19	18	06	06	21	17	59	05	55	17	51
21	06	38	18	41	05	48	17	51	06	07	18	10	05	59	18	03	05	16	17	52
Oct 1	06	25	18	48	05	37	17	57	05	57	18	16	05	50	18	0.8	05	39	17	54
11	06	12	18	55	05	25	18	03	05	45	18	22	05	39	18	12	05	30	17	57
21	05	58	19	04	05	12	18	09	05	33	18	27	05	27	18	17	05	23	17	59
Nov 1	05	46	19	13	0.5	02	18	17	05	24	18	35	05	19	10	24	05	16	18	03
11	05	38	19	23	0.4	55	18	26	05	17	18	44	05	13	18	32	05	14	18	08
21	05	31	19	33	04	49	18	34	05	12	18	52	05	08	18	39	05	11	18	13
Dec 1	05	29	19	43	04	48	18	12	05	11	19	00	05	07	18	46	05	12	16	19
11	05	28	19	50	04	48	18	50	05	11	19	07	05	08	18	53	05	14	18	25
21	05	32	19	57	04	52	3.6	57	05	15	19	14	05	12	19	00	05	18	18	31
									_									10	40	21

## SOLAR ECLIPSES

The Total Eclipse of the Sun on February 26th will be seen only from North America and Greenland. Mid eclipse will be at  $19^h22^m$ .

The Annular Eclipse of the Sun on August 22 will be seen only from South America and Antarctica. Mid eclipse will be at  $19^h53^m$ .

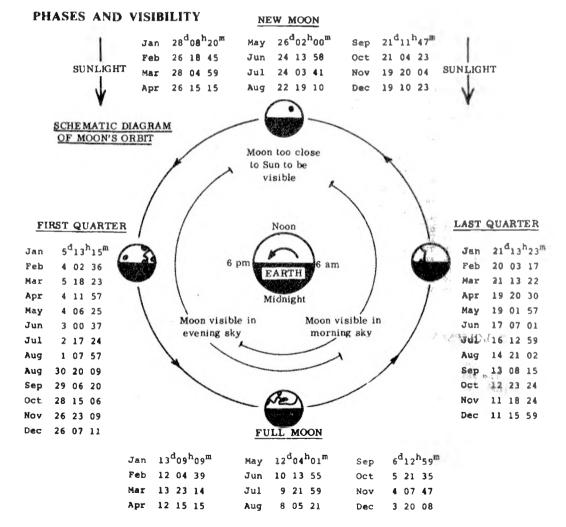
## **THE MOON 1979**

## **BASIC DATA**

Diameter: 3 480 km (0,27 of Earth) Mass:  $7,35 \times 10^{22}$  kg (1/81 of Earth)

Surface Gravity: 0,16 of Earth

Average distance from Earth: 384 000 km



### THE MOON'S ORBIT

Dates of Apogee, when the Moon is furthest from the Earth (approx. 407 000 km) and of Perigee, when the Moon is closest to the Earth (approx. 357 000 km) are given below.

	MO	TA NO	PERI	GEE		TA NO	APOG				
Jan	28	May	18	Oct	4	Jan	15	May	5	Sep	19
Feb	26	Jun	13	Nov	1	Feb	11	Jun	1	Oct	16
Mar	26	Jul	11	Nov	29	Mar	10	Jun	29	Nov	13
Apr	23	Aug	8	Dec	23	Apr	7	Jul	27	Dec	11
		Sep	6					Aug	23		

As a result of its motion around the Earth, the Moon appears to make a complete circuit of the heavens in just under a month. It occasionally passes in front of bright stars (details given in Occultation section - page 35) and close to visible planets (details given in Diary pages 4 and 5).

## TIMES OF MOONRISE AND MOONSET

Times for Bloemfontein, Cape Town, Durban, Johannesburg and Port Elizabeth can be obtained from the tables on pages 11 to 16.

## LUNAR ECLIPSE

There will be a partial eclipse of the Moon visible from Southern Africa on March 13/14.

Moon enters penumbra	13 <sup>0</sup> 20 <sup>n</sup> 11.7
Moon enters umbra	13 21 29.7
Middle of eclipse	13 23 08.8
Moon leaves umbra	14 00 08.1
Moon leaves penumbra	14 02 05.9

Magnitude of the eclipse 0.858. First contact of Umbra with Limb of Moon is at 139° East of North point of Moon. Last contact of Umbra with limb of Moon at 104° West of North point of Moon.

The total eclipse of the Moon on September 6th will be about local noon and hence will not be visible from Africa.

## THE SURFACE OF THE MOON

In common with the inner planets of our solar system, the Moon's surface suffered bombardment by numerous minor bodies during period 4,5 to 3,0 billion years ago. This has produced the heavily cratered topography now visible. Some particularly large impacts caused large circular depressions, which were flooded by molten lava from the Moon's interior. These are the maria basins which appear smoother and darker then the rest of the surface (the latin words mare and maria come from older times when they were mistaken for seas). The maria surfaces, being younger have fewer large craters, but the entire surface is peppered with tiny craters produced by tiny bodies which have also served to plough up the ground thus forming the regolith - a layer of loose material a metre or so deep.



## LIBRATIONS



Jan 25 Feb 21 Mar 21 Apr 17 May 15 Jun 10 Jul 7 Aug 4/31 Sep 27 Oct 24 Nov 20 Dec 17



Jan 11 Feb 7 Mar 6 Apr 2/30 May 27 Jun 23 Jul 20 Aug 16 Sep 12 Oct 10 Nov 7 Dec 3

Dates of Maximum Exposure of Indicated Limbs

Jan 22 Feb 19 Mar 18 Apr 15 May 11 Jun 8 Jul 6 Aug 3/31 Sep 28 Oct 25 Nov 20 Dec 5

Jan 7 Feb 4 Mar 4 Apr 1/29 May 26 Jun 22 Jul 19 Aug 15 Sep 12 Oct 11 Nov 8 Dec 17





# The Moon 1979

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# CAPE TOWN – TIMES OF MOONRISE AND MOONSET FOR PORT ELIZABETH SUBTRACT 28 MINUTES

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# CAPE TOWN — TIMES OF MOONRISE AND MOONSET FOR PORT ELIZABETH SUBTRACT 28 MINUTES

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# The Moon 1979

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# DURBAN – TIMES OF MOONRISE AND MOONSET FOR BLOEMFONTEIN ADD 19 MINUTES

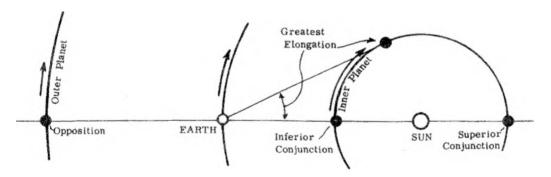
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		2 <sub>m</sub>			•	S	EO.	<b>~</b>	_		×	7	1	~		٠,	.0	7	~		_			7	_	-	м			_	_	44
	Set	$3^{h}$ 2		-	0	0	03 0	0	0	0	0	5	4	M	-	2	543	-	N	4	7	-	0	S	4	4	₩)	7	-	0	0	5
JULY		III 2															-															
7		h.	4	2	S	M	4 19	0	0	0	0	-	7	7	M	М		М	M	M	7	2	_	0	4	7 26	0	M	-	47	7	0 55
	×					-		-	=	-	1			7				ŏ	0	0	0	ŏ	õ	ŏ	ŏ	0 7	0	õ	90	50	1	10
		_	7	M	₹	S	9	7	00	6	0	_	7	~	4	S	9	7	<b>6</b>	6	0	_	2	2	4	S	و	7	00	6	0	-
											-	_	-	-	-	-	-	-	-	-	(.4	~	N	7	4	N	171	~	2	N	14.]	M

# THE PLANETS 1979

		BAS	SIC DATA			
	Dist from Sun	Period of Revolution	Mass	Diameter	Rotation Period	Inclination of Equator
	10 <sup>6</sup> km	уеагв	(Earth = 1)	10 <sup>3</sup> km		to Orbit
Mercury	58	0,24	0,056	4,98	594	?
Venus	108	0,62	0,817	12,4	244 . h . m	23027'
Earth	150	1,00	1,000	12,8	23"56"	23 27'
Mars	228	1.88	0,108	6,76	24 37	23 59
Jupiter	778	11,9	318,0	142,7	09 51	03 04
Saturn	1426	29,5	95,2	120,8	10 14	26 44
Uranus	2868	84.0	14,6	47,1	10 49	97 53
Neptune	4494	164.8	17,3	44,6	14 ?	28 48
Pluto	5896	247.6	0,9?	?	6d?	?

## **GENERAL**

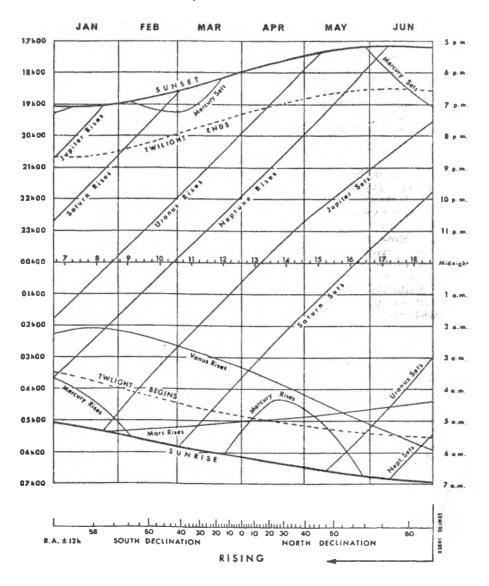
Apart from Uranus, Neptune and Pluto, the planets of our solar system are amongst the brightest objects in the night sky. Unlike the distant stars, their relative positions do not remain fixed, but continually change as, like the Earth, they orbit around the Sun. Their apparent movements against the starry background are complicated as they result from a combination of their own motion and the Earth's motion. Their brightnesses also vary considerably, as both their distances from the Earth and the visible portions of their sunlit hemispheres change. Since the period of a planet increases with increasing distance from the Sun, so we find that the inner planets - Mercury and Venus - appear to "overtake" the Earth in their orbits, while the Earth in turn "overtakes" the outer planets - Mars, Jupiter and Saturn. The terms given in astronomy to the various Sun-Earth-Planet configurations are illustrated in the accompanying diagram. Dates of such configurations occuring in 1978 are listed chronologically in the Diary (pages 4 and 5) and are also mentioned in the text below.



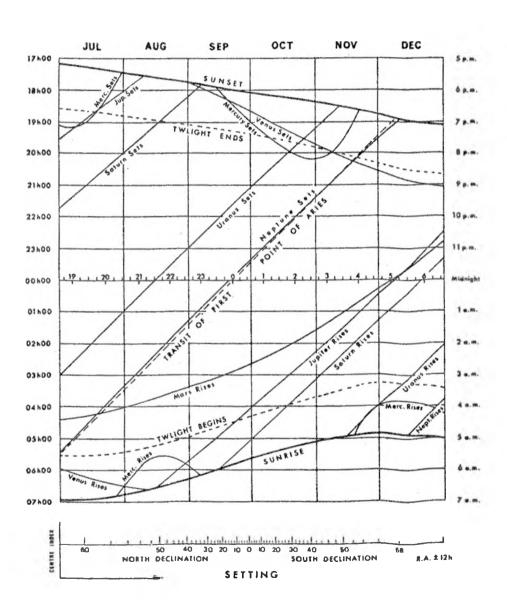
## TIMES OF RISING AND SETTING

The times of rising and setting given by the diagram are accurate for position 30°East. 30° South, and approximately correct for other places in Southern Africa. Strictly speaking, corrections for latitude and longitude should be applied, but the latitude correction is, in general sufficiently small to be ignored and in no case will exceed 15 minutes. Longitude corrections are given on page 44.

The scales at the bottom of these pages are for finding rising or setting times of any object of which the right ascension and declination are known. Set dividers or a strip of paper from the index at the centre of the scale to the object's declination and in the direction desired for either



rising or setting. Measure this same distance and direction along the midnight line, beginning at the objects right ascension indicated by the numerals. (Should this end point fall outside the chart, 12 hours should be added to or subtracted from the right ascension. Reset the dividers using the end of the scale instead of the centre index, and measure in the opposite direction to that first used). Through the point established draw a line parallel to the First Point of Aries transit line (indicated by the dashed line on the chart).



## OBSERVING THE PLANETS

To the naked eye, planets appear as virtually pinpoint sources of light. However, their disks can be readily resolved with the aid of a small telescope. Even so, their angular diameters are of the order of 10 seconds of arc - roughly 1/200 of the Moon's angular diameter - so it is not always possible to distinguish details on their disks. The disks of Mercury and Venus are only seen fully illuminated when they are furthest from us - as they draw closer, their disks grow larger but the phase changes to a crescent as we see more of their dark hemispheres. In contrast, the disks of the outer planets are always seen fully or near fully illuminated.

## **MERCURY**

The innermost planet, Mercury, resolves around the Sun faster than any of the other planets. Being close to the Sun it can only be seen just after sunset or just before sunrise, when it is near greatest elongation (greatest angle between Mercury and Sun as seen from the Earth). Except when in transit, it can never be seen near inferior conjunction (passing between Earth and Sun) or near superior conjunction (passing round the far side of the Sun). The angular diameter of Mercury's disc rarely exceeds 10 seconds of arc so it is difficult to make out any features on the disc, but phases (like those of the Moon) might just be visible. Close up photography by the Mariner 10 spacecraft has shown that the planet's surface is incredibly similar to that of the moon.

		d	h		d	h		d	h
Superior conjunction	Feb.	9	08	May.		01	Sept.	13	07
Greatest elongation E	Mar.	8	03(18°)	July	4	00(26)	Oct.	29	18(24°)
Stationary	Mar.			July	17	04	Nov.		20
Inferior conjunction	Mar.	24	16	July	31	19	Nov.	20	06
Stationary	Apr.		03	Aug.	10		Nov.	29	
Greatest elongation W	Apr.	21	15(27)	Aug.	19	06 (19°)	Dec.	7	18(210)

## **VENUS**

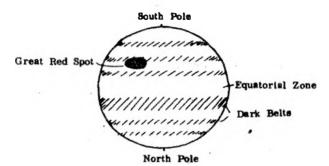
Venus will be seen in the morning sky till June but after that will be too close to the sun for viewing till about the beginning of October after which it will be in the evening sky till the end of the year. Its apparent diameter will change from 31"in January to 10" in August and to 12" at the end of the year. It will be brightest -4.3 at the beginning of January and -3.5 in early August.

## MARS

Mars is too close to the sun for viewing until about April and is a morning star for most of the rest of the year. Its magnitude will stay at about +1.4 till October and it will brighten to +0.3 at the end of December, while its apparent diameter veries from 4" in January to 10" in December and its distance from the earth varies between 2.4 AU in January and 1.0 AU in December.

## **JUPITER**

Jupiter is a prominent object in the evening sky till July and from late August it will be a morning sky object, moving into the evening sky in December. It is at its brightest (magnitude -2)in January and December. Because of its large angular size (43 seconds of arc at opposition), Jupiter makes an excellent object for a small telescope. It is often possible to see features on the disc: dark and light cloud bands running parallel to the equator and spots, in particular the famous Great Red Spot. These are indicated in the diagram below. The Great Red Spot is not always visible because of the rotation of the planet. Also clearly visible are four of Jupiter's fourteen moons. An entire section of this Handbook is devoted to the movements of these satellites and the Phenomena associated with them (see page 23).



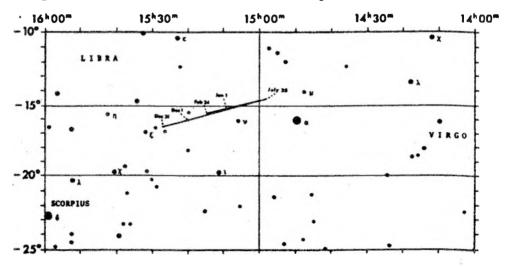
## **SATURN**

Saturn, in the constellation of Leo till late October and in Virgo thereafter, will be clearly seen in the evening sky till August and in the morning sky from October until the end of the year. Unfortunately it will never be very high above the horizon because of its northerly declination. It is at its greatest brightness (magnitude 0.5) at opposition on March 1. The diagram below shows its appearance through a small telescope (the scale is the same as for the Jupiter diagram) - including the spectacular ring system. The angle of the ring varies from 4° in January to 7° in May and to being end-on on October 27 after which the angle will open up to 2° by the end of the year, with the north side of the rings then visible.



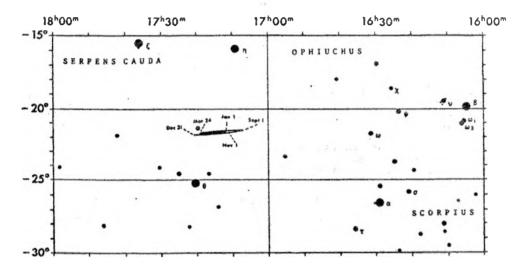
## **URANUS**

Uranus, at opposition on May 10 and conjunction on November 14, is in the constellation of Libra all year. It is on the border of naked eye visibility (magnitude 5.7 at opposition) and may be found with the aid of binoculars and references to the accompanying finding chart, which shows all of the stars in the region down to the same faintness. With a small telescope, its disc (angular diameter 4 seconds of arc) may just be distinguished. Uranus is now known to have rings similar to those of Saturn.



## **NEPTUNE**

Neptune lies in the constellation of Ophiuchus close to the conspicuous constellation of Scorpius. It is far too faint to be seen with the naked eye - magnitude 7,7 at opposition on June 10 - but may be located using the finding chart below (which shows all stars down to magnitude 7,7) and a small telescope. Its angular diameter is only 2,5 seconds of arc, but its non-stellar appearance should aid in identification.



## **PLUTO**

During the year Pluto moves in the constellation of Virgo between the coordinates  $RA~13^h26^m$  and  $13^h44^m$  and  $Dec~+7^0.31^t$  and  $+9^0.47^t$ . Since it is very faint, magnitude 14, it can only be found using a large telescope and specially prepared finding charts.

May	26	18 <sup>h</sup>	44 <sup>m</sup>	III	Tr	E	Jun	6	19 <sup>h</sup>	4 7 <sup>m</sup>	I	Sh	E	Jun	29	178	01 <sup>m</sup>	I	Tr	I
		19	37	III	Sh	I		7	17	08	1	Ec	R			17	25	IV	Oc	R
		20	40	IV	0c	D		10	18	05	II	Tr	I			17	44	1	Sh	I
	27	17	37	II	Sh	Ε			19	58	II	Sh	I			19	19	I	Tr	E
	28	20	02	I	Tr	1			20	54	II	Tr	E			19	22	IA	Ec	D
		21	06	I	Sh	I		12	17	03	II	Ec	R		30	17	21	1	Ec	R
	29	17	22	I	Oc	D		13	17	40	III	0c	R	Jul	5	17	01	II	Sh	I
		20	45	I	Ec	R			17	45	III	Ec	Ø			18	37	-II	Tr	E
	30	17	52	I	Sh	E			18	30	I	Tr	I		8	17	15	III	Tr	I
Jun	1	20	07	II	Oc	D			19	25	I	Sh	I							-
	2	19	25	III	Tr	I		14	19	03	I	Ec	R	Dec	15	00	21	I	Tr	E
	3	17	22	II	Sh	1		19	19	40	II	Ec	R		17	00	59	III	Ec	D
		18	09	II	Tr	Ε		20	18	23	III	00	D		20	23	14	III	Tr	E
		20	12	II	Sh	E		21	17	52	1	Oc	D		22	23	20	1	0c	R
	4	17	04	IV	Sh	I		22	17	19	I	Tr	E		27	00	56	II	Tr	E
	5	19	22	I	Oc	D			18	07	I	5ħ	E			22	29	III	Sh	E
	6	17	25	III	Ec	R		26	17	51	II	0c	D			23	30	III	Tr	I
		17	30	I	Sh	1		28	17	16	II	Sh	E		29	00	40	1	Sh	I
		18	48	I	Tr	E			19	52	I	0c	D		30	22	28	I	Tr	E

# THE MOONS OF JUPITER AND SATURN 1979

## JUPITER'S MOON

One of the most popular sights for an observer with a small telescope is Jupiter and its moons. Four of Jupiter's fourteen moons are large enough to be seen very easily - they would even be visible to the naked eye were it not for the glare of the mother planet. In order of increasing distance from Jupiter, the four moons are Io, which orbits once around Jupiter in less than 2 days; Europa,  $3\frac{1}{2}$  days; Ganymede, 7 days; and Callisto which takes 17 days for a full circuit. All the orbits lie in Jupiter's equatoria! plane and the system is seen almost edge on. As the moons circle Jupiter, they appear to oscillate from side to side alternatively passing in front of and behind the planet. Their configurations change from night to night and are shown in the diagrams on pages 26 and 27.

# EVENTS RELATED TO THE MOONS PASSING IN FRONT OF AND BEHIND JUPITER

The table below lists all events occurring between the end of twilight and just after midnight when the planet is above the horizon in Southern Africa. Explanation of table:

Date and predicted times are given - these are for mid-phenomena and are not instantaneous.

The moons concerned are 1 - lo

III - Ganvmede

II - Europa

IV - Callisto

Phenomena - the abbreviations used are:

Ec.	-	Eclipse: the satellite passes through the shadow of Jupiter	D	-	Disappearance
Oc.	-	Occultation: the satellite is obscured by the disk of Jupiter	$\mathbf{R}$	-	Reappearance

Tr. - Transit: the satellite crosses the disk of Jupiter I Ingress
Sh. - Shadow transit: the shadow of the satellite transits the disk E Egress

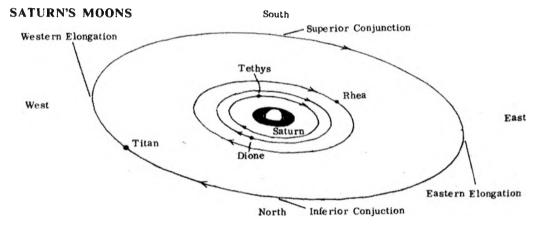
Jan	1	21 <sup>1</sup>	<sup>n</sup> eo <sup>r</sup>	II	Sh	I	Jan 1	.0 2	21	101 <sup>m</sup>	II	0c	R	Jan	21	22	13 <sup>m</sup>	IV	Tr	E
		22	16	II	Tr	I		2	2	07	I	Sh	E		24	23	38	I	Tr	I
		23	29	1	Sh	I		2	2	28	1	Tr	E			23	39	I	Sh	I
	2	00	01	II	Sh	Ε	1	7 0	0	29	I	Ec	D			23	39	II	0 <b>c</b>	D
		20	40	I	Ec	D		1	9	48	III	Sh	I			23	44	III	Tr	I
		23	31	I	Oc	R		2	0	29	III	Tr	I			23	46	III	Sh	I
	3	19	47	II	0¢	R		2	ı	06	II	Ec	D		25	20	50	Í	Oc	D
		20	14	I	Sh	E		2	L	45	I	Sh	I			23	11	I	Ec	R
		20	44	I	Tr	E		2	Ł	55	ı	Tr	I		26	20	20	I	rr	E
	4	22	58	IV	Sh	I		2	3	22	III	Sh	E			20	23	I	Sh	E
	8	23	45	II	Sh	I	1	8 0	0	01	I	Sh	E			21	03	II	Tr	E
	9	00	32	II	Tr	I		0	0	04	III	Tr	E			21	10	II	Sh	E
		22	35	I	Ec	D		0	0	11	I	Tr	E		29	23	34	IV	0c	D
1	0	19	22	III	Sh	E		0	0	14	II	Oc	R	Feb	1	22	34	1	0c	D
		19	51	I	Sh	E		1	В	58	I	Ec	D		2	19	48	I	Tr	I
		20	12	I	Tr	I		2	L	25	1	Oc	R			20	01	I	Sh	I
		20	47	III	Tr	E	2	1 2	L	31	IV	Sh	E			20	25	11	Tr	I

ine moons of Jupiter and Saturn 19/9

```
7eb 2 20<sup>h</sup>54<sup>m</sup>
                              Mar 6 21h 43m
                                             II Tr E
                II Sh I
                                                      Apr 6 1747
                                                                             IV Oc R
        22 04
                                      23 35
                                              II Sh E
                I Tr E
                                                                      18 22
                                                                              I Ec R
        22 17
                 I Sh E
                                   8 19 53 III Tr I
                                                                     19 15
                                                                            III Sh E
        23 19
                                      23 26 III Tr E
                                                                  7 17 59
              II Tr E
                                                                             II Tr I
        23 49
                II Sh E
                                      23 41 III Sh I
                                                                     20 30
                                                                              II Sh I
     3 19 34
                                   11 23 03
                                              I Tr I
                 I Ec R
                                                                     20 50
                                                                              II Tr E
    4 21 26
               III Ec R
                                   12 00 02
                                              I Sh I
                                                                     23 21
                                                                              II Sh E
     9 00 19
                                      18 13
                                             IV Tr E
                                                                  9 17 34
                 I Oc D
                                                                              II Ec R
       21 32
                 I Tr I
                                      20 19
                                              I Oc D
                                                                 10 00 43
                                                                            III Oc D
       21 55
                                      22 57
                                             IV Sh I
                 I Sh I
                                                                 11 00 56
                                                                               I Tr I
       22 42
                                              I Ec R
                II Tr I
                                      23 38
                                                                     22 13
                                                                              I Oc D
       23 31
               II Sh I
                                   13 18 30
                                              I Sh I
                                                                 12 19 24
                                                                              I Tr I
        23 49
                                      19 46
                 I Tr E
                                             I Tr E
                                                                     20 39
                                                                              I Sh I
    10 00 12
                 I Sh E
                                      20 46
                                              I Sh E
                                                                     21 39
                                                                              I Tr E
        21 29
                 I Ec R
                                      21 16
                                              II Tr I
                                                                     22.55
                                                                               I Sh E
    11 20 02
              III Oc D
                                      23 19
                                              II Sh I
                                                                 13 18 05
                                                                            III Tr E
        20 53
                                  14 00 07
                                              II Tr E
               II Rc R
                                                                     19 39
                                                                             III Sh I
    15 18 54
                IV Ec D
                                   15 20 30
                                              II Ec R
                                                                     20 17
                                                                               I Ec R
        23 33
                                      23 27 III Tr I
               IV Ec R
                                                                     23 14 III Sh E
    16 23 16
                I Tr I
                                   19 00 53
                                               I Tr I
                                                                 14 17 23
                                                                              I Sh E
                                            III Ec R
        23 50
                I Sh I
                                      21 25
                                                                     20 33
                                                                            II Tr I
    17 00 59
                II Tr I
                                      22 09
                                              I Oc D
                                                                     23 07
                                                                             II Sh I
        20 30
                I Oc D
                                  20 19 20
                                              I Tr I
                                                                     23 11
                                                                             IV Tr I
        23 24
                 I Ec R
                                      20 25
                                              I Sh I
                                                                     23 24
                                                                             II Tr E
    18 19 28
                II Oc D
                                      20 27
                                             IV Oc D
                                                                 16 20 10
                                                                             II Ec R
        19 59
                 I Tr E
                                            I Tr E
                                      21 36
                                                                 19 21 18
                                                                              I Tr I
        20 34
                 I Sh E
                                      22 41
                                               I Sh E
                                                                     22 34 I Sh I
        23 23
              III Oc D
                                      23 42
                                             II Tr I
                                                                     23 34
                                                                              I Tr E
        23 27
                II Ec R
                                  21 20 02
                                              I Ec R
                                                                 20 18 29
                                                                           III Tr I
    22 19 17
               III Sh E
                                  22 18 02
                                              II Oc D
                                                                     18 37
                                                                              I Oc D
    23
       22 20
               IV Tr I
                                      23 05
                                              II Ec R
                                                                     22 02
                                                                            III Tr E
        22 16
                I Oc D
                                  24 18 07
                                              II Sh E
                                                                     22 12
                                                                              I Bc R
    25 19 29
                I Tr I
                                  26 20 40 III Oc R
                                                                     23 38 III Sh I
        20 13
                 I Sh I
                                      21 48
                                             III Ec D
                                                                 21 18 02
                                                                             I Tr E
        21 44
                I Tr E
                                      23 59
                                               I Oc D
                                                                     19 18
                                                                              I Sh E
        21 45
                                  27 21 11
               II Oc D
                                               I Tr I
                                                                     23 10
                                                                            II Tr I
        22 29
                I Sh E
                                      22 20
                                              I Sh I
                                                                 23 17 22
                                                                             II Oc D
    26 19 48
                I Ec R
                                      23 26
                                              I Tr E
                                                                     19 08
                                                                             IV Ec D
    27 19 22
                II Tr E
                                  28 00 36
                                              I Sh E
                                                                     22 46
                                                                             II Ec R
        20 58
                II Sh E
                                      18 27
                                              I Oc D
                                                                 24 17 26 III Ec R
War 1 19 41
               III Sh I
                                              I Ec R
                              Mar 28 21 57
                                                                 25 17 53
                                                                             II Sh B
        19 56
               III Tr E
                                  29 17 54
                                              I Tr E .
                                                                 26 23 13
                                                                             I Tr I
        23 16
               III Sh B
                                      19 05
                                                                 27 20 32
                                              I Sh E
                                                                              I Oc D
     4 00 03
                 I Oc D
                                              II Oc D
                                      20 29
                                                                     22 30
                                                                           III Tr I
        21 15
                 I Tr I
                                                                 28 17 42
                                      21 41
                                              IV Sh E
                                                                              I Tr I
        22 07
                 I Sh I
                                  31 17 52
                                              II Sh I
                                                                     18 57
                                                                              I Sh I
        23 31
                 I Tr E
                                      18 17
                                              II Tr E
                                                                     19 58
                                                                              I Tr E
     5 00 04
                                      20 44
                                              II Sh E
                                                                 29 18 36
               II Oc D
                                                                              I Ec R
        00 23
                              Apr 2 20 51
                                             III Oc D
                                                                 30 19 59
                I Sh E
                                                                             II Oc D
                I Oc D
                                   3 00 27
                                             III Oc R
                                                             May 1 17 47
                                                                            III Ec D
        18 30
                                      23 03
                                               I Tr I
                                                                     21 26 III Ec R
        21 43
                I EC R
                                   4 00 15
                                              I Sh I
                                                                     21 53
                                                                             IV Tr E
     6 18 51
                II Tr I
                                      20 20
                                              I Oc D
                                                                  2 17 39
        18 52
                I Sh E
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                                      23 53
                                               I Ec R
        20 42
                II Sh I
                                                                     17 58
                                                                             II Tr E
```

Ma	y		20 <sup>h</sup>	30 <sup>m</sup>	II	Sh	E	May 9	2 0 <sup>‡</sup>	, 38 <sub>tu</sub>	ıτ	Tr	E	May	16	20 <sup>1</sup>	30 <sup>m</sup>	11	Tr	I
		4	22	29	1	0c	D	10	18	01	ΙV	Ec	R		18	19	54	II	Ec	R
		5	19	39	r	Tr	I	11	17	17	II	Ec	R		19	19	14	III	Sh	E
			20	52	I	Sh	1	12	21	36	I	Tr	I		20	20	54	I	0c	D
			21	55	1	Tr	E		22	47	I	Sh	1		21	18	03	1	Tr	I
		6	20	31	I	Ec	R	13	18	56	1	0c	D			19	11	I	Sh	I
		7	17	38	I	Sh	E	14	17	16	r	Sh	I		21	20	20	I	Tr	E
		ß	20	23	III	0c	R		18	21	I	Tr	E			21	28	I	Sh	E
			21	47	III	Ec	D		19	33	I	Sh	E		22	18	50	1	Еc	R
		9	17	48	II	Tr	I	15	20	56	111	0c	D		25	17	23	II	0 <b>c</b>	D
			20	16	II	Sh	I													

The Moons of Jupiter continued on page 22



Saturn's moons are considerably fainter than the four Galilean moons of Jupiter. The diagram shows the orbits of four of Saturn's ten moons. The easiest to find is Titan (magnitude 8,5), according to the diagram and the information in the table below (which covers the period when Saturn is clearly visible in the evening sky).

**TITAN 1979** 

Eas	Eastern				ferio	r	We	stern	1	Sup	erio	r
Elon	gat1	on		Conj	unct	ion	Elon	gati	on	Conj	unct	ion
Jan	7,	23		Jan	11,	26	Jan	14,	30	Jan	3,	19
Feb	8,	23		Feb	11,	27	Feb	15		Feb	4,	19
Mar	11,	27		Mar	15,	31	Mar	3,	19	Mar	7,	23
Apr	12,	28		Apr	16		Apr	4,	20	Apr	8,	24
May	14,	30		May	2,	18	May	6,	22	May	10,	26
Jun	15			Jun	3,	19	Jun	7,	23	Jun	11,	27
Jul	1,	17		Jul	5,	21	Jul	9,	25	Jul	13,	29
Aug	2			-	-	-	-	-	-	-	-	-
	-	-		-	-	-	-	4	+	14	-	-
Oct	21		1	Oct	25		Oct	29		Oct	17	
Nov	6,	22		Nov	10,	26	Nov	14,	30	Nov	2,	18

Dec 16

Dec

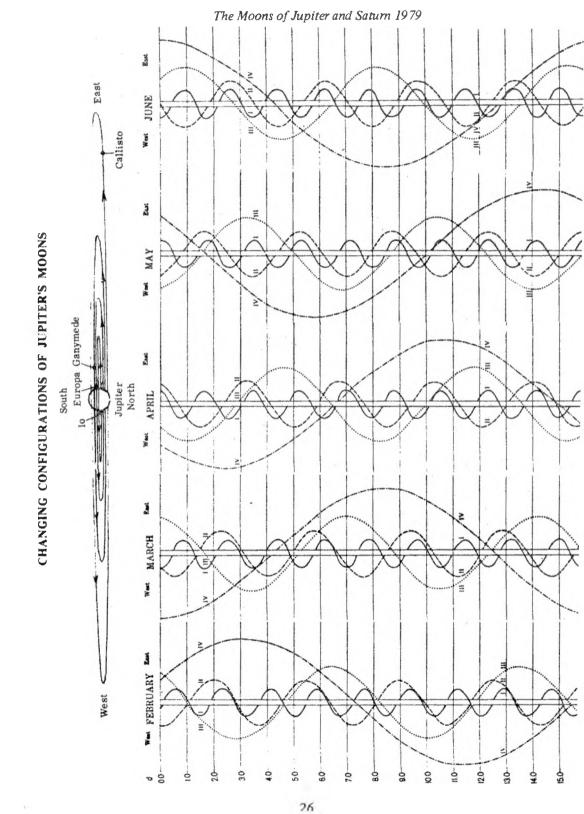
4, 20

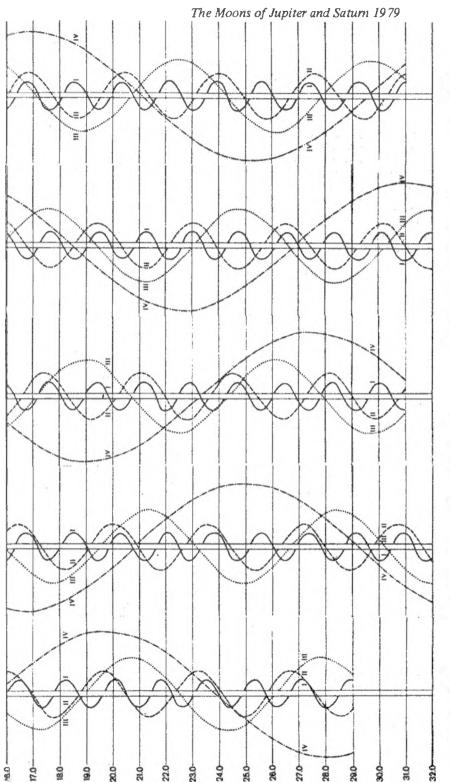
12, 28

Dec

Dec

8, 24





The four bright moons of Jupiter always appear close to a straight line passing through the planet since, as shown in the drawing at the top, their orbits are seen nearby edge on. The main part of the diagram then shows how their positions along such a straight line chang during the live months when Jupiter is prominent in the evening sky. For each month, time increases downward; the disk of Jupiter is stretched to make the central column and horizontal lines, representing 2 a.m. (0 brs. Universal time), are shown for every day of the month. The wavy lines show how the Moons appear to oscillate from one side of the planet to the other.

# **COMETS AND METEORS**

## **COMETS**

Comets are celestial bodies moving around the sun, mostly in very elongated orbits. The typical comet consists of a nucleus surrounded by a hazy aura of gas and dust called the coma, and in many cases there is a tial stretching away from the sun. Faint comets, several of which are discovered each year, usually appear only as fuzzy patches without nucleus or tail, though there may be a central condensation.

While they are believed to be true members of the solar system, comets differ radically from the planets in that their orbits, besides being highly eccentric, are inclined at all angles to the plane of the ecliptic, and their motion may be direct (like that of the planets) or retrograde. Compared with that of a planet, a comet's mass is almost negligible; nearly all this mass is concentrated in the nucleus, which is believed to be not one solid piece but composed of many separate particles of various sizes.

Comets are the most mysterious and capricious of solar system objects and the nature of the physical changes which they exhibit is still not fully understood.

Observers with quite modest equipment, say a refracting telescope of not less than 7.5 cm, can do useful work by following known comets and reporting on their appearance. The ability to make accurate brightness estimates is especially useful and well worth cultivating. Sweeping the sky for new comets, though requiring considerable patience and perseverance, is also well within the scope of the equipment mentioned.

Many of the fainter comets are undoubtedly escaping detection, particularly in the Southern skies which are not being as intensively searched as the Northern. There is a need for more amateurs to undertake this work.

Interested persons are asked to contact the Director of the Comet and Meteor Section at 90 Malan Street, Riviera, Pretoria 0084.

## **METEORS**

Meteors or "shooting stars" result from small bodies entering the Earth's upper atmosphere, and are generally seen in greater abundance after midnight (due to the direction of the Earth's motion) than in the early evening.

There are two categories of meteors - the sporadic ones and the showers. A meteor shower comes from a certain direction in space (the Radiant) and is thought to be associated with the remains of a comet. When the Earth passes close to the comet's original orbit, such a shower can be expected. A list of these predicted showers is given in the table opposite.

The term "shower" is perhaps misleading (as the table indicates); the most prolific of these showers normally yields an average of less than one meteor per minute. On rare occasions however, as in the case of the Leonids, there is a phenomenal rise in the number of meteors observed.

There is always the possibility of new showers occurring, and any large-scale meteor activity observed on dates other than those mentioned should be reported without delay.

Reports by a reliable observer of the number of meteors seen coming from a particular radiant in a given period are always useful, but the best work is done by organised teams making a full sky coverage.

"Fireballs" are meteors of a luminosity equalling or exceeding that of the brightest planets. Accurate reports of their path among the stars, or their altitude and azimuth, at specific times, are of great value, particularly if made by observers at different places along the trajectory. Details of brightness (compared with Venus, Moon etc.) size and form, colours and any train or streak, are also important.

PREDICTED METEOR SHOWERS 1979

18   Corona Australids   16 <sup>n</sup> 20 <sup>m</sup>			Shower	Radi	a a	Date	ar town \$ A C M	HOUTLING	JisasT JasibsA Xorqqs)	Recommended Time of Watch	Conditions at Maximum
24 April Lyrids 18 08 +32 Apr 22 15 04 15 07h - 02h 12 Eta Aquarids 22 24 00 May 5 18 07 30 02h - 02h 12 June Lyrids 18 00 -30 Jun 14 7 00 30 20h - 22h 25 0 June Lyrids 18 32 435 Jun 16 8 01 00 20h - 23h 29 Cetids (?New) +02 00 -15 Jun 26 8 21 30 20h - 23h 20 Cetids (?New) +02 00 -15 Jun 28 7 07 40 03h - 0ah 25 Capricornids 21 00 -15 Jul 25 8 00 50 23h - 02h 24		18	Corona Australids	16,20			16	'n	Od h45m	02h30F - dawn	_
12 Eta Aquarids	Apr	24	April Lyrids	18 08	+32		22	15	04 15	onh - ozh	Favourable
36 Sco-Sgr System 18 00 -30 Jun 14 ? 00 30 $20^h - 22^h$ 27 Junne Lyrids 18 32 +35 Jun 16 8 01 00 $20^h - 23^h$ 28 Cetids (New) -02 00 -15 Jun 28 8 02 00 -24 00 03 -04 00 03 -04 00 03 -04 00 03 -04 00 03 -04 00 03 -04 00 03 -04 00 03 -04 00 03 -04 00 00 -15 Jun 28 8 00 00 03 00 03 -04 00 00 03 -04 00 03 -04 00 03 -04 00 03 -04 00 03 -04 00 03 -04 00 00 03 -04 00 00 03 -04	May	12	Eta Aquarids	22 24	8	May	ν,	18	07 30		Favourable
21 June Lyrids 18 32 +35 Jun 16 8 01 00 20h - 23h 26 Ophiuchids 17 20 -20 Jun 20 8 23 30 20h - 24h 5 Cetids (?New) -02 00 -15 Jun 28 7 07 40 03h - dawn 15 Delta Aquarids 22 36 () Jul 27 35 02 10 22h - 24h 25 Alpha Capricornids 22 40 -30 Jul 30 11 02 10 00h - 04h 24 Iota Aquarids (22 04 -30 Jul 30 11 02 10 00h - 04h 24 Iota Aquarids (22 04 -15 Aug 6 12 (01 10 22h - 24h 27 Orionids 06 24 +15 0ct 21 35 04 30 02h - dawn 5 Taurids (03 44 +14 Nov 8 16 () 50 50 21h - 23h 20 Iota Aquarids (03 44 +21 Nov 17 10 06 30 02h - dawn 5 Phoenicids (03 44 +21 Nov 17 10 06 30 02h - dawn 15 Phoenicids (03 44 +21 Nov 17 10 06 30 02h - dawn 15 Phoenicids (03 44 +21 Nov 17 10 06 30 02h - dawn 15 Phoenicids (03 44 +21 Nov 17 10 06 30 02h - dawn 17 Oct 21 35 04 00 02h - dawn 17 Oct 21 35 04 00 02h - dawn 17 Oct 21 35 04 00 02h - dawn 17 Oct 21 35 04 00 02h - dawn 17 Oct 21 35 04 00 02h - dawn 17 Oct 21 35 04 00 02h - dawn 17 Oct 21 35 05 05 05 05 05 05 05 05 05 05 05 05 05	301	30	Sco-Sgr System		-30		14	<i>ر</i> .,	80 30	1	Unfavourable
26 Ophluchids 17 20 -20 Jun 20 8 23 30 $20^h - 24^h$ 29 Cetids (?New) +02 00 -15 Jun 28 7 07 40 03h - dawn  5 Caprifornids 22 36 (-0.7) Jul 25 8 00 50 $23^h - 02^h$ 15 Delta Aquarids 22 36 (-1.7) Jul 27 35 02 10 $22^h - 24^h$ 20 Fisces Australids 22 40 -30 Jul 30 11 02 10 $00^h - 04^h$ 24 Iota Aquarids 20 36 -10 Aug 2 10 00 00 01h - 03h  27 Orionids 06 24 +15 0ct 21 35 04 30 $02^h - 04^h$ 28 Taurids 03 44 +14 Nov 8 16 (00 50 $02^h - 04^h$ 29 Leonids 01 00 +22 Ncv 17 10 06 30 $02^h - 04^h$ 20 Leonids 01 00 -55 Dec 4 7 20 10 $21^h - 24^h$ 15 Geminids 07 28 +32 Dec 13 55 02 00 $21^h - 04^h$ 7 Velads 09 56 -51 Dec 29 7 03 30 $03^h - 04^h$	Jun	23	June Lyrids	18 32	+35		91	8		ı	Favourable
29	Jun	26	Ophiuchids	17 20	-20		20	60	23 30	1	Favourable
5 Capricornids 21 00 -15 Jul 25 8 00 50 23h - 02h 15 Delta Aquarids 22 36 (-17	Jun	58	Cetids (?New)		-15		28	~	07 40	1	Favourable
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20 Pisces Australids 22 40 -30 Jul 30 11 02 10 $\cos^{h} - \cot^{h}$ 25 Alpha Capricornids 20 36 -10 Aug 2 10 $\cos \cos \cos$	Aug	15	Delta Aquarids		(-17		27	35		4	Favourable
25 Alpha Capricornids 20 36 -10 Aug 2 10 00 00 01h - 03h 24 Iota Aquarids (22 04 -6 Aug 6 12 (01 10 22h - 24h 27 Orionids (22 32 -15 Aug 6 12 (01 40 22h - 24h 27 Orionids (03 44 +14 Nov 8 16 (00 50 20 Leonids (03 44 +14 Nov 8 16 (00 50 20 Leonids (03 44 +12 Nov 17 10 00 50 20 Leonids (03 49 +22 Nov 17 10 00 50 20 Leonids (01 00 -55 Dec 4 7 20 10 21h - 24h 215 Geminids (07 28 +32 Dec 14 55 02 00 21h - 01h 27 Velaids (09 56 -51 Dec 29 7 03 30 (03h - dawn 11)	Aug	20	Pisces Australids	22 40	-30		စ္က	11		1	Favourable
24 Iota Aquarids (22 04 -6 Aug 6 12 (01 10 22h - 24h 27 Orionids (22 32 -15 Aug 6 12 (01 40 22h - 24h 27 Orionids (03 44 +14 Nov 8 16 (00 50 21h - 23h 10 00 10 00 +12 Nov 17 10 06 30 02h - dawn 18 Phoenicids (10 08 +22 Nov 17 10 06 30 02h - dawn 19 00 07 28 +32 Dec 14 7 20 10 21h - 24h 15 Geminids (10 08 55 Dec 14 55 02 00 21h - 01h 17 Velaids (10 56 -51 Dec 29 7 03 30 03h - dawn 19 00 56 00 5	Aug	25	Alpha Capricornids	20 36	-10	Aug	7	10	8	1	Unfavourable
27 Orionids	Aug	24	Iota Aquarids		- 6	Aug	9	12	(01 10 (01 40	22h - 24h	Unfavourable
5 Taurids (03 44 +14 Nov 8 16 (20 50 21h - 23h 20 Leonids 10 08 +22 Nov 17 10 06 30 02h - dawn 15 Geminids 07 28 +32 Dec 13 55 02 00 21h - 24n 15 Geminids 07 28 +32 Dec 13 55 02 00 21h - 01h 17 Velaids 09 56 -51 Dec 29 7 03 30 03h - dawn 17 Velaids	8	27	Orionids		+15		7.7	35	94 30	02h - dawn	Favourable
20 Leonids 10 08 +22 Nov 17 10 06 30 $02^{\rm h}$ - dawn 5 Phoenicids 01 00 -55 Dec 4 7 20 10 $21^{\rm h}$ - $24^{\rm h}$ 15 Geninids 07 28 +32 Dec 14 55 02 00 $21^{\rm h}$ - $01^{\rm h}$ 7 Velaids 09 56 -51 Dec 29 7 03 30 $03^{\rm h}$ - dawn 10	Dec	M	Taurids	(03 44	+14	Nov	00	16	8 %	1	Unfavourable
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	Nov	20	Leonids		+22		17	10		1	Favourable
15 Geninids 07 28 +32 Dec 13 55 02 00 $21^{\rm h}$ - $01^{\rm h}$ 7 Velaids 09 56 -51 Dec 29 ? 03 30 $03^{\rm h}$ - dawn	Dec	2	Phoenicids		-55	Dec	4	۴.		1	Unfavourable
7 Velaids 09 56 -51 Dec 29 ? 03 30 03h - dawn	- Dec	15	Geninids		+32	Dec 1	E 4	55		,	Unfavourable
	- Jan	1	Velaids	99 60	-51		63	64		03h - dawn	Favourable

\* Attention is drawn to the suspected new shower in Cetus.

# THE STARS

### CONSTELLATIONS

Apart from our Sun all the stars that we see are so incredibly distant that, despite their high speed velocities, their apparent positions change by only minute amounts each year. Consequently the patterns that they form appear unchanged. The Greeks and other ancient civilisations identified these patterns, or constellations, with various mythological characters and creatures, and most of the names they gave are still used today.

In all there are 88 constellations, roughly one half of which would be above the horizon at any one time. Some contain distinctive patterns of bright stars and are relatively easy to find; others are difficult to locate, even with suitable maps. The Southern Cross and Centaurus, Orion and Taurus, Scorpius and Sagittarius, are featured later in this section. Detailed information on other consellations is beyond the scope of this handbook and interested observers are advised to obtain a suitable star atlas.

### STAR NAMES

Within each constellation, the brightest star is generally labelled  $\alpha$  (Alpha), the next  $\beta$  (Bets) and so on through the Greek alphabet. Most of the brightest stars also have their own names - usually of arabic origin. For example  $\alpha$  Canis Majoris, otherwise known as Sirius, is the brightest star in the constellation Canis Major.

## STELLAR MAGNITUDES

The apparent brightness of a star - which depends both on its true luminosity and its distance - is indicated by its magnitude. Equal intervals of magnitude represent equal ratios in light intensity. A star of magnitude 1,0 (typical of the brightest stars in the night sky) would be exactly one hundred times more luminous than a star of magnitude 6,0 (about the limit of visibility to the naked eye). The maps in this section show stars down to magnitude 4.5.

## STELLAR DISTANCES

Distances are often expressed in units of light years - the distance light would travel in a year (equal to  $9.5 \times 10^{12}$  km).

## DOUBLE STARS

It now appears that single stars such as our Sun are the exception, the majority of stars being double or multiple - two or more suns in orbit around one another.

## STAR CLUSTERS

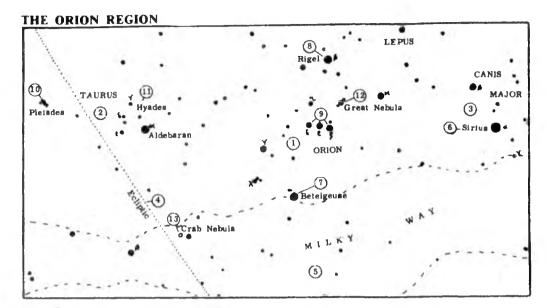
These are of two completely different sorts. Galactic clusters, having of the order of 100 stars, are found close to the plane of the Milky Way. The ones we can see are relatively nearby. Globular clusters are much larger and far more distant. They contain of the order of 100 000 stars each and are seen above and below the Milky Way on that side of the sky towards the centre of our galaxy. So great is their distance that small telescopes fail to resolve individual stars - instead they appear as fuzzy balls.

## NEBULAE

Possibly one third of the matter in our region of the Galaxy is in the form of gas and dust (the remainder being contained in stars, plus a tiny amount in planets). Condensations of this material are called nebulae, some of which are illuminated by nearby stars while others are dark. They are usually referred to by their numbers in Messier's catalogue (M) or the New General Catalogue (NGC).

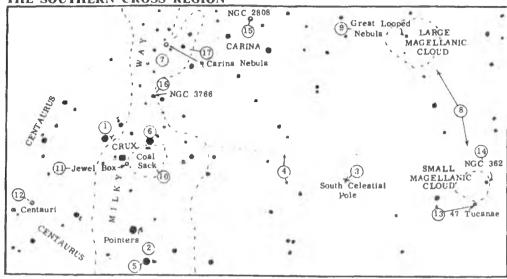
## THREE POPULAR REGIONS

The dominating constellation of the summer skies is Orion, that of the winter skies is Scorpius, while the Southern Cross is conspicuous for most of the year. The regions around these constellations are also rich in interesting objects - visible either to the naked eye, or with the aid of binoculars or a small telescope - and are featured in the maps and text below. It may be necessary to rotate the maps to match the orientation of the constellations in the sky.



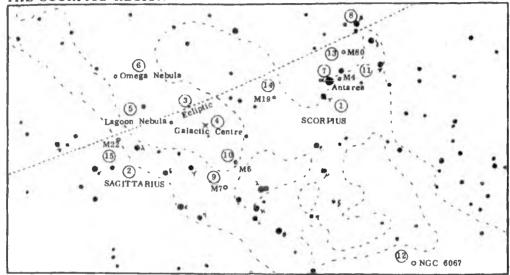
- 1) The constellation of Orion. The figure of the legendary hunter of Greek mythology is unfortunately upside down when seen from Southern Africa. The faint stars by  $\lambda$  represent the head,  $\alpha$  and Y the shoulders,  $\delta \epsilon n$  the belt, and  $\theta$  and  $\theta$  the legs. Orion forms part of the "great hunting scene" in which he faces the one-laught of  $\theta$  Taurus, the bull. Only the forepart of the bull is depicted and, like Orion, it is upside down.  $\theta$  and  $\theta$  are the eyes, Y the nose. Orion is accompanied by  $\theta$  Canis major, the large dog, and the small dog (off map) while Lepus, the hare, crouches at his feet.
- A section of the Ecliptic a line encircling the entire sky and representing the plane of the Earth's orbit. As the Earth revolves around the Sun, the Sun appears to move along the ecliptic through the constellations of the Zodiac, of which Taurus is one.
- A portion of the Milky Way (looking out towards the edge of our Galaxy).
- 6 Sirius the brightest star in the night sky. It is somewhat brighter than our Sun and relatively close by at a distance of 9 light years. It is a double star but the companion is a white dwarf (only slightly larger than the Earth, and with a mass comparable to our Sun ) and is only visible through a large telescope.
- (7) Betelgeuse most famous of the red giant stars. Its diameter is of the order of the size of the Earth's orbit and its luminosity is nearly 10 000 times that of our Sun. Its red colour should be obvious to the eye. It is 520 light years distant.
- (8) Rigel, despite being physically smaller than Betelgeuse, is more luminous (higher surface temperature bluish colour) and more distant.
- (9) The stars in Orion's belt are distant hot blue stars.
- (10) The Pleiades or Seven Sisters form the best known nearby star cluster. Six or seven stars are visible to the naked eye, binoculars or a small telescope show more.
- (1) The Hyades is another nearby galactic cluster, but Aldebaran is not a member (it lies closer to us).
- The Great Nebula in Orion, just visible to the naked eye, shows up as a fan shaped mass of luminous gas through binoculars or a telescope. A telescope will also show a tiny "Trapezium" of four stars in the centre.
- (13) The Crab Nebula, the remnant of a supernova recorded by the Chinese in 1054, requires a moderate sized telescope for observation. In its heart is located the extraodinary pulsar which emits a double flash of light 30 times every second. The current belief is that it is a rapidly rotating neutron star a star with the mass of our sun but with a diameter of only 10 km.





- (I) Crux, the Southern Cross, is one of the most compact patterns of bright stars to be found in the sky. It lies on the border of that region of the sky which never sets as seen from Southern Africa.
- (2) The two "Pointer" stars lie close to the Cross. (A similar pattern to the Southern Cross called the False Cross lies just outside and above the map, but has no accompanying pointer stars).
- (3) The South Celestial Pole: This is one of two opposite points in space towards which the Earth's axis of rotation is directed. As the Earth rotates so the sky appears to pivot about this point. It always lies above the south point on the horizon, elevated by an angle equal to the observer's southern latitude. (The north celestial pole lies below the northern horizon and can never be seen from the Earth's southern hemisphere)
- 4 The intersection of a line extended through the major axis of the Cross and the perpendicular bisector to the Pointers indicates the approximate position of the South Celestial Pole.
- (5) a Centauri has the distinction of being the closest star to our solar system at a distance of approximately 40 million million km or 4,3 light years. A small telescope readily shows that it is a double star the two components take 80 years to revolve about one another. A much fainter third star also belongs to the system.
- (6) a Crucis can also be resolved as a double star by a small telescope (separation 5 sec of arc).
- 7) The region indicated is one of the brightest sections of the entire Milky Way.
- (8) The Large and Small Magellanic Clouds are the nearest of the external galaxies (see also next section). They can be seen with the naked eye provided the sky is reasonably dark.
- (9) The Great Looped Nebula possibly the remnant of a supernova explosion in the Large Magellanic Cloud. (Naked eye or binoculars).
- (1) The "Coal Sack" a dark mass of gas and dust obscuring a part of the Milky Way. (Naked eye or binoculars).
- Herschel's "Jewel Box" a galactic cluster containing stars of different colours. (Small telescope or binoculars).
- 12 ω Centauri and 13 47 Tucanae are perhaps the best known globular clusters. Binoculars will show their fuzzy appearance. (14) NGC 362 and (15) NGC 2808 are fainter globular clusters.
- (16) NGC 3760 a fine galactic cluster. (Binoculars or small telescope).
- (17) The n Carinae nebula site of a slow supernova that brightened to magnitude -0,8 in 1843 and is now of magnitude 6,4.

## THE SCORPIUS REGION



- The constellation of Scorpius. The creature is depicted with  $\alpha$  in the centre of the body and  $\beta$  and  $\eta$  the size. The distinctive tail  $\epsilon \zeta = \theta$  curls round to the sting  $\lambda$ .
- (2) Sagittarius the figure of the centaur archer is very difficult to make out.
- (3) A section of the Ecliptic. Like Taurus, Scorpius and Sagittarius are constellations of the Zodiac.
- The direction of the centre of our Galaxy the Milky Way is that part of our Galaxy visible to us. Unfortunately the central nucleus is obscured by foreground gaseous and dusty matter both dark and luminous hence the irregular shape of the Milky Way in this region. Luminous nebulae include (5) the Lagoon nebula and (6) the Omega nebula. These are best seen with the aid of binoculars.
- Antares a distant red giant, several hundred times the diameter of our Sun is so named because its red colour rivals that of the planet Mars.
- β Scorpli can be resolved as a double star (separation 16 sec of arc) with a small telescope. In fact
  the brighter component is in itself a triple star, and the fainter component a double star:

This region includes a number of galactic clusters including (9 M7, (19 M8, (1) M4 and (12 NGC 6067. (Use binoculars or a small telescope).

Further from the plane of the Milky Way are some globular clusters: (13) M80 (14) M19 and (15) M22.

## NOVA SEARCHING

On rare occasions a star may undergo a nova outburst, its light increasing tremendously. The result is that a "new" star appears where previously no star was visible to the naked eye, or even with a small telescope. The light of the nova may fluctuate for a time, then gradually fades over a period of days, weeks or months.

Even observers having no telescopes can perform a useful task in keeping a watch for such novae in an allocated area of the sky. A good knowledge of the constellations is a recommendation, since part of the procedure is to scan the whole sky for bright novae before the more detailed search in the allocated area is begun. However, anyone can be given training in star recognition.

Interested persons should contact the Director of the Nova Search Section, Mr. J.C. Bennett, 90 Malan Street, Riviera, Pretoria 0084.

#### VARIABLE STAR OBSERVING

The General Catalogue of Variable Stars lists some 20 000 stars. Professional observatories cannot possibly monitor all of these and the observation of variable stars is a field the therefore in which amateurs can make real contributions to astronomical knowledge.

Of the 20 000 stars, at least 2 000 are suitable for monitoring by Southern Hemisphere observers but less than 200 are in fact observed from South Africa and a still smaller number receive adequate attention from southern observers.

The Variable Star Section of the ASSA exists for the purpose of encouraging observers and of acting as a medium of communication. The Section disseminates incoming information amongst observers and will forward (on request) the observations of individuals to various variable star bodies. These include the American Association of Variable Star Observers and the Variable Star Section of the Royal Astronomical Society of New Zealand. These bodies combine the South African observers' light estimates with those from other parts of the world. The resulting light curves and tables are sent to a large number of professional observatories where astronomers are interested in investigating certain of the stars more fully.

In addition to the international work, the VSS of the ASSA supplies information direct to certain South African astronomers. It is in a position to warn observatories of sudden changes in certain "VIP" stars before the overseas bodies can do so. Some of these VIP stars are:

R. Compage Porestis, variables (decrease in brightness).

A Coronae	Doleans varian	ies (decreas	e in origin	ressi:	
123753	UW Cen	181146	RS Tel	191033	RY Sgr
145971	S Aps	181824	GU Sgr		
Increase in	n brightness:				
U Geminoi	rum variables:	040971	VW Hyi	132554	BV Cen
Novae:		063462a	RR Pic	174406	RS Oph
Flare star	:	013418	UV Ceti		

Most of the above stars could be observed every night and would still not be over-observed.

Many long period variable stars vary through a large range of brightness. This means that even approximate estimates by beginners can generate quite acceptable light curves. Some of these stars are relatively easy to locate and identify and this makest them very suitable for new observers. Examples of this class of stars are:

025050	R Hor	043262	R Ret	134236	RT Cen	174162	W Pav
034625	U Eri	054331	S Col	161122 a	R Sco	180222a	VX Sgr
035124	T Eri	091868	RW Car	161122 b	S Sco	180363	R Pav
040725	W Eri	094953	Z Vel	165030 a	RR Sco	191019	R Sgr
_		-		_		214247	R Gru

Several variable stars are situated near the South Celestial Pole and can be observed throughout the year during the evening hours. Three such stars are:

055686 R Oct 131283 U Oct 172486 S Oct

055686 R Oct 131283 U Oct 172486 S Oct Others will need early morning observing sessions for fuller coverage.

A few variables are so bright that they can be monitored throughout their cycles without optical aid,  $071044 L^2$  Puppis is such a star. Norton's Star Atlas names several more.

Beginners are encouraged to practice variable star observing by following the easy-toobserve stars listed below through the brighter parts of their variations, using naked eye or binoculars.

Approximate dates of maxima 1979

Appr	oximate magnitud	e range		with a	
021403	o Ceti (Mira)	2.0 - 10.1	September September	28	
092962	R Carinae	3.9 - 10.0	Jan. 28.		25.
100661	S Carinae	4.5 - 9.9	November	20	

From the above it can be seen that there are variable stars to suit amateurs with different amounts of experience, ability, optical aid and spare time. Amateurs who want to start observing stars in any of the above catagories should contact the Director of the Variable Star Section, Mr. J. Hers, P.O. Box 48, Sedgefield 6573, Telephone 113. New observers will be given charts of a few easy objects and (if possible) a certain amount of instruction at the eyepiece. When writing, prospective observers should give brief details of their equipment.

## **ORDINARY OCCULTATIONS**

An occultation occurs when the moon passes in front of a star. The disappearance or reappearance of the star is an instantaneous event, or almost so, and is therefore easily timed. Time signals are broadcast continuously through the 24 hours from station ZUO (see page 43) and, radio propagation permitting, occultation timings can be made, either by eye and ear or by stop watch.

The essential record of a occultation observation consists of

- (i) Date and observed time of the event (to a tenth of a second if possible). Universal Time is preferable.
- (ii) The Z.C. number of the star.
- (iii) Whether disappearance or reappearance.
- (iv) The observer's estimate of the accuracy of the observation.
- (v) Whether the time given in (i) has been corrected for personal error, and if so by how much.

Also required are the size and type of telescope used, the method of timing used, and the position of the telescope.

The telescope position must be to an accuracy of 1" or better in latitude and longitude and 30 metres or better in altitude. An observer can usually read his position off a 1/50 000 map published by the Director General of Trig. Survey, Mowbray, or his equivalent in other countries. Positions determined astronomically are NOT acceptable, no matter how carefully determined.

Timings of occultations are very valuable in determining the moon's shape and motion and constitute a field in which the amateur astronomer, often with only a small, unsophisticated telescope, can make an extremely useful contribution. Interested persons are urged to contact the Director of the Society's Occultation Section, Mr. A.G.F. Morrisby (c/o Dept. of Surveyor General, P.O. Box 1580, Bulawayo, Rhodesia.

Predictions of occultations of stars brighter than magnitude 7,5, supplied by H. M. Nautical Almanac Office, are reproduced on the next two pages.

Explanation of Table:

Z.C. - the number of the star in the "Catalogue of 3539 Zodiacal Stars for the Equinox 1950.0" by James Robertson (U.S. Naval Observatory, 1939). A short index of the brighter stars is given on page 47.

Mag. - the visual magnitude of the star

P - Phase. 1 = Disappearance 2 = Reappearance

El. of

Moon - the elongation of the Moon. 0° = New Moon, 90° = First Quarter, 180° = Full Moon, 270° = Last Quarter.

U.T. - The predicted time in Universal Time which is exactly two hours behind South African Standard Time. For example 0h55,9 UT is 2h55,9 a.m. SAST.

a,b — the approximate time of an occultation at a place  $\Delta\lambda$  degrees west and  $\Delta\phi$  degrees north of the city concerned is

Predicted time +  $a.\Delta\lambda$  +  $b.\Delta\phi$ 

where a and b are in minutes of time.

P - the position angle on the Moon's limb, measured eastward from the north point.

N - no occultation A - Moon at very low altitude

S - sunlight interferes G - grazing occultation

## Occultation Predictions

	_						-	APE T				JOE	AFRES		_		5.	ALISB		
Dect.	•	z.c.	Hag.			z.	18.	500,	8. <u>3</u> 3	.900	E.	28,	000,	8. 26	.200	E.	31,	100, 1		700
		Io.			Моов	U	.7.		ъ	P	ι	ı.T.	4	ъ	P	Ū	.1.	•	Þ	P
Jan.	16	50 832 836 1439 1441	6.0 4.7 5.5 5.9 6.4	1 1 1 2 2	82 193 153 209 209	1		H H -2.7	•1.2 0.0					-1.2 +0.1		21 1		A -1.0		
Feb .	17	1442 1 15504 3530 404 650	5.8 7.0 5.2 5.7	\$2 2 1 1	210 220 49 88 111		56,1 a1.a	8	42.5	349 60	l		1	+1.6 +2.1		18		-1.8 H -0.1		
		7874 943 9474 1921 1924	7.5 6.2 5.2 5.9 5.8	1 1 2 2	123 135 135 23k 23k	23 22	24.7 05.3	# -2.0	+3.0 -1.6 -0.9	18 286	23	03.2 58.5	-0.2 И -0.8	-0.1 -1.6 -2.2 -1.7	146 312	22	59.1 37.4	-3.1 -0.9 H -0.3 -1.2	-0.1 -3.5	111 343
Mag.	20 4 6	2279 2433 609 886 1038	6.2 6.5 7.5 7.0 6.8	2 1 1 1	270 283 80 103 115	18	23.7	-1.9 ■	-1.0 +1.4 +0.7	71	24 18	15.2 55.7	-0.3 -1.7	-0.6 -1.6 +2.6 +2.8	298 46	23		-0.1 +0.3 N		
	9 12 18	1158 1258 14864 2128 30154	5.2 6.7 4.6 5.8 5.3	1 1 2 2	127 136 161 229 306	17	40.4		+0.5	159	17	37.8 10.4 33.5	-2.5 -0.6 -3.2	+1.9 -2.4 -0.6 +0.6	132 125 261	1	13.6	-2.9 -0.9 -2.7 -0.9	+0.6 -1,2	91 295
Apr.	5	832 836 9854 1235 1238	5.5 6.9 7.4 6.1	1 1 1 1	72 72 84 107 107	18	29.5	a *	-1.2 +0.4	156	18	31.1	-1.0	+0.1 -0.1 -0.5	116	18	37.7	-1.9 -1.3 -1.7 -0.2	+0.8	94
	7		5.9 6.4 5.8 6.4		129 129 129 140 152	21	08.6	-0.8	-1.2 -2.3 +1.9	158	21	11.8	-1.7	+0.3 -0.4 -1.6	118	21 21	28.6	-2.4 -1.7 H -0.3	-0.h	1 12
	16 16 17	23724 23724 24954 2647 2826	6.0 6.4 4.0	2 2 1	224 224 235 248 262				-2.7 +2.6		3	49,4	-1.7	-0.5 +1.0 0.0	265	22 21	19.9 54.4	-2.5 8 -1.4 -1.7 -0.8	-1.1 +2.1	283 220
May	20 21 22	2826 3109 3112 3270 1405	4.0 6.5 6.2 6.1 7.0	2 2 2 1	262 288 288 302 98				~0.5 +0.6					-1.3 +2.9		23	43.9	-2.3 -0.4 -0.8 G	+0.2	245
June	10 14	1921 1926 1941 2460 1875d	5.9 5.8 4.8 6.1 6.5	1 1 2 1	153 154 156 206 122	18	43.2 53.3	-1.h	-1.5 -1.9 +0.4	124	18 0 1	51.4 46.3 53.8	-2.1 -0.9 -2.2	+0.1 -0.6 -0.8 -0.7 -3.2	90 132 297	1	40.8	N H -0.9		341
	6 8 13	1994d 2008 2128 3015d 3188	6.5 6.6 5.8 5.3 5.4	1 1 2 2	13h 135 1h8 226 2h1	20	31.9	+0.2	-1.3 -0.7 -1.9 -2.4	134 299				+0.1 +0.3		23		-1.0 -0.4		
July	19 27 29	33224 219 1323 1531 1732	6.4 5.1 6.3 5.9 7.0	2 1 1	254 296 35 57 79	1		8	+1,1 +1,5		- %	09.3	-1.6	-1.8 +1.1 -0.3	229	Į.	•	3 -1,4 3 -1,5		

## Occultation Predictions

					El.		C	APE T		0		KOL	ANDES				LISBURY		
Date	0	Z.C.	Mag.			E.	18,	500 . :	8. 33		E		000.			R. 31.1	00. 8.		
		po.			Modifi	υ	т.		ь	P		J.T.	6.	ъ	P	U.T.	B.	ъ	P
July	7 8	2072 2184 2460 2495d 3109	6.7 7.0 6.1 6.0 6.5	1 1 1 2	115 126 152 154 207	18 17 0	14.5 37-5 14.2	-1.h -0.5 -1.3	-1.2 -2.9 -2.9 +1.0 -3.7	145 145 89	18 17	49.2 30.7	-2.5 -1.5	+0.1 -0.9 -1.4 +1.5	110	21 42.5 18 58.2 17 31.0 0 52.1	-3.4 +1 -2.3 +0	.1	76 76 74 40
Aug.	11 13 1	3112 3131 34304 2128 2133	6.2 5.5 5.7 5.8 5.6	2 2 1 1	208 209 237 94	23	18.7 26.5	-2.2	-1.6 -1.9 +1.0	294	22	40.1	H -1.2 H	+0.3	242	17 51.5			271 171
	la la 5	2279 2433 25554 2731 3353	6.2 6.5 7.5 6.5 3.8	1 1 1 1	108 122 133 147 203	0	2.9		+0.7 +2.0 +0.9		21		-0.1 A -1.6	+1,1	56				
	10 28 30	3353 3383 2089 2365 2485	3.8 6.5 6,6 7.1 7.4	2 1 1 1	203 206 64 89 100		18.0 38.0		-0.1	173 121	17 20	48.9	H -1.7 -0.7	-1.1 +0.5 -3.7	101	20 12.0 2 06.0 17 16.1 20 57.6 16 23.1	-1.8 +0 -0.3 +1	.3	173 97 74
fleyt.	1	2495d 2640d 2647 2649d 2653d	6.0 6.1 6.4 6.6	1 1 1 1 1	100 113 113 113 113			R	-1.1 -1.9		16 17	50.5	-2.1 -2.6	+0.6 -3.9 0.0	144 92	18 43.3 16 39.4 18 05.6 18 30.3 18 50.4	-2.7 -1 -2.6 +1 -3.2 -3	.0	60
	3 4 9	2826 2986 3131 219 2184	6.4 5.5 5.1 7.0	1 1 2 1	127 142 155 215 46	23 I	00.2 39.4	-1.6	-0.4 +0.1 +2.2 +1.7	108 32	23 20 3	18.3 23.6 30.9	-17.3 -1.4	+0.8 +0.6 +1.5 -1.5	97 357 245	20 29.3 23 28.7 17 49.2	-0.9 +1 B S	٥.	56 75
Oct.	30 1 2	2454 2922 3066 30834 3353	7.2 7.4 6.0 7.3 3.8	1 1 1 1	70 109 122 124 149	21 19	14.9	-2.3 -2.3	-1.0 -1.2 +0.1	131	20	28.0	-1.5 -2.1 A	0.0 -0.2 +0.9	116	19 13.3 21 33.7 20 20.8 17 04.0	-0,9 +0 -1.6 +1 A	-7	87 91 57
	15 18 18	3360 1371 4006 4006 2279	6.3 6.4 1.4 1.4 6.2	1 2 1 2 1	149 292 325 325 28	03	10.5	-1.2 R -0.6	-2.2 -2.6 -1.1 +1.4	316 264	02	59.1 32.6	-0.3	-0.6 -1.6 -1.4		18 22.0 02 24.0 03 30.6	R -0.≩ -0	.6	63 89 308
	27 30 31	25554 2889 33224 3461 3465	7.5 7.1 6.h 6.k 6.5	1	53 79 119 131 132	21 21	51.5 55.0	0.0	•2.2	73	22 18	18.0	A-0.4	+1.9 +2.b +0.7	35 121	22 42.7 18 39.5 20 13.3	-3.1 -0	1.1	6 91 61
Bov.	13		5.1 4.8 -1.6 -1.6 6.4	1 2 1 2 1	161 202 285 285 285	7 7 7	01.4 10.6 26.7	:	-0.2	38 18		45.0 10.0	-1.6 N N		86 328		-1.4 +1 N N M	. 4	65
Dec.	21 27 9	3120 3405 3430d 1487d 6006	7.0 7.0 5.7 1.3	1 1 2 1	73 100 102 251 276	22 (	յն, ե	H S -0.9 A N	+0,6	106	22 21	15.9 55.1	N 3 -0.4 -0.6	+0.8 -2.0	92 306	17 50.7 17 29.2 21 35.7 05 44.6	-1.2 +3 A -0.9 -3	6	24 336
	23 24 24	4006 3238 3379 3383 3528	1.3 7.0 6.4 6.5 7.5	2 1 1 1	276 58 70 71 85	19	6.5	H +0.3 N -0.6 -0.5	+1.7	18 63 76	20	03.6	H S G -0.2 A	<b>+</b> 1 . B	52	06 50.0 19 31.8	5		
	27	2694	7.3	1	111	22	2.0	-0.1	+3.0	22			3						

4005 = JUPITER 4006 = SATURE

## **GRAZING OCCULTATIONS**

when a star moves tangentially to the limb of the Moon, and is occulted for a very short period only - a few minutes, or even seconds - a grazing occultation is said to occur. Because the limb, as seen from the Earth, is in fact the outline of numerous mountains and valleys, there may be several disappearances and reappearances, which are not only fascinating to observe, but which may be accurately timed to yield valuable data on the relative positions of star and Moon, in both right ascension and declination, as well as on the shape of the Moon. Some of these data cannot readily be obtained in any other way.

The maps on the following pages have been prepared by M.M. Nautical Almanac Office to show the tracks of stars brighter than 7.5 magnitude which will graze the limb of the Moon when it is at a favourable elongation from the Sun and at least 10° above the observer's horizon (2° in the case of bright stars). Each track starts in the West at some arbitrary time given in the key and ends beyond the area of interest, except where the letters "A", "B" or "S" are given. "A" denotes that the Moon is at a low altitude, "B" that the bright limb interferes, and "S" that sunlight interferes. The tick marks along the tracks denote 5 minute intervals of time which, when added to the time at the beginning of the track, give the approximate time of the graze at places along the tracks.

The tracks as shown on the maps are approximate only. Since the observer's location is very critical, successful observations call for very accurate predictions. Such predictions, which include graphical representations of the probable profile of the Moon, are computed annually for a number of centres in southern Africa. By plotting the predicted graze track on a reliable survey map (e.g. South African 1:50 000 series) it is usually possible to select a convenient site from where the graze may be observed. Ideally a team of observers would be stationed at intervals along a line runing at right angles to the graze track - say, along a main road - each with his own telescope and timing equipment. Each observer will see a different sequence of events, the combined results forming an accurate picture of the limb of the Moon.

The equipment needed is similar to that used for ordinary (or 'total') occultations, but must, of course, be portable. A 75mm refractor is ideal for average events, but better instruments with a larger aperture have often shown their superiority under difficult conditions. Timing is best carried out with a portable tape recorder and radio receiver tuned to ZUO or other time signal station.

It will be seen from the maps that many grazing occultations occur in regions which are rather far removed from the main cities, and which cannot casily be reached by teams of observers from one of the ASSA centres. It is worth remembering, however, that a team of many observers, while ideal, is by no means essential; that a single good observer is worth more than many unsuccessful ones, and that one good observation is worth infinitely more than no observation at all.

Interested observers - especially those living in the more distant regions - who wish to be informed of favourable grazes occurring in their neighbourhood, are therefore invited to contact the co-ordinator for grazing occultations: Mr. J. Hers, P.O. Box 48, Sedgefield 6573. Telephone: 113.

#### EXPLANATION OF COLUMN HEADINGS IN TABLES

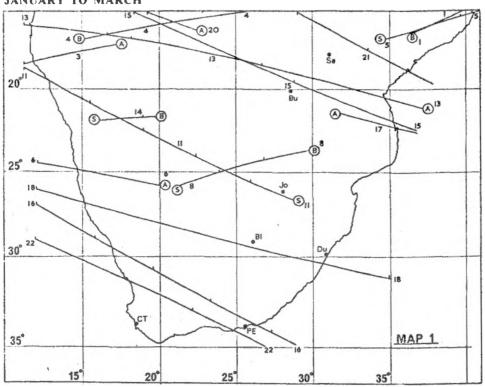
- No. the number of the track on the map. An asterisk denotes that the same is double notes are given below.
- Z.C. the number of the star in the Zodiacal Catalogue.

Date

Beginning an arbitrary time of the beginning of the track in the west.

- Sunlit the percentage of the Moon's disk lit by the Sun
- Limit N = northern limit (a complete occultation takes place south of the track)
  - S = southern limit (complete occultation north of the track)

#### JANUARY TO MARCH



	ZC	MAG.	OF THE T		IN TH		PERCETT SUNLIT	N OR S LIMIT		2C	MAG.	TIME (U.T OF THE	TRACK II		WEST	PERCENT SUMLIT	LIMIT
1	3146	6.5	JANUARY	1	16	17	12	n	1%	886	7.0	MARCH	6	17	k3	62	8
3	3385	6.6	JANUARY	30	18	36	9	11	15	1158	5.2	MARCH	8	22	50	80	И
k	3530	7.0	JANUARY	31	18	32	17	н	16	2128	5.8	MARCH	18	٥	21	83	8
5	109	6.5	PEDRUARY	1	16	43	26	8	17	3008	6.9	MARCH	24	0	$I_0I_2$	20	В
6	404	5.2	FESRUARY	3	21	35	49	И	184	3015	5.3	MARCH	24	3	34	50	R
8.	650	5-7	PEBRUARY	5	17	51	68	8	20	1238	6.1	APRIL	5	22	ko	65	M
11	2167	7.5	FEBRUARY	19	2	ji0	59	8	21	1439	5.9	APHIL	7	20	51	Вэ	21
13	609	7.5	HARCE	1,	19	10	12	N	22*	2495	6.0	APRIL	16	21	47,	78	9
	Trac	k No	>	Zc													

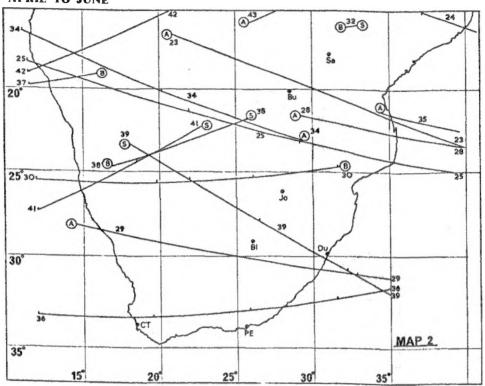
8 650 is a spectroscopic binary.

18 3015 is the mean of the double star Aitken 14099. The components are of magnitude 5.8 and 6.3; separation 0"2 in p.a. 115°.

22 2495 is the mean of the two brightest components of the

is the mean of the two brightest components of the triple star Aitken 10465. These components are of magnitude 6.3 and 7.4; separation 1"9 in p.a. 262°. The third component is of 11th magnitude at a wide separation.

#### APRIL TO JUNE

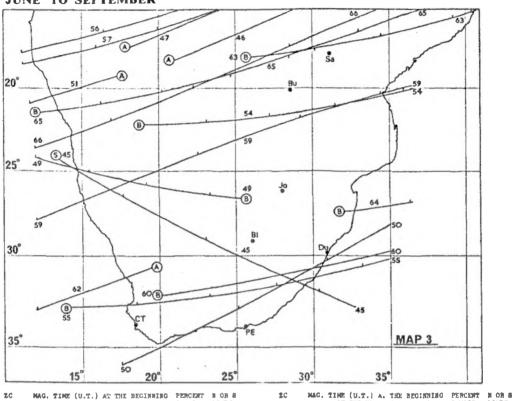


	ZC	MAG.	OF THE TRA		IN TH		PERCENT SUNLIT	N OR S		ZC	MAG.	OF THE	T.) AT	IN TH	ECIRNING E WEST MIN	PERCENT SUNLIT	N OR S	
53	2647	6.4	APRIL	17	21	33	68	8	35	3066	6.0	MAY	17	21	19	61	8	
5#a	2653	6.4	APRIL	17	22	20	68	9	364	3083	7.3	HAY	18	1	19	59	5	
25	2826	4.0	APRIL	19	0	2	56	5	37	3238	7.0	MAT	19	2	59	h7	8	
28	3112	6.2	APRIL	20	23	47	34	8	38	109	6.5	MAY	22	4	7	16	У	
29	3119	6.7	APRIL	21	0	45	33	В	39	1360	7.5	YAK	31	16	34	30		
30	3131	5.5	APRIL	21	3	0	33	8	la 1	354	5.5	JUNE	20	4	36	18	N	
32	3280	7.4	APRIL	22	3	3h	22	Ħ	42	475	7.4	JUNE	21	1	27	11	18	
34	1405	7-0	KAY	L	21	19	57	N	43	618	7.2	JUNE	22	4	1	5	8	

24 2653 is the mean of the two brightest components of the triple star Aitken 11240. These components are of magnitude 6.9 and 7.4; separation 0.4 in p.a. 131°. The third component is of magnitude 9.4 at a separation of 17.2 in p.a. 52° from the mean position.

36 3083 is the brightest component of the double star Aitken 14638. The companion is of magnitude 11.3; separation 4"9 in p.a. 170°.

#### JUNE TO SEPTEMBER

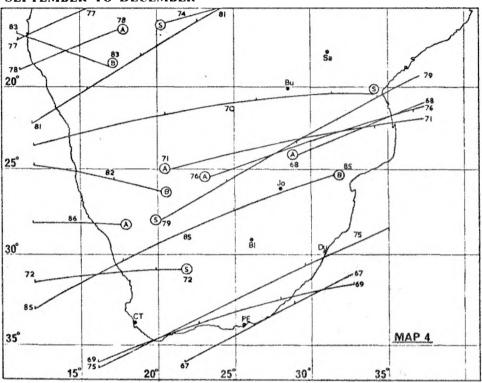


	20	MIN .	OF THE TR		RT U		SUMLIT	LIMIT		-		OF THE TRACK		E WEST	SUMLIT	LIMIT
45	1531	5.9	JUNE	29	16	55	23	Ħ	56	2658	5.4	SEPTEMBER 1	20	28	71	8
46	662	7-5	JULY	21	3	35	8	II.	57	2828	6.0	SZPTEMBER 2	20	16	81	8
hŢ	863	6.7	JULY	21	3	46	8	18	59	1060	7.4	GEPTEMBER 15	2	31	33	18
kg	2133	5.6	AUGUST	1	18	38	54	1	60	218k	T-0	SEPTEMBER 25	18	7	15	8
50	269	7.3	AUGUBT	13	۵	6	70	H	62	2460	6.1	SEPTEMBER 27	21	13	34	3
51	2372	$k_{a}k_{b}$	AUGUST	30	22	12	50	8	63	2578	6.4	SEPTEMBER 28	16	54	43	8
544	2649	6.6	S EPTENGER	1	18	21	70	8	64	2745	6.9	SEPTIONER 29	16	32	54	В
550	2653	6.4	SEPTEMBER	1	18	48	70	8	65	2755	6.6	SEPTIOGER 29	17	14	5h	8
									66	2764	6.3	SEPTEMBER 29	19	50	55	8

54 2649 is the brightest component of the triple star Aitken 10086. The brighter companion is of magnitude 9.6 at a wide separation; the second companion of magnitude 10.1 is at a separation of 10"4 in p.a. 132°.

is the mean of the two brightest components of the triple star Aitken 11240. These components are of magnitude 6.9 and 7.4; separation 0,4 in p.a. 131. The third component is of magnitude 9.4 at a separation of 17,2 in p.a. 52° from the mean position.

## SEPTEMBER TO DECEMBER



	ZC	MAG.	TIME (U.T.) OF THE TR		IN TH		PERCENT SUMLIT	N OR S LIMIT		ZC	MAG.	OF THE T		IN T		PERCENT SUNLIT	N OR B LIMIT
67	2922	7.4	SEPTEMBER	30	21	34	67	3	76	1203	7.1	ROVEMBER	9	22	36	67	N
68	1006	6.9	OCTOBER	11	22	62	60	N	77	2833	7.0	NOVEMBER	23	18	15	16	5
69	1025	7.4	OCTOBER	12	1	45	59	М	78	2846	6.9	BOVEMBER	23	19	42	17	8
70	1260	7.0	OCTOBER	14	2	15	40	N	79	3120	7.0	NOVEMBER	25	17	43	36	S
71	1371	6.4	OCTOBER	15	1	32	31	25	81	3412	h_4	NOAEMBEK	27	19	12	59	8
72	1375	5.6	OCTOBER	15	3	18	30	N	82	1501	7-3	DECEMBER	10	1	16	65	Ħ
74	2865	5.9	OCTOBER	27	17	3	39	s	83	2137	6.4	DECEMBER	16	3	29	11	N
75	940	5.7	NOVEMBER	7	21	57	85	К	85	3379	6.4	DECEMBER	24	19	18	34	8
									86	269	7.3	DECEMBER	27	23	6	69	N

I THE TICKS ARE AT 10 MINUTE INTERVALS.

# TIME SYSTEMS AND TELESCOPE SETTINGS

This section is intended to serve established amateurs and professional astronomers - i.e. those having some knowledge of time and coordinate systems. Space in the booklet does not permit full explanation, which in any case would appear complicated to the layman.

#### TIME SIGNALS FROM RADIO STATION SUO

Radio signals of mean solar time are generated by the Precise Physical Measurements Division of the National Physical Research Laboratory in Pretoria. They are broadcast by the Post Office, the 2.5 and 5 MHz signals from Olifantsfontein, and the 100 MHz signals from Johannesburg.

Carrier Frequency	Radiated Power	Time of Transmission
2,5 MHz	4 Kw	2000 - 0600 SAST
5 MHz	4 Kw	Continuous
100 MHz	80 w	Continuous

The signals consist of one pulse per second, each pulse consisting of 5 cycles of 1000 Hz tone. The first pulse in every minute is lengthened to 500 milliseconds. Morse code announcements are made during the minute preceding every fifth minute. They consist of the call sign ZUO (repeated 3 times) and the Universal Time (formally known as Greenwich Mean Time) at the next minute. (A special coding indicating UTI minus UTC is also indicated in the first 15 seconds of the minute by slightly lengthened second pulses)

#### SOUTH AFRICAN STANDARD TIME

South African Standard Time (as in everyday use) is mean solar time for the 30° East meridian (which runs east of Johannesburg and just west of Durban) and is exactly 2 hours ahead of Universal Time

#### TIME OF SUN'S TRANSIT OVER 30° MERIDIAN

The table below gives the SAST when the Sun transits the  $30^{\circ}$  meridian - and a sundial on that meridian reads noon.

Jan	1	12 <sup>h</sup> 03 <sup>m</sup> 24 <sup>s</sup>	May 11	11 <sup>h</sup> 56 <sup>m</sup> 21 <sup>s</sup>	Sep 18	11 <sup>h</sup> 54 <sup>m</sup> 19 <sup>8</sup>
	11	12 07 48	21	11 56 29	28	11 50 50
	21	12 11 13	31	11 57 32	Oct B	11 47 42
	31	12 13 26	Jun 10	11 59 15	18	11 45 17
Feb	10	12 14 16	20	12 01 21	28	11 43 52
	20	12 13 49	30	12 03 29	Nov 7	11 43 41
Mar	2	12 12 18	Jul 10	12 05 13	17	11 44 53
	12	12 09 56	20	12 06 15	27	11 47 29
	22	12 07 04	30	12 06 24	Dec 7	11 51 16
Apr	1	12 04 03	Aug 9	12 05 31	17	11 55 53
	11	12 01 12	19	12 03 42	27	12 00 51
	21	11 58 48	29	12 01 04	31	12 02 48
May	1	11 57 08	Sep 8	11 57 50		

#### SIDEREAL TIME ON THE 30° MERIDIAN

Sidereal Time is given by the line of Right Ascension coinciding with the meridian.

		At 0 h	rs A	t 21 SAS				hrs AST		21 h Sast	rs			) hrs		21 hrs
Jan	1	6 <sup>h</sup> 41	n 3	h <sub>44</sub> π	а мау	11	15	1 3 m	12	17 <sup>m</sup>	Sep	18	23 <sup>t</sup>	146 <sup>m</sup>	20	149 <sup>m</sup>
	11	7 20	4	23		21	15	53	12	56		28	2	25	21	29
	21	8 00	5	03		31	16	32	13	36	Oct	8	1	05	22	80
	31	B 39	5	43	Jun	10	17	12	14	15		18	1	44	22	48
Feb	10	9 19	6	22		20	17	51	14	55		28	2	24	23	27
	20	9 58	7	01		30	18	31	15	34	Nov	7	3	03	0	06
Mar	2	10 37	7	41	Jul	10	19	10	16	13		17	3	42	0	46
	12	11 17	8	20		20	19	50	16	53		27	4	22	1	25
	22	11 56	9	00		30	20	29	17	32	Dec	7	5	01	2	05
Apr	1	12 36	9	39	Aug	9	21	09	18	12		17	5	41	2	44
	11	13 15	10	19		19	21	48	18	51		27	6	20	3	24
	21	13 55	10	58		29	22	27	19	31		31	6	36	3	39
May	1	14 34	11	37	Sep	8	23	07	20	10						

#### CORRECTION FOR PLACES NOT ON THE 300 MERIDIAN

Approximate longitude corrections from the 30° East Meridian are provided below.

To find time of Sun's transit over local meridian, apply the longitude corrections to the data in the table above.

To find the side real times at SAST 0 hrs and SAST 21 hrs apply the corrections with the sign reversed to the data in the table.

Bloemfontein		East London	+ 8 <sub>m</sub>	Port Elizabeth	+18 <sup>m</sup>
Bulawayo	+ 6 <sup>11</sup>	Grahamstown	+14 <sup>m</sup>	Pretoria	+ 7 <sup>m</sup>
Cape Town	+46 <sup>m</sup>	Johannesburg	+ 8 <sup>m</sup>	Salisbury	- 4 <sup>m</sup>
Durban	- 4 <sup>m</sup>	Kimberley	+21 <sup>m</sup>	Windhoek	+52 <sup>m</sup>

#### TELESCOPE SETTING

When a telescope equipped with setting circles is aimed on the meridian, Its R.A. circle should read the sidereal time. Thus one can calculate the sidereal time and then set the circle, but is is usually simpler to aim the telescope at one of the well known stars given below and then to adjust the R.A. circle.

#### A LIST OF BRIGHT STARS FOR CHECKING TELESCOPE CIRCLES

Star	R. A.	Dec.	Mag.	Sp.	l Star	R. A.	Dec.	Mag.	Sp.
Achernar	1h36m8	-57°21'	0.6	<b>B</b> 5	Procyon	7 <sup>h</sup> 38 <sup>m</sup> 1	+5° 17'	0.5	F5
Aldebaran	4 34 6	+16 28	1,1	K5	Regulus	10 07 .1	+12 05	1.3	B8
Rigel	5 13 ,4	- 8 14	0,3	B8	Spica	13 24 0	-11 02	1,2	B2
Bete lge use	5 53 ,9	+ 7 24	0,4	M0	Arcturus	14 14 .6	+19 18	0.2	KO
Canopus	6 23 ,4	-52 41	-0.9	F0	Antares	16 28 ,0	-26 23	1.2	M1
Sirius	6 44 .1	-16 41	-1.6	An	Altair	19 49 7	+8 48	0 9	A5

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## **PAST PRESIDENTS**

1922-23	S.S. Hough	1941-42	H. Knox Shaw	1960-61	G.G. Cillie
1923-24	R.T.A. Innes	1942-43	A.F.I. Forbes	1961-62	M.D. Overbeek
1924-25	J.K.E. Halm	1943-44	W.H. van den Bos	1962-63	A.J. Wesselink
1925-26	W. Reid	1944-45	A.W.J. Cousins	1963-64	A.G.F. Morrisby
1926-27	H. Spencer Jones	1945-46	R.H. Stoy	1964-65	H.C. Lagerwey
1927-28	A.W. Roberts	1946-47	W.P. Hirst	1965-66	A. Menzies
1928-29	A.W. Long	1947-48	J. Jackson	1966-67	G.R. Atkins
1929-30	H.E. Wood	1948-49	A.E.H. Bleksley	1967-68	J. Hers
1930-31	D. Cameron-Swan	1949-50	W.S. Finsen	1968-69	J.C. Bennett
1931-32	H.L. Alden	1950-51	H.E. Krumm	1969-70	J. Churms
1932-33	H. Spencer Jones	1951-52	A.D. Thackeray	1970-71	W.C. Bentley
1933-34	D.G. McIntyre	1952-53	J.C. Bentley	1971-72	A.H. Jarrett
1934-35	J.K.E. Halm	1953-54	David S. Evans	1972-73	K.J. Sterling
1935-36	J. Jackson	1954-55	P. Kirchhoff	1973-74	G.A. Harding
1936-37	H.E. Houghton	1955-56	W.H. van den Bos	1974-75	C. Papadopoulos
1937-38	J.S. Paraskevopoulos	1956-57	S.C. Venter	1975-76	P.A.T. Wild
1938-39	T. Mackenzie	1957-58	M.W. Feast	1976-77	S. Booysen
1939-40	R.A. Rossiter	1958-59	H. Haffner	1977-78	B. Warner
1940-41	E.B. Ford	1959-60	P. Smits		

## **HONORARY SECRETARIES**

1922	H.W. Schonegevel	1923	H.E. Houghton	1934	H.W. Schonegevel
1922	T. Mackenzie	1930	S. Skewes	1935	A. Menzies
1923	C.L. O'Brien Dutten	1931	H. Horrocks	1965	T.W. Russo

## **HONORARY MEMBERS**

Prof. A.E. Bleksley	Dr W.S. Finsen	Dr R.H. Stoy
Dr A.W.J. Cousins	Mr H.E. Krumm	Dr A.G. Velghe
Mr R.P. de Kock	Dr J.H. Oort	Dr A.J. Wesselink
Dr David S. Evans	Mr M.D. Overbeek	Sir Richard Woolley
Prof Ch Febranbach	De 7 Cabille	1

## **GILL MEDALLISTS**

1956	H. Knox Shaw	1960	W.H. van den Bos	1967	W.S. Finsen
1957	W.P. Hirst	1963	A.W.J. Cousins	1970	J.C. Bennett
1958	J. Jackson	1965	R.H. Stoy	1976	A.D. Thackeray

## **JULIAN DATE AT 1400 HOURS**

	JAN	FEB	MAR	APR	MAY	JUN	JUL	AUG	SEP	OCT	NOV	DEC
	2443	2443	2443	2443	2443	2444	2444	2444	2444	2444	2444	2444
1	875	906	934	965	995	026	056	087	118	148	179	209
2	876	907	935	966	996	027	057	088	119	149	180	210
3	877	908	936	967	997	028	058	089	120	150	181	211
4	878	909	937	968	998	029	059	090	121	151	182	212
5	879	910	938	969	999	030	060	091	122	152	183	213
6	880	911	939	970	*000	031	061	092	123	153	184	214
7	881	912	940	971	*001	032	062	093	124	154	185	215
8	882	913	941	972	*002	033	063	094	125	155	186	216
9	883	914	942	973	*003	034	064	095	126	156	187	217
10	884	915	943	974	*004	035	065	096	127	157	188	218
11	885	916	944	975	*005	036	066	097	128	158	189	219
12	886	917	945	976	*006	037	067	098	129	159	190	220
13	887	918	946	977	*007	038	068	099	130	160	191	221
14	888	919	947	978	*008	039	069	100	131	161	192	222
15	889	920	948	979	*009	040	070	101	132	162	193	223
16	890	921	949	980	*010	041	071	102	133	163	194	224
17	891	922	950	981	*011	042	072	103	134	164	195	225
18	892	923	951	982	*012	043	073	104	135	165	196	226
19	893	924	952	983	*013	044	074	105	136	166	197	227
20	894	925	953	984	*014	045	075	106	137	167	198	228
21	895	926	954	985	*015	046	076	107	138	168	199	229
22	896	927	955	986	*016	047	077	108	139	169	200	230
23	897	928	956	987	*017	048	078	109	140	170	201	231
24	898	929	957	988	*018	049	079	110	141	171	202	232
25	899	930	958	989	*019	050	080	111	142	172	203	233
26	900	931	959	990	*020	051	081	112	143	173	204	234
27	901	932	960	991	*021	052	082	113	144	174	205	235
28	902	933	961	992	*022	053	083	114	145	175	206	236
29	903		962	993	*023	054	084	115	146	176	207	237
30	904		963	994	*024	055	085	116	147	177	208	238
31	905		964		*025		086	117		178		239