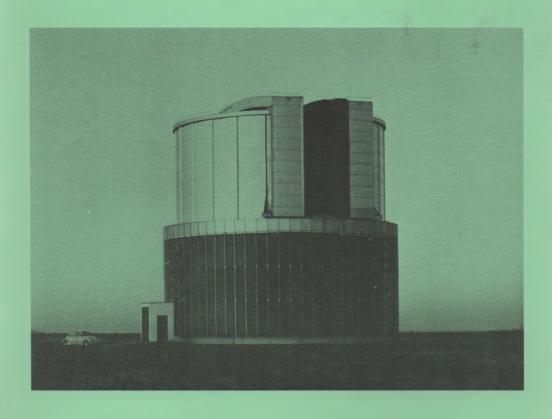
ASTRONOMICAL HANDBOOK FOR SOUTHERN AFRICA

1984



published by the Astronomical Society of Southern Africa.

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PREDICTED PERIHELION PASSAGES OF COMETS 1984

Periodic comet	Perihelion Date		Revolution Period years	Perihelion Distance au	Magnitude
Tritton	Mar	3	6.36	1.44	18.9
Encke	Mar	27	3.30	0.34	8.7
Wolf	May	31	8.21	2.42	19.1
Schuster	June	2	7.25	1.53	19.7
Gehrels 3	June	3	8.14	3.44	18.4
Tsuchinshan 2	July	21	6.85	1.79	20.0
Wild 2	Aug	20	6.18	1.49	11.9
Wolf-Harrington	Sep	23	6.53	1.62	16.6
Neujmin 1	0ct	8	18.21	1.55	12.5
Arend-Rigaux	Dec	1	6.84	1.45	16.0
Haneda-Campos	Dec	26	6.27	1.22	15.2

The expected magnitude at the time when the comet will be at its brightest after perihelion passage is as quoted above.

ASTRONOMICAL HANDBOOK FOR SOUTHERN AFRICA 1984

This booklet is intended both as an introduction to observational astronomy for the interested layman — even if his interest is only a passing one — and as a handbook for the established amateur or professional astronomer.

FRONT COVER

The 1.8m reflecting telescope at Sutherland. This is the largest telescope in South Africa and was formerly at Radcliffe Observatory, Pretoria.

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NOTE

This handbook is produced for the Astronomical Society of Southern Africa. The data it contains has been adapted for Southern Africa from data obtained from the Royal Greenwich Observatory, Herstmonceux, from the British Astonomical Association, the International Lunar Occultation Centre, Tokyo and the Hydrgrapher of the South African Navy.

All correspondence concerning this booklet should be addressed to the Handbook Editor, Astronomical Society of Southern Africa. 8 Glebe Road, Rondebosch. 7700 from whom further copies are available at R2.00 per copy.

Although every care has been taken in the compilation of the Handbook, it is sold and distributed on the explicit understanding that neither the Society or any of its members accepts any responsibility for errors therein.

Dit spyt ons dat as gevolg van beperkte fondse en produksiefasiliteite dit nie moontlik is om die handboek in Afrikaans te laat druk nie.

> R F HURLY EDITOR

ASTRONOMY IN SOUTHERN AFRICA

As one of the few parts of the Earth having both access to the rich southern skies and a suitable climate, Southern Africa holds a favoured position in astronomy. Consequently, it has seen the establishment of a number of professional observatories engaged in research while many individuals have become enthusiastic amateur astronomers. Planetaria and visiting nights at observatories convey to the general public much of what goes on in this field.

Observatories

The South African Astronomical Observatory (S.A.A.O.), was established in 1972, as a joint venture between the Council for Scientific and Industrial Research of South Africa and the Science Research Council of Great Britian, combining the facilities of the former Royal, Republic and Radcliffe Observatories, and is directed by Dr M W Feast. Its excellent observing site near Sutherland, in the Karroo, has four instruments, namely the 1.8m, 1.0m, 0.75m and 0.5m telescopes. The headquarters are in Cape Town, where also a limited amount of observing continues. Research is undertaken in many areas, with considerable effort being put into the study of variable stars, the Magellanic Clouds, the Galactic centre and globular clusters as well as optical studies of celestial X-ray sources. These studies involve the use of spectroscopic, photometric and infrared techniques. Besides providing research facilities for its own staff S.A.A.O. observing time is allocated to astronomers from Great Britian and from South African Universities.

The 0.67m visual refractor of the former Republic Observatory, Johannesburg, is maintained by the National Institute for Telecommunications Research (N.I.T.R.).

The National Radio Astronomy Observatory at Hartebeeshoek, near Krugersdorp, is, under the direction of Dr G Nicolson, operated by the NITR. The telescope, a 26m dish, is used for observations of extragalactic radio objects such as quasars and X-ray sources. The Rhodes University Radio Astronomy Group led by Prof E.E. Baart use this telescope, currently in a survey of the entire southern sky at 13cm wavelength.

Boyden Observatory, situated at Mazelspoort, 25km from Bloemfontein, is operated by the Department of Astronomy of the University of the Orange Free State, under the leadership of Prof A. H. Jarrett. The site offers good conditions, without being remote from a large centre, and observing facilities include the 1.52m Rockefeller Reflector, and the 0.41m Nishimura Reflector. Research covers essentially the two areas of flare stars, and interferometry of the sun.

In addition to the professional observatories mentioned above, South Africa and Zimbabwe have numerous private observatories, built and operated by amateur astronomers.

Observatories Open to the Public

S.A.A.O. Headquarters, Observatory, Cape are open to visitors on the second Saturday of each month at 8.00 pm. It is unnecessary to give prior notification of intention to visit, unless there are more than ten persons in a party. Day visits are possible to the S.A.A.O. observing site near Sutherland, and enquiries should be made to Sutherland prior to the intended visit.

Visiting nights at Boyden Observatory are generally held around the time of first quarter. Enquiries should be made to the Observatory. Numbers are restricted to twenty persons on each ocassion.

Planetaria

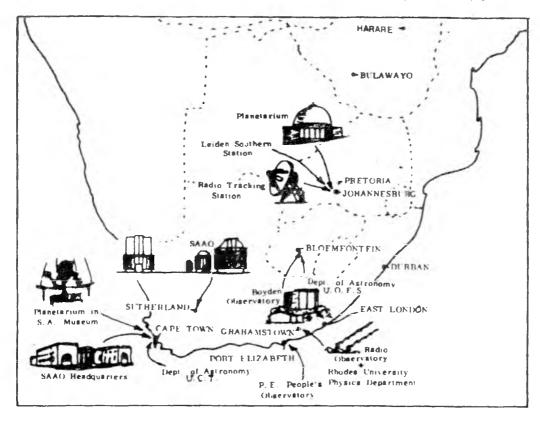
The major planetarium in South Africa is that situated on the grounds of the University of Witwatersrand (entrance in Yale Road, alongside MI). It is equipped with a highly complex Zeiss projector and seats over 400 persons.

A small planetarium, with a Spitz projector and seating approximately 70, is located within the South African Museum, Cape Town. Shows are given each weekend, at 3.00 pm on Saturday, and 3.30 pm on Sunday, and at 11.00 am and 3.30 pm on public holidays. Further information can be obtained by phoning the museum at 41 2668.

Astronomy in Southern Africa

Teaching Departments

Both the University of Cape Town and University of the Orange Free State have departments of astronomy. The chair of astronomy at U C T is occupies by Prof Brian Warner, whose department uses the S.A.A.O. observing facilities at Sutherland. Professor G F R Ellis of the Department of Applied Mathematics, U C T heads a group carrying out research in theoretical cosmology. The U O F S department, incorporated with the Boyden Observatory is headed by Prof A H Jarrett. The Physics Department of Rhodes University specialises in radio astronomy, and has its own observatory outside Grahamstown. The Department of Mathematics, Applied Mathematics and Astronomy at U N I S A offers a number of courses in astronomy and astrophysics.



THE ASTRONOMICAL SOCIETY OF SOUTHERN AFRICA

The Astronomical Society of Southern Africa is a body consisting of both amateur and professional astronomers. Membership is open to all interested persons, regardless of knowledge or experience. In addition to this handbook, the Society issues "The Monthly Notes of the Astronomical Society of Southern Africa" (MNASSA). Members also receive copies of "Sky and Telescope", an excellent and very popular monthly magazine published in the United States. It provides up to date information on both professional and amateur activities, together with news of space research and other related subjects. The Society's annual subscription is R25-00 and there is an entrance fee of R5-00. Information can be obtained from the Honorary Secretary, Astronomical Society of Southern Africa, c/o The South African Astronomical Observatory, P O Box 9, Observatory, 7935, Cape.

ASTRONOMY IN SOUTHERN AFRICA

LOCAL CENTRE OF THE SOCIETY

Autonomus local Centres of the Society hold regular meetings in Cape Town, Durban, Johannesburg, Bloemfontein, Pietermaritzburg, Pretoria and Harare. Visitors are very welcome at meeting and may, if they wish, join a Centre, without becoming a full member of the Society. Centre members do not receive society publications nor Sky & Telescope.

CAPE CENTRE (Cape Town) - meetings on 2nd Wednesday of the month (except January and December) at the South African Astronomical Observatory at 8.00 pm.

The Centre possesses a small observatory housing the 30cm Ron Atkins Telescope. Secretarial address: c/o S.A.A.O., P. O. Box 9, Observatory, 7935. Information on meetings also available from telephone (day time) 69 8531 Ext 256, 2102532 evenings 614534.

TRANSVAAL CENTRE (Johannesburg) - general meetings are held on 2nd Wednesday of each month excluding December in the Sir Herbert Baker Bulding, Republic Observatory, Observatory at 20h00 when a formal lecture is delivered. On this site the Christos Papadopoulos dome houses a combined instrument, comprising of 30 cm cassegrain telescope arranged for photo-electric photometry, and two refracting telescopes of 18 and 15 cm aperture, which is available to members of the Centre. Public observing evenings are held occasionally and there are sections catering for the observation of variable stars and grazing occultations. Secretarial Address: W G C Jacobs, Private Bag X7, Parkview. 2122 Telephone: Oll/6465959

NATAL CENTRE (Durban) - regular monthly meetings are held on the 3rd Wednesday of each month in St. Paul's Church Hall, Church Street, Durban (opposite the GPO) at 7.45 pm. The Natal Centre publishes a monthly booklet "NDABA", which contains news and views and current information on astronomical and related topics. Secretarial Address: c/o P O Box 840, Durban. 4000 Telephone: 842321 (W) and 844751 (H)

NATAL MIDLANDS CENTRE (Pietermaritzburg) - meetings are held on the 2nd Wednesday of the month (except January) at the College for Further Education and Training, Havelock Road at 7.45 pm and occasionally at private homes. Information on meetings and membership is available from the Secretary, P. O. Box 330, Hilton 3245. Telephone No: 0331/24 074 or 54 038

BLOEMFORTEIN CENTRE - meetings are held every 3rd Wednesday of the month. For further information contact Ar F C Neser, Telephone 051/221142, 108 Japie Neser Street, Uitsig.Bloemfontein.

PRETORIA CENTRE - meetings are held on the 4th Wednesday of the months - January, March, May, July, September and November at 8.00pm, the venue alternating between the Main Academic Building of the University of South Africa and the Christian Brother's College, Silverton Road. The Centre's observatory containing a 32cm reflector is situated on the latter site. For further information contact the Secretary, Mr J C Bennett, 90 Milan Street, Riviera, Pretoria,0084.Telephone 012/704895

HARABE CENTRE— the centre holds fairly frequently meetings, usually at 8.00 pm at which talks on various subjects are given and/or films shown. In addition, social "star-gazing" sessions are arranged at intervals, at which telescopes are set up by those members who possess them and made available for observing by all members present. The address of the Harare Centre is P. O. Box UA 428, Union Avenue, Harare and the Hon. Secretary (to whom communication should be addressed) is Mr W 1 Stedman 3 Essex Road, Awandale, Harare.

OBSERVING SECTIONS OF THE SOCIETY

These sections exist to co-ordinate and encourage constructive observing programmes. Mention of the type of observations and equipment involved are made in the appropriate parts of this handbook together with the names and addresses of the directors.

Comets and Meteors	see page 27
Grazing Occultations	see page 38
Nova Search Section	see page 32
Ordinary Occultations	see page 35
Variable Stars	see page 33

DEARY OF PHENOMENA 1984

COMFIGURATIONS OF SUR, MOON AND PLANETS

3800.007			
4 1		APRIL	
30 00 00 30 02 30 23 12 30 30 23 31 23 31 20	MEW MOON Earth at perihelion Venus 1°.8 N of Antares Venus 1°.8 N of Uranus Neccury stationary FIRST QUARTER FULL MOON Juniter 0°.9 S of Nepture Neccury greatest elong, W (24°) Spica 7° S of moon LAST QUARTER Mars 1°.6 S of Moon Venus 0°.03 N of Nepture Saturn 0°.2 S of Moon Venus 0°.8 N of Juniter Uranus 0°.2 s of Moon Venus 0°.8 N of Moon Jupiter 1°.8 N of Moon Venus 3° N of Moon Venus 3° N of Moon Mercury 3° N of Moon Mercury 3° N of Moon Mercury 3° N of Moon	d h 1 14 2 16 2 02 3 05 5 04 1 12 02 1 14 1 5 2 7 01 1 8 00	NEW MOCH Neuture stationary Mercury 5° N of Moon Mercury greatest elong. E (19°) Mars stationary Vesta 1°.2 S of Moon FIRST QUARTER Mercury stationary Soica 6° S of Moon FULL MOON Saturn 0°.6 N of Moon Occ Mars 0°.04 S of Moon Occ (visible north of lire Grootfontein - Durban) Uranus 0°.6 N of Moon Neptune 3° N of Moon Pluto at apposition Jupiter 3° N of Moon Mercury in inferior conjunction LAST QUARTER
	Vesta stationary	29 22	Jupiter stationary
FEBRUAR	Y	30 02	Mercury 0°.7 N of Venus
d h		MAY	
2 02 6 1: 9 2! 10 06 15 15 17 03 20 16 22 16 22 16 23 19 24 12 25 08 25 22 26 10 29 05	Uranus 0ª,2 N of Moon Occ"	d h 1 06 3 10 4 15 8 14 11 12 23 14 10 14 21 15 06 16 13 17 22 18 19 19 13 19 22 22 20 28 20 30 19	NEW MOON Saturn at opposition Mercury stationary FIRST QUARTER Mars at opposition Spica 6° S of Moon Saturn 0°.5 N of Moon Occ Mars 2° S of Moon FULL MOON Penumbral Eclipse Uranus 0°.6 N of Moon Uranus 0°.6 N of Moon Jupiter 3° N of Moon Mars closest approach Mercury greatest elong. W (26°) LAST QUARTER Mercury 1°.0 S of Moon Occ NEW MOON Eclipse
_	NEL MAN	•••	new room acrigat
2 21 8 19 10 16 10 29 17 12 18 08 13 03 20 20 21 15 22 20 22 22 24 05 24 23 30 14	Mercury in superior conjunction Vesta 1°.1 S of Moon Occ FIRST QUARTER FULL MOON Uranus stationary Spica 6° S of Moon Equinox Saturn 0°.6 N of Moon Occ Mars 0°.4 N of Moon Occ Uranus 0°.5 N of Moon Occ Ceres in conjunction with Sun	JUNE d h 2 000 6 19 9 06 10 15 10 15 12 20 13 03 13 17 14 06 15 00 16 01 20 12 21 07 21 08 21 13 23 04 29 05 29 18	Uranus at opposition FIRST QUARTER Spica 6° S of Moon Saturn 0° 2 N of Moon Occ Mars 4° S of Moon Uranus 0° 5 N of Moon Occ Mercury 5° N of Aldebaran FULL MOON: Penumbral Eclipse Neptune 3° N of Moon Jupiter 3° N of Moon Jupiter 3° N of Moon Venus in superior conjunction Mars stationary Solstice Neptune at opposition LAST QUARTER Mercury in superior conjunction NEW MOON Jupiter at opposition

DIARY OF PHENOMENA 1984

CONFIGURATIONS OF SUN, MOON AND PLANETS

JUI	LY		OCT	OBER	
d	h		d	h	
3	08	Earth at aphelion	1	05	Neptune 4° N of Moon
3	16	Hercury 5° S of Pollux	1	16 00	Jupiter 3° N of Moon FIRST QUARTER
5	23	FIRST QUARTER Spica 6° S of Moon	2	16	Mars 3° S of Reptune
6 7	12	Saturn 0°.1 N of Moon Occ	8	19	Venus 3° S of Saturn
á	00	Mars 4° S of Moon	10	02	FULL MOON
10	01	Uranus 0°.4 N of Moon Occ	10	20	Mercury in superior conjunction
		(visible south of line	14	01	Mars 1°.9 S of Jupiter
11	15	Port Nolloth - Komatipoort) Neptune 3° N of Moon	21	23	LAST QUARTER Vesta 0°.4 5 of Moon Oc.
H	22	Pallas stationary	23	23	Spica 5° S of Moon
12	01	Juniter 3° N of Moon	24	14	NEW MOON
13	04	FULL MOON	25	10	Pluto in conjunction with Sun
14	06	Saturn stationary	26	17	Pallas stationary
16	04	Pluto stationary	27	02 08	Venus 0°.3 S of Moon Occ ⁿ Uranus 1°.3 N of Moon
21 26	06 07	LAST QUARTER Mercury 0°.8 S of Regulus	27	18	Venus 3° N of Antares
28	14	NEW MOON	28	15	Neptune 4° N of Moon
30	09	Mercury 7° S of Moon	29	07	Jupiter 3° N of Moon
			29	22	Mercury 3° S of Saturn
AUGU	ST		29	23	Mars 2" N of Moon Venus 1°.5 S of Uranus
a	t.		30 31	15	FIRST QUARTER
1		Marcury greatest elong. E (27°)			
2	16	Spica 6" S of Moon Saturn 0".3 N of Moon Occ	d	EMBER h	(
4	05	Saturn 0°.3 N of Moon Occ" FIRST QUARTER	6	00	Spica 5° S of Moon
5		Mars 3° S of Moon	8	20	FULL MOCN Penumbral Eclipse
6	06	Uranus 0°.5 N of Moon Occ	10	14	Ceres at opposition
7	07	Venus 1°.1 N of Regulus	1.1	09	Saturn in conjunction with Sun
7	1.7	Neptune 3° N of Moon	13	00	Hercury 2° N of Antares
7 8	17 03	Juno in conjunction with Sun Jupiter 2° N of Moon	13 15	21 16	Venus 3° S of Neptune Mercury 2° S of Uranus
10	04	Vesta in conjunction with Sun	16	09	LAST QUARTER
ii	18	FULL MOON	19	05	Yesta 1°.1 N of Moon Occ ⁿ
14	03	Mercury stationary	20	09	Spica 5° S of Moon
16	20	Mercury 6° S of Venus	23	01	NEW MOON Eclinse
18		Uranus stationary	24	16	Mercury 0°.1 S of Moon Occ ⁿ Venus 2° S of Jupiter
19 26	22	LAST QUARTER NEW MOON	24 25	23 02	Meptune 4" N of Moon
28	05	Venus 4° S of Moon	25	20	Mercury greatest elong. E (22°)
28	17	Mercury in inferior conjunction	26	01	Jupiter 4° N of Moon
29	25	Jupiter stationary	26	03	Venus 1°.6 N of Moon
30	02	Spica 5° S of Moon	27	23	Mars 4° N of Moon
31	11	Saturn 0°.5 N of Moon Occ	30	10	FIRST QUARTER
SEPT	EMBI	ER .	DEC	EMBER	₹
d	h		d	h	
2	09	Mars 1°.7 S of Moon	2	17	Mercury 3° S of Meptune
2	12	FIRST QUARTER	4	23	Mercury stationary
2	12	Uranus 0°.8 N of Moon Occ ⁿ	5	20	Uranus in conjunction with Sun
3	05	Mars 2° N of Antares	6	17	Mercury 1°.8 S of Meptune
3 4		Neptune 4° N of Moon Mercury 3° S of Regulus	8	13 16	FULL MOON
4	07	Jupiter 3° N of Moon	15	17	Mercury in inferior conjunction LAST QUARTER
4		Mars 2° S of Uranus	17	16	Spica 6° S of Moon
6	10	Mercury Stationary	19	23	Saturn 1°8 N. of Moon
6	- 11	Pallas at opposition	21	18	Solstice
8 10	14	Mercury 1°6 S. of Regulus Neptune Stationary	22	14	NEW MOON
10	01	Neptune Stationary FULL MOON	22	21	Neptune in conjunction with Sun-
14	03	Mercury greatest elong. W.(18*)	24	00	Jupiter 4° N. of Moon Mercury 3° N. bf Uranus
18	12	LACT NHADTED	24	19	Hercury Stationary
19	13		26	04	Hercury Stationary Venus 3° N. of Moon
22		Ceres Stationary	26	0.9	Mercury 3° N. of Uranus
25		Equinox NEW MOON	27 28	02	Mercury 3° N. of Uranus Hars 4° N. of Moon Pallas 1°2 N. of Moon Occ*
27	12	Venus 2° S. of Moon	30	07	FIRST QUARTER
28	00	Saturn 0°9 N. of Moon Occ".			·
29	21	Branus 1°1 N. of Moon Gcc".			

THE SUN 1984

Basic Data

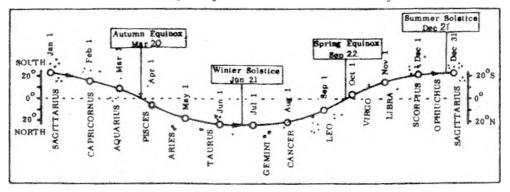
Diameter: 1 392 000 km (1 09 times Earth diameter) Mass: 1.99×10^{30} kg (330 000 times Earth mass)

Surface Temperature: Approx. 6000°C

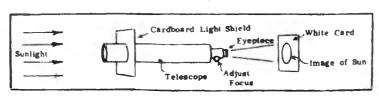
Temperature at centre: Apprx, 10 million°C

The Sun is our nearest star. It is composed chiefly of hydrogen and is in a gaseous state throughout. So hot and dense is its interior that nuclear reactions occur there - thus producing the energy that is eventually radiated from its surface. At times its surface is disturbed by sunspots (which may persist for some weeks) and flares (short lived).

The Earth's orbit round the Sun is not quite circular. In 1983 we will be closest to the Sun on January 2 (perihelion - approx. distance 147 million km) and furthest from the Sun on July 6 (aphelion - approx. 152 million km). During the year, the Sun appears to us to make a complete circuit of the sky (i.e. relative to the starry background) as indicated in the diagram.



Permanent damage to the eye can be caused by looking directly at the Sun. The diagram below shows how a small telescope (or half a binocular) may be used to project an image of the solar disk onto a piece of white card. It may also be advisable to stop down the telescope aperture so that the eyepiece is not damaged by the intense light passing through it. Tiny black sunspots are generally visible on the otherwise white solar disk - if monitored over a period of a week or so, the rotation of the Sun should be apparent.



SUN'S DECLINATION SUN'S DECLINATION SUN'S															
Jan	1	23	05 S	Apr	5	6	04 N	Jul	4	22	53 N	0ct	2	3	34 S
	11	21	58 S		15	9	45 N		14	21	41 N		12	7	23 5
	21	20	08 S		25	13	11 N		24	19	53 N		22	11	02 S
	31	17	40 S	May	5	16	15 N	Aug	3	17	31 N	Nov	1	14	25 S
Feb	5	16	14 5		15	18	52 N		13	14	41 N		11	17	25 S
	15	13	02 S		25	20	57 N		23	11	28 N		21	19	54 S
	25	9	28 \$	Jun	4	22	26 N	Sep	2	7	56 N	Dec	1	21	48 S
Mar	6	5	4 0 S]4	23	16 N		12	4	12 N		11	23	00 S
	16	i	45 S		24	23	25 N		22	0	20 N		21	23	26 5
	26	2	12 N										31	23	06 S

THE SUN 1984

TIMES OF SUNRISE AND SUNSET FOR THE MAIN CITIES OF SOUTHERN AFRICA

		CAPE TOWN	ÖÜRBAN	BLOEMFONTEIN	JOHANNESBURG	HARARE
	_	Sunrise Sunset	sunrise sunset	sunrise sunset	sunrise sunset	sunrise sunset
Jan	1 11 21	05 46 20 02 05 55 19 59	04 ⁿ 58 ^m 19 ^h 01 ^m 05 06 19 02 05 14 19 00	05 ⁿ 21 ^m 19 ⁿ 18 ^m 1 05 29 19 18 05 37 19 17	05 ^h 18 ^m 19 ^h 04 ^m 05 25 19 05 05 33 19 04	05 ⁿ 24 ^m 18 ^r 35 ^m 05 29 18 37 05 37 18 38
Feb	1	06 07 19 52	05 24 18 55	05 46 19 13	05 42 19 00	05 42 18 36
	11	06 17 19 44	05 32 18 48	05 54 19 06	05 49 18 55	05 47 18 32
	21	06 26 19 33	05 41 18 39	06 02 18 57	05 54 18 47	05 52 18 27
Mar	1	06 33 19 23	05 46 18 3 0	06 08 18 48	06 00 18 39	05 55 18 21
	11	06 41 19 11	05 53 18 19	06 13 18 38	06 04 18 29	05 57 18 15
	21	06 49 18 58	05 59 18 06	06 18 18 27	06 11 18 19	06 00 18 06
Apr	1	06 58 18 41	06 06 17 53	06 25 18 13	06 17 18 06 1	06 02 17 57
	11	07 04 18 30	06 11 17 43	06 30 18 03	06 21 17 56 1	06 04 17 50
	21	07 13 18 17	06 17 17 31	06 35 17 52	06 25 17 47	06 07 17 43
May	1	07 20 18 06	06 24 17 22	06 42 17 44	06 31 17 38 1	06 10 17 37
	11	07 28 17 57	06 31 17 14	06 49 17 36	06 37 17 31	06 13 17 32
	21	07 34 17 50	06 36 17 06	06 54 17 30	06 41 17 26	06 16 17 29
Jun	1	07 43 17 45	06 43 17 04	07 01 17 27	06 47 17 23	06 20 17 28
	11	07 48 17 44	06 48 17 03	07 05 17 26	06 52 17 22	06 23 17 27
	21	07 51 17 44	06 51 17 04	97 06 17 27	06 55 17 24	06 26 17 29
Jul	1	07 53 17 48	06 53 17 07	07 10 17 30	06 57 17 27	06 27 17 32
	11	07 51 17 52	06 51 17 11	07 06 17 34	06 55 17 30	06 27 17 35
	21	07 47 17 58	06 48 17 16	07 05 17 39	1 06 53 17 35	06 26 17 40
Aug	1	07 39 18 06	06 42 17 22	07 00 17 45	06 48 17 41	06 23 17 42
	11	07 30 18 13	06 34 17 29	06 53 17 51	06 41 17 46	06 18 17 46
	21	07 19 18 20	06 24 17 35	06 42 17 55	06 32 17 50	06 11 17 48
Sep	11 21	07 06 18 27 06 52 18 34 06 38 18 41	06 12 17 40 06 00 17 46 05 48 17 51	06 31 18 01 06 19 18 06 06 07 18 10	06 21 17 54 06 11 17 59 05 59 18 03	06 04 17 49 05 55 17 51 05 46 17 52
0ct	11 21	06 25 18 48 06 12 18 55 05 58 19 04	05 37 17 57 05 25 18 03 05 12 18 09	05 57 18 16 05 45 18 22 05 33 18 27	05 50 18 08 05 39 18 12 05 27 18 17	05 39 17 54 05 30 17 57 05 23 17 59
Nov	1	05 46 19 13	05 02 18 17	05 24 18 35	05 19 18 24	05 16 18 03
	11	05 38 19 23	04 55 18 26	05 17 18 44	05 13 18 32	05 14 18 08
	21	05 31 19 33	04 49 18 34	05 12 18 52	05 06 18 39	05 11 18 13
Dec	1	05 29 19 43	04 48 18 42	05 11 19 00	05 07 18 46	05 12 18 19
	11	05 28 19 50	04 48 18 50	05 11 19 07	05 08 18 53	05 14 18 25
	21	05 32 19 57	04 52 18 57	05 15 19 14	05 12 19 00	05 18 18 31

Solar Eclipsés

The Annular Eclipse of the Sun on May 30 will be seen from Mexico, the south-eastern United States, Morocco and Algeria, while a partial eclipse will be seen in a belt some 30° on either side of this line. From Southern Africa nothing of the eclipse will be seen.

About midnight on the 22 - 23 November there will be a total eclipse of the Sun which will be visible from New Guinea and the South Pacific Ocean. Partial phases eill be seen from Australia and New Zealand and from Antartic and the southern tip of South America.

The Moon 1984

BASIC DATA

3 480 km (0,27 of Earth) 7,35 x 10²²x kg (1/81 of Earth) Diameter:

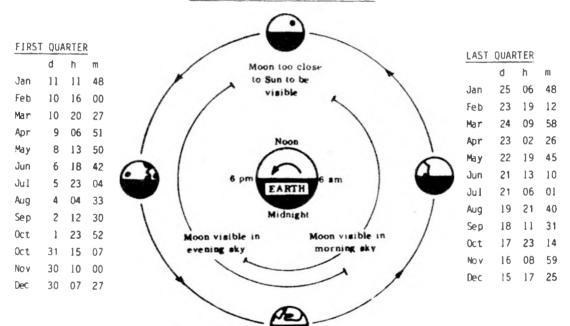
Mass:

Surface Gravity: 0,16 of Earth

Average distance from Earth 384 000 km

PHASES AN	D V	SIB	LITY		NEW	MOC)N					
	đ	ħ	m		d	h	តា			d	h	m
Jan	3	07	16	May	1	05	45	:	Sep	25	05	11
Feb	2	01	46	May	30	18	48	1	0ct	24] 4	08
Mar	2	20	31	Jun	29	05	18	1	Nov	23	00	57
Apr	1	14	10	Jul	28	13	51	ı	Dec	22	13	47
				Aug	26	21	25					

SCHEMATIC DIAGRAM OF MOON'S ORBIT



FULL	MOC	N									
	d	h	m		d	h	m		d	h	m
Jan	18	16	05	May	15	06	29	Sep	10	09	01
Feb	17	02	4]	Jun	13	16	42	0ct	10	01	58
Mar	17	12	10	Ju l	13	04	20	Nov	8	19	43
Apr	15	21	11	Aug	11	17	43	Dec	8	12	53

THE MOON'S ORBIT

Dates of Apogee, when the Moon is furthest from the Earth (approx. 407000 km) and of Perigee, when the Moon is closest to the Earth (approx. 357000 km) are given below.

	M	on at	Perigee					Moon at	Apogee		
Jan	d 20	h 0	Jul	d 3	h }	Jan	d 7	h 22	Jul	d 18	h 16
Feb	17	13	Ju l	30	14	Feb	4	11	Aug	15	7
Mar	16	23	Aug	27	19	Mar	2	13	Sep	11	15
Apr	14	8	Sep	25	5	Mar	29	18	Oct	8	19
May	12	5	0ct	23	16	Apr	26	9	No v	5	1
Jun	7	13	Nov	20	23	May	24	3	Dec	2	17
			Dec	18	12				Dec	30	14

As a result of its motion around the Earth, the Moon appears to make a complete circuit pf the heavens in just under a month. It occasionally passes in front of bright stars (details given in Occultation section - page 35) and close to visible planets (details given in Diary pages 4 and 5).

TIMES OF MOONRISE AND MOONSET

Times for Bloemfontein, Cape Town, Durban, Johannesburg and Port Elizabeth can be obtained from the tables on pages 11 to 16.

THE SURFACE OF THE MOON

In common with the inner planets of our solar system, the Moon's surface suffered bombardment by numerous minor bodies during a period 4,5 to 3,0 billion years ago. This had produced the heavily cratered topgraphy now visible. Some particularly large impacts caused large circular depressions, which were flooded by molten lava from the Moon's interior. These are the maria basins which appear smoother and darker than the rest of the surface (the latin words mare and maria come from older times when they were mistaken for seas). The maria surfaces, being younger have fewer large craters, but the entire surface is peppered with tiny craters produced by tiny bodies which have also served to plough up the ground thus forming the regolith - a layer of loose material a metre or so deep.

Lunar Eclipses

There will be no full eclipse of the Moon in 1984. Three penumbral eclipses will occur.

On May 15 the Moon enters the penumbra (half shadow) at 4.42 S A S T and leaves at 8.38. This phenomenon will be seen from all of Southern Africa till interupted by moonset.

On June 13 the penumbral eclipse will occur in mid-afternoon and so will not be seen in our longitudes.

When the Moon rises on November 8 it will already be in the partial shadow and it will leave the shadow completely by 22.12.

THE MOON 1984



LIBRATION

Due to the fact that the Moon's orbit is an elipse the speed in the orbit varies and this combined with the inclination of its orbit has the effect that we see rather more than half the entire surface of the Moon. This effect is termed Libration. Below are tabulated the days of maximum libration and the amount and direction of each libration. The direction is measured from North eastwards.

Date	Size o	P . A	Date	Size	P . A .	Date	Size	P . A .
Jan 12	7.9	5.1	May 8	7.5	144	Sept 6	8.0	312
2 4	8.1	227	2 1	7.4	327	2 1	8.9	130
Feb 11	7.9	8 0	June 5	7.0	160	Oct 3	8.9	308
2 2	8.0	2 4 5	1 9	7.2	3 3 5	19	9.1	135
Mar 12	8.1	111	July 3	6.7	175	3 1	9.2	312
2 3	7.5	277	1 5	7.0	3 3 3	Nov 15	8.6	138
Apr 10	8.0	4.3	2 7	7.0	127	2.8	8.9	3 1 7
2 3	7.4	3 1 9	Aug 11	7.2	328	Dec 13	7.8	148
			2.3	8.1	118	2 6	8.3	3 2 4

	4	914	49	19	49	19	49	22	58	39	92		20	7	28	33	47	56	0	10	-	2	2	2	7		5	.4		5	7	7 5
MARCH	Se		18	19	19	20	20	21	21	22	23		00	0	05	03	04	05	07	08	60	10	Ξ	12	13	10	14	-	7	=		_
M	Rise	04 41 11	05 54	06 26	07 17	60 80	00 60	09 53	10 49	11 46	12 46	13 47	14 47	15 44	16 37	17 24	18 07	18 47	19 25	20 04	20 44	21 26	22 11	23 00	23 52		4	4	3	03 29	2	2
ARY	Set	19 ^h 00 ^m	19 34	20 15		21 17	_	2	N	0		00 00	00 45	01 36	02 36	03 42	04 53	06 04	07 14	08 21	92 60	10 30	11 32	12 34	13 34	14 31	15 25	16 15	17 00	17 40		
FEBRUARY	Rise	O4 hs 7m	05 52				_	_			12 56	13 56	14 58	16 01	17 03	17 59	18 51	19 36	20 18	20 56	21 33	22 11	22 50	23 32		0	01 07	5	2 5	3.4		
JANUARY	Set	17h 45m											00 17														13 41		15 39	16 35	17 28	18 17
JAN	Rise	03h 36m	04 24	05 15	60 90	07 04	07 59	08 52	09 45	10	_	12	13	14	15	16	17	18	19	20	21	21	22	23	23		000	00 2	01 3	02 21	03 1	04 0
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	Set	8 ^h 5	9 6	1 0	22 09	3		00 18	01 20	02 22	03 25	04 29	05 33	06 36	07 38	08 34	09 25	10 10	10 49	11 23	11 55	12 24	12 54	13 23	13 55	14 29	15 09	15 54	16 47	17 46	18 51
JUME	Rise	08h27m	12 60	10 22	11 12	11 56	12 36	13 14	13 49	14 25	15 03	15 43	16 28	17 17	118 111	19 07	20 05	21 02	21 57	22 50	23 42		3	2	-	_	_	_	genera	07 15	_
	Set	745	3	9 2	0 0	21 05	2 0	3		91 00	01 21	02 25	03 29	04 33	05 38	06 44	07 49	08 52	09 52	10 46	11 33	12 15	12 51	13 24	13 55	14 24	14 54	15 24	15 57	16 34	17 15
MAY	Rise	49	~	80	6	10 35					-				17 01			-	-	_	N	\sim		10 00	5	5	4	(~)	C		06 24
	Set	8 h2	00	9 2	5	20 37	1 2	2	3		00 13	01 19	02 27	03 34	04 41	05 47	06 53	07 59	09 04	10 08	11 09	12 05	12 55	13 39	14 17	14 52	15 24	15 54	16 23	16 53	17 24
APRIL	ise	6 h04m	99 9	7 49	8 44	9 41	0.40	1 40	2 39	3 36	4 28	5 16	5 59	6 39	17 17	17 55	18 34	91 61	10 02	20 50	21 42	22 37	23 33		2	2	-	3 0	G,	04 50	S

TIMES OF MOON RISE AND

	JULY	>	AUGUST		SEPTE	SEPTEMBER
	Rise	Set	Rise	Set	Rise	Set
	m70460	9 ^h 5	946		10 ^h 20 ^m	
. ~	09 54	21 05	10 27	23 12	11 05	00 12
. ~	0	2	0		11 53	01 15
<	_	3	1	00 14	12 45	ni
- 40	11 51		2	01 17	13 41	197
9	12 26	_	(*)	02 19	14 37	03 59
1	13 03	_	13 57	03 21	15 33	04 43
00	13.41	- CVI	14 50	04 19	16 29	05 22
5	14 23	(4	15 46	05 12	17 22	05 56
01	15 10	. 4	16 43	06 01	18 14	06 28
-	16 01		17 39	06 43	19 05	06 57
12	16 56		18 34	07 21	19 56	07 26
2	17 54		19 27	07 55	20 47	07 55
14	18 51	08 04	20 19	08 25	21 40	08 25
15	19 47	_	21 09	08 55	22 35	08 57
16	20 41		22 00	09 27	23 31	09 33
17	21 34		22 52	09 53		•
18	22 25		23 46	10 24	00 30	
6	23 16			10 58	62 10	
20			ব	11 36	02 26	
21	U		A	12 21	03 20	
22	0		a	13 13	04 10	
23	Œ1		A	14 12	04 55	
24	LL 1		42	15 17	05 36	
25	421		6.1	16 27	06 15	
56			-	17 36	06 54	
27	_		_	18 45	07 32	
28			_	19 53	08 14	
52	07 47	00	08 23	20 58	08 58	
30	മ		_	22 04	09 46	
31	ch.			23 08		

The Moon 1984

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DECEMBE	e	m64	40	30	22	15	_	60	60	08	90	57	44	56		0	4	_	5	2	-	41	41	4	4	121	4	4	*	31	
	Ris	22	3	14	15	16	17	18	19	20	21	23	22	23		00	00	0	0	02	03	03	04	90	90	07	08	60	10	=	12
		22m	56	33	03	32	00	53	00	34	14	99	45	4	4	44	48	52	56	00	05	12	20	28	35	91.	30	17	57		32
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OCTOBER		130 H		-	- 00	4	11	0	10	52	47	36	30	26	24	22	,	6	13	02	47	28	07	45	23	03	47	34	27	23	25
0	Rise	0	-	12 3	13.2	14	15	16	1	17	18	6	20	21	22	23		00	0	05	0.5	03	04	0.4	0.5	90	90	07	080	60	10

					1000	ONE SAN	MOON CET	CADE T	TOUN			
				IMES	UF MUUN K	KISE AND						
				FOR P	PORT ELIZ	IZABETH SU	JBTRACT 2	8 MINUTES	S			
	JAN	JANUARY	FEB	FEBRUARY	MAI	MARCH	APRIL	III	£	MAY	JUNE	
	Rice	Set	Rise	Set	Rise	Set	Rise	Set	Rise	Set	Rise	Set
	Meallen	40	mr.rhan	1 chekim	W	197	06 h 42m	18 ^h 59 ^m	07 ^h 26 ^m	18 ^h 26 ^m	09 ^h 26 ^m	19h16#
	03 57	10 40	16 16	20 32	00 90	19 35	07 38	19 25	08 27	00 61	10 26	20 17
	06 45	200	07 12	35		202	08 35	19 53	62 60	19 40	11 20	21 23
	50 50	250	OR OR	7		20	09 33	20 24	10 32	20 28	12 07	22 32
	92 20	21	04 03	7		20	10 34	21 00	11 34	21 23	12.47	23 41
	08 23	22	95 60	3		21	11 36	21 42	12 31	22 24	13 23	
	67 60	35	10.53	22	_	2	12 38	22 31	13 22	23 31	13 56	65 00
	10 16	36	49	23	~	22	13 38	23 28	14 07		14 27	01 57
	11 10	36	2 47	23		23	14 34		14 46	00 40	14 59	03 04
		1	13 47	ì	m	23	15 24	00 32	15 21	01 50	15 32	04 11
25		9		00	S		16 08	01 41	15 54	02 59	16 08	61 90
12	-	88		0	15 46	00	16 47	02 53	16 26	04 07	16 49	06 27
10	-	3 2		0	16 41	10	17 22	04 05	16 59	91 50	17 36	07 33
25		5 2	-	02	17 31	02	17 56	05 16	17 34	06 26	18 28	08 36
1 4	·	3 6		040	18 14	04	18 29	06 27	18 13	07 35	19 25	09 33
2 4		0.5		05	18 53	99	19 04	07 38	18 57	08 44	20 25	10 22
20		0		90	19 28	90	19 6	08.48	19 47	09 80	21 24	11 05
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000		3 8		08 57	20 36	90	21 09	11 05	21 39	11 44	23 20	12 12
200		5 6		2	21 11	0	22 00	12 07	22 38	12 30		12 40
35		88		Ξ	21 50	7	22 55	13 03	25 37	13 09	00 15	13 06
22		10		12	22 32	2	23 52	13 52		=	01 10	13 31
210		=		13	23 19	3		14	00 34	4	02 05	13 57
200		12		7		4	00 20	15	01 30	14	03 05	14 25
25	6	15	00	15	11 00	2	01 48	15	02 25	15	04 01	14 56
36	36	10	30	16	01 05	45	02 44	91	03 21		05 02	15 32
20	35	1	05	-	02 02	1	03 40	-	04 17	=	90 90	16 14
10	5 2	-	03	_	02 59	-	04 35	-	05 15	=	07 11	17 05
10	20	-	04 08	=	03 56	-	05 30	-	91 90	_	08 14	18 04
JM	030	-			04 52	18 07	06 27	-	07 18	17 37	21 60	19 10
3.5	0.0	22 19 14			05 47	-			08 22			

TIMES OF MOON RISE AND MOON SET - CAPE TOWN

FOR PORT ELIZABETH SUBTRACT 28 MINUTES

Dice	SEPTEMBER		
2000	Rise Set	2	Set
	h45m		22h47
=	11 26 01 0		23 55
12	12 02		
13	03 03		01 02
3 14	13 59 04 0		02 09
15	57 04		03 15
91 8	56 05		04 18
4 17	54 06		05 18
5 18	51 06		0.0
3 19	47 07		06 57
9 20	42 07		07 37
14 21	37 08		08 12
9 22	32 08		08 42
55 23	28 08		01 60
54	27 09		09 35
27 00	60		10 00
35 01	27 10		10 26
20 0.2	11		10 53
14 0/2	28 12		11 23
16 0.3	25 13		11 59
24 0.4	18 14		12 40
36 04	04 15		13 30
50 05	45 16		14 29
03 05	21 18		15 36
16 06	56 1		16 49
28 07	29 2		18 03
40 07	03 2		19 17
51 08	40 2		20 29
60	20		21 40
00 01 10 39			22 50
1			23 59

FOR BLOEMFONTEIN ADD 19 MINUTES

	Set	18h35m	19 36	20 41	21 48	22 55		10 00	01 05	02 09	03 14	04 19	05 25	06 30	07 32	08 29	61 60	10 03	10 41	11 15	11 45	12 13	12 40	13 08	13 38	14 11	14 49	15 33	16 24	17 23	18 29
JUNE	Rise	08 ^h 21 ^m	09 21	10 16	11 05	11 48	12 27	13 02	13 36	14 10	14 45	15 24	16 07	16 55	17 48	18 45	19 43	20 41	21 38	22 33	23 26						04 02				
	0,	17h4111	18	18	19	20	21	22	23		0	02	03	04	90	90	07	08	50	0	Ξ	2	2	13.15	13 44	14 12	14 40	15 09	15.40	16 15	16 55
MA	Rise	06 h 28 m	07 27	08 27	65 60	10 29	11 27	12 19	13 06	13 48	14 26	15 01	15 36	16 11	16 43	17 30	18 16	19 06	20 01	20 58	21 56	22 53	23 48		-					•	91 90
	Set	18h08m	18 36	19 07	19 40	20 17	21 01	21 51	22 48	23 51		85 00	02 07	03 16	04 25	05 33	06 41	07 49	98 80	10 02	11 03	11 59	12 49	13 32	14 10	14 44	15 14	15 43	16 11	16 39	17 08
APKIL	Rise	05 ^h 50 ^m	06 44	07 38	08 35	09 33	10 33	11 34	12 34	13 30	14 22	15 08	15 50	16 28	17 04	17 40	18 17	18 57	19 40	20 28	21 20	22 14	23 11				01 57				
T.	Set		18	19	10	2	202	23	21	22	23	23		00	05	03	04	05	90	07	60	10	11	12	13	14	14 53	15.34	16 10	16.42	17 12
MARCH	Rise	04h21m	05 16	90 90	07 07	07 55	08 48	09 43	10 40	11 39	12 40	13 41	14 42	15 38	16 30	17 15	17.57	18 35	11 61	19 48	20 26	21 06	21 50	22 38	23 30			-	02 15		
		18	0	20	200	25	3	22		23	23	1	00	10	02	03	04	050	90	08	60	2	-	12	13	10	15 20	16 10	16 54	17 33	
FEBRUAR	Rise	OA hasm	č	5 6	5 0	0 0) C	- 0			-	-	-		-												00	0	02 30	03	
IRY	Set	17	10	- 0	100	2 6	2 5	200	27	10	20	1	O	č	5 0	0) C	, _	, –	. –							3 32				
JANUARY	Rice	mathren	00 00	20 40	200	14 50	00 40	000	2000	10 01	11 12	12 07	13 00	10 01	15 02	16 06	17 19	18 17													
																			, a	0 0	n ç	2 =	- 2	4 5	3 5	<u>.</u> 18	3 15	25	36	3 5	30

TIMES OF MOON RISE AND MOON SET - DURBAN

MES OF MOON RISE AND MOON SET - DURBAN
FOR BLOEMFONTEIN ADD 19 MINUTES

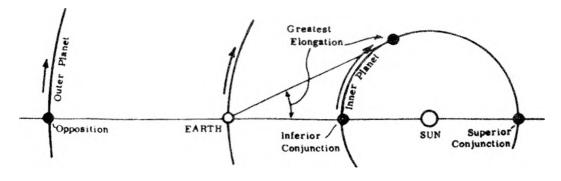
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DECEMBER	Rise	mcrhc1								19 03								NO 28	20	0 35	02 11	02 51	03 37	04 28	05 25	06.26	07 28	08 24	00 27	10 22	11 15	12 08	200
BER	.Set	O1h16m	01 52	25 50	47 70	02 53	03 20	03 47	04 14	04 43	05 15	05 52	06 34	07 22	08 17	09 18	10.22	11 28	12 34	13 40	14 46	15 53	17 02	18 12	19 22	20 29	21 31	22 25	23 11	23 50	1	00 24	
NOVEMBER	Rise	11 h56m	12 53	13.07	2 4 5	14 40	15 33	16 25	17 19	18 15	19 13	20 12	21 12	22 10	23 04	23 53		00 37	01 17	01 53	02 28	03 02	03 38	04 16	05 00	05 49	06 44	07 43	08 44	09 44	10 43	11 38	
2	Set	m10,00	00 10	01 52	00 22	02 37	03 16	03 20	04 21	04 49	05 16	05 43	11 90	06 41	07 14	07 52	08 36	92 60	10 23	11 26	12 32	13 40	14 49	15 57	17 06	18 16	19 27	20 37	21 45	22 48	23 45		400
OCTOBER	Rise	10h16m		12 09	12 07	1000	14 04	14 59	15 53	16 46	17 38	18 31	19 25	20 21	21 19	22 18	23 17		00 14	01 07	95 10	05 40	03 19	03 56	04 32	05 08	05 46	06 27	07 13	08 04	00 60	69 60	01 01
SEPTEMBER	Set		00 02	60 10	00 00	20 00	60 50	03 54	04 37	05 14	05 48	06 18	06 46	07 13	07 40	08 08	08 39	09 14	09 54	10 40	11 34	12 35	13 41	14 51	16 02	17 13	18 22	19 32	20 41	21 50	22 57		
	Rise	10h01	10 44	11 31						16 09	17 04	17 58	18 51	19 43	20 36	21 30	22 26	23 25		00 24	01 24	02 21	03 15	04 03	04 47	05 26	06 03	06 39	07 16	07 55	08 38	09 25	
AUGUST	Set	21 h54m	22 59		00 00	00 04	01 08	02 12	03 15	04 13	05 07	05 55	06 37	07 13	07 46	08 15	08 43	01 60	09 38	10 07	-		_	_	_		_		•			21 52	
	Rise	09h38m	10 13	10 47	11 22	11 23	12	12	13		15	91	17	89	5	50	2	2	2	λí		00	0	05	03	04	05	90	90	0	90	08 45	-
JULY	Set	19h37m						00 03	90 10	02 10	03 15	04	05	90	6	0	0	С	Ö		_				_	_						19 39	
	Rise	mto460		10 27			11 38	12 12	12 46	13 23	14 03	14 49	15	9	17	8	6	20	21	22	23	5		ö	0	0	0	0	0	0		0 08 23	
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THE PLANETS 1984

		BAS	SIC DATA			
	Dist from Sun	Period of Revolution	Mass	Diameter	Rotation Period	Inclination of Equator
	10 ⁶ km	years	(Earth = 1)	10 ³ km		to Orbit
Mercury	58	0,24	0,056	4,98	591	7
Venus	108	0,62	0,817	12,4	244 n m	_?
Earth	150	1,00	1,000	12,8	23 ^h 56 ^m	23°27'
Mars	228	1.88	0,108	6,76	24 37	23 59
Jupiter	778	11,9	318,0	142,7	09 51	03 04
Saturn	1426	29,5	95,2	120,8	10 14	26 44
Uranus	2868	84,0	14,6	47,1	10 49	97 53
Neptune	4494	164,8	17,3	44,6	14 ?	28 48
Pluto	5896	247,6	0,9?	?	6d?	?

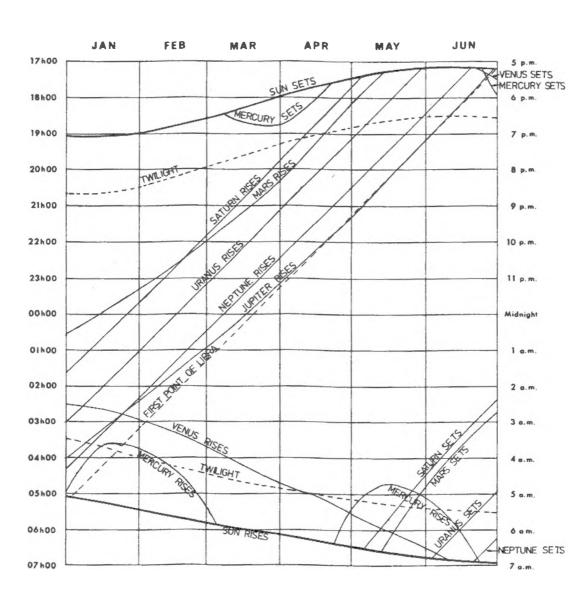
GENERAL.

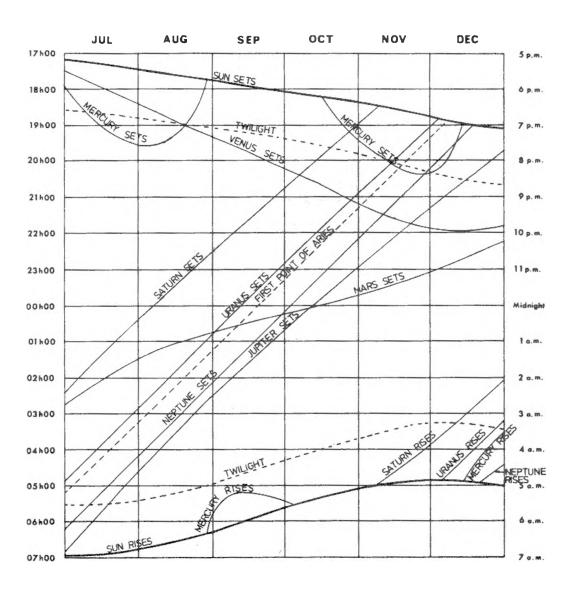
Apart from Uranus, Neptune and Pluto, the planets of our solar system are amongst the brightest objects in the night sky. Unlike the distant stars, their relative positions do not remain fixed, but continually change as, like the Earth, they orbit around the Sun. Their apparent movements against the starry background are complicated as they result from a combination of their own motion and the Earth's motion. Their brightnesses also vary considerably, as both their distances from the Earth and the visible portions of their sunlit hemispheres change. Since the period of a planet increases with increasing distance from the Sun, so we find that the inner planets - Mercury and Venus - appear to "overtake" the Earth in their orbits, while the Earth in turn "overtakes" the outer planets - Mars, Jupiter and Saturn. The terms given in astronomy to the various Sun-Earth-Planet configurations are illustrated in the accompanying diagram. Dates of such configurations occurring in 1984 are listed chronologically in the Diary (pages 4 and 5) and are also mentioned in the text below.



TIMES OF RISING AND SETTING

The times of rising and setting given by the diagram are accurate for position 30° East. 30° South, and approximately correct for other places in Southern Africa. Strictly speaking, corrections for latitude and longitude should be applied, but the latitude correction is, in general sufficiently small to be ignored and in no case will exceed 15 minutes. Longitude corrections are given on page 44.





OBSERVING THE PLANETS

To the naked eye, planets appear as virtually pinpoint sources of light. However, their disks can be readily resolved with the aid of a small telescope. Even so, their angular diameters are of the order of 10 seconds of arc - roughly 1/200 of the Moon's angular diameter - so it is not always possible to distinguish details on their disks. The disks of Mercury and Venus are only seen fully illuminated when they are furthest from us - as they draw closer, their disks grow larger but the phase Changes to a crescent as we see more of their dark hemispheres. In contrast, the disks of the outer planets are always seen fully or near fully illuminated.

Mercury

The innermost planet, Mercury, moves rapidy among the stars and can be seen only in the early evening sky or just before dawn. Its greatest angular distance from the Sun is 28 but at some elongations this may be as little as 18. The best times for viewing the planet are within a few days of Jan 22, May 19 and Sep 14 in the morning sky.

In the evenings the best dates will be near April 13, Aug 1 and Nov 25.

Greatest Elongation East	Apr 3(19 ⁰)	Aug 1(27 ⁰)	Nov 25(22°)
Stationary	Apr 12	Aug 14	Dec 4
Inferior Conjunction	Apr 22	Aug 28	Dec 14
Stationary	May 4	Sep 6	Dec 24
Greatest Elongation West	Jan 22(24°)	May 19(26°)	Sep 14(18 ⁰)
Superior Conjunction	Mar 4	Jun 23	Oct 10

Venus

Venus willbe a morning object from January to the beginning of June when it will pass into the twilight. It is in conjunction with the Sun on June 15 and from the beginning of July it will be an evening sky object for the rest of the year.

Venus at magnitude -3.6 at the beginning of the year varies little during the year fading to -3.3 at mid-year and brightening again to -3.8 in December.

Mars

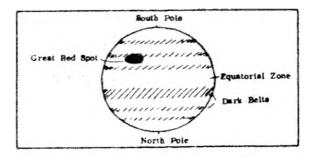
Mars rises after mid-night at the beginning of the year and moves slowly forward till on May 9 it rises at sunset. It will then be visible in the evening sky for the rest of the year. Its magnitude varies from +1.4 in January to -1.7 in May and then fades to +1.1 at the end of the year.

In Virgo in January the planet will pass into Libra at the end of that month, into Scorpius in August and into Ophiuchus at the beginning of September and into Sagitarius a month later and into Capricornus in mid-November and ends the year in Aquarius.

Jupiter

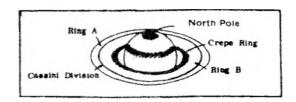
Jupiter emerges from the morning twilight in mid-January (mag -1.4). It rises earlier each night and reaches opposition on June 29 when it will rise at midnight. By then it will have brightened to mag -2.2. It will then be in the evening sky till the end of the year when having faded to mag -1.4 it will pass into the evening twilight.

It will be in Sagittarius for the whole of the year, falling back among the stars till it is stationary on April 29. Thereafter it will advance among the stars till Aug 30 when it is again stationary. It will then fall back again till the end of the year.



SATURN

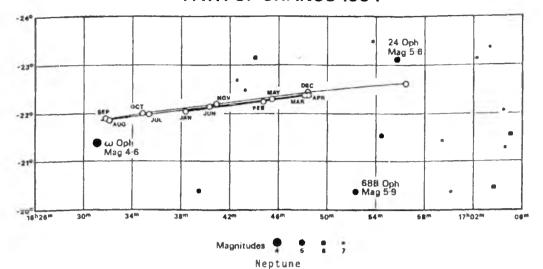
Saturn (mag 1.0) rises at 3 am at the beginning of the four year and moves steadily towards the evening sky till it rises at sunset at the beginning of May and is visible in the evening sky until October when it disappears in the evening twilight. From mid-December it will be in the morning sky. It will at all times be in the constellation of Libra.



Uranus

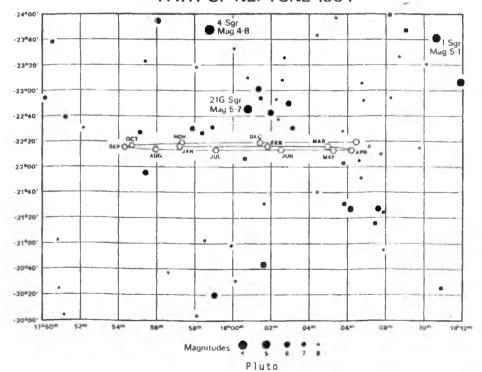
Uranus is in Ophiuchus and will be at opposition on June 2 at mag 5.8. Its diameter then will be 3\$8. Binoculars will assist to identify it.

PATH OF URANUS 1984



Neptune is in Sagittarius all year. It is at opposition on June 21 and its magnitude then will be 7.7. Its drameter varies from 2."35 to 2."5 and on dark nights binoculars will serve to identify this planet.

PATH OF NEPTUNE 1984

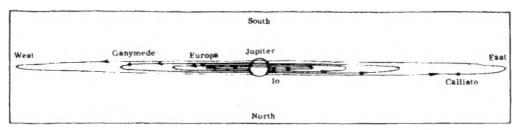


Pluto at mag 14 needs a large telescope before it can be seen. It will be at opposition on April 20.

THE MOONS OF JUPITER AND SATURN 1984

MOONS OF JUPITER

Une of the most popular sights for an observer with a small telescope is Jupiter and its Moons. Four of the fifteen - Io. Europa, Ganymede and Callisto are generally clearly visible - they would just be visible to the naked eye were it not for the glare from the mother planet. As the diagram below indicates. the system



is seen almost edge-on so the moons always lie close to a straight line extending from the planet's equator. As they orbit, so they appear to oscillate from one side to the other, alternately passing in front and behind the planet. This motion is represented in the diagrams on pages 26 which cover the period when Jupiter is clearly visible in the evening sky. The norizontal lines show their relative configurations at 2 am each day.

When the moons pass in front and behind the planet, transits, occultations and eclipses occur. Details of such phenomena occuring between the end of astronomical twilight and just after midnight (and when the planet is above the horizon in Southern Africa) are given in the table below.

Explanation of Table.

Date and predicted times are given - these are for mid-phenomena and are not instantaneous.

The moons concerned are I - lo III - Ganymede II - Europa IV - Callisto

Phenomena - the abbreviations used are:

Eclipse: the satellite passes through the shadow of Jupiter Oc. Occultation: the satellite is obscured by the disk of Jupiter

D. Disappearance R. Reappearance

Ir. Transit: the satellite crosses the disk of Jupiter

Shadow Transit: the shadow of the satellite transits the disk Sh.

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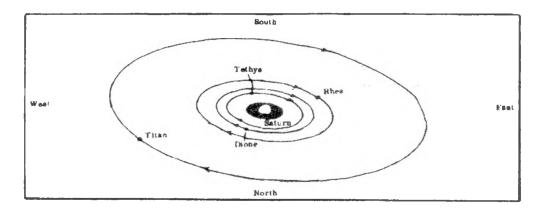
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THE MOONS OF JUPITER AND SATURN 1984

SATURN'S MOONS

Saturn's moons are considerably fainter than the four Galilean moons of Jupiter. The diagram shows the orbits of four of Saturn's ten moons. The easiest to find is Titan (magnitude 8,5), according to the diagram and the information in the table below.

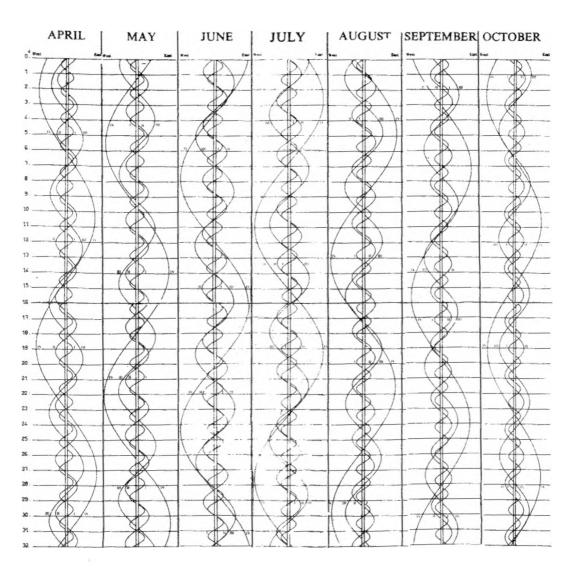


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The Moons of Jupiter and Saturn 1984

CHANGING CONFIGURATIONS OF JUPITER'S MOONS



The four bright moons of Jupiter always appear close to a straight line passing through the planet since, as shown in the drawing at the top, their orbits are seen nearby edge on. The main part of the diagram then shows how their positions along such a straight line change during the five months when Jupiter is prominent in the evening sky. For each month, time increases downward; the disk of Jupiter is stretched to make the central column and horizontal lines, representing 2a.m. (0 hrs Universal time), are shown for every day of the month. The wavy lines show how the Moons appear to oscillate from each side of the planet to the other.

COMETS AND METEORS

COMETS

Comets are celestial bodies moving around the sun, mostly in very elongated orbits. The typical comet consists of a nucleus surrounded by a hazy aura of gas and dust called the coma, and in many cases there is a tail stretching away from the sun. Faint comets, several of which are discovered each year, usually appear only as fuzzy patches without nucleus or tail, though there may be a central condensation.

While they are believed to be true members of the solar system, comets differ radically from the planets in that their orbits, besides being highly eccentric, are inclined at all angles to the plane of the ecliptic, and their motion may be direct (like that of the planets) or retrograde. Compared with that of a planet, a comet's mass is almost negligible; nearly all this mass is concentrated in the nucleus which is believed to be not one solid piece but composed of many separate particles of various sizes.

Comets are the most mysterious and capricious of solar system objects and

Comets are the most mysterious and capricious of solar system objects and the nature of the physical changes which they exhibit is still not fully understood.

Observers with quite modest equipment, say a refracting telescope of not less than 7.5 cm, or in the case of bright comets, a good pair of binoculars, can do useful work by following known comets and reporting on their appearance. The ability to make accurate brightness estimates is especially useful and well worth cultivating. Sweeping the sky for new comets, though requiring considerable patience and perseverance, is also within the scope of the equipment mentioned.

Many of the fainter comets are undoubtedly escaping detection, particularly in the Southern skies which are not being as intensively searched as the Northern. There is a need for more amateurs to undertake this work.

Interested persons are asked to contact the Director of the Comet and Meteor Section at 90 Malan Street, Riviera, Pretoria, 0084. TEL: 704-895

METEORS

Meteors or "shooting stars" result from small bodies entering the Earth's upper atmosphere, and are generally seen in greater abundance after midnight (due to the direction of the Earth's motion) than in the early evening.

There are two categories of meteors - the sporadic ones and the showers. A meteor shower comes from a certain direction in space (the Radiant) and is thought to be associated with the remains of a comet. When the Earth passes close to the comet's original orbit, such a shower can be expected. A list of these predicted showers is given in the table opposite.

The term "shower" is perhaps misleading (as the table indicates); the most prolific of these showers normally yields an average of less than one meteor per minute. On rare occasions however, as in the case of Leonids, there is a phenomenal rise in the number of meteors observed.

There is always the possibility of new showers occuring, and any large-scale meteor activity observed on dates other than those mentioned should be reported without delay.

Reports by a reliable observer of the number of meteors seen coming from a particular radiant in a given period are always useful, but the best work is done by organised teams making a full sky coverage. In the latter case, care must be taken to avoid the overlapping of individual reports, i.e. where more than one observer reports the same meteor(s), giving a false total for the group.

"Fireballs" are meteors of a luminosity equalling or exceeding that of the brightest planets. Accurate reports of their path among the stars, or their altitude and azimuth, at specific times, are of great value, particularly if made by observers at different places along the trajectory. Details of brightness (compared with Venus, Moon etc) size and form, colours and any train or wake, are also important.

Radiant

Shower

						05	R.A.	Dec.	Ü
Mar		-	14 - Mar 18	18	Corona Australids	16	16 ^h 20 ^m	-48°	
Apr	19	ī	- Apr 24	24	April Lyrids	18	88	+32	
May		Y	1 - May 12	12	Eta Aquarids	22	24	00	
Apr	20	1	- Jul	30	Sco-Sgr System	18	00	-30	
Jun	10	í	Jun	21	June Lyrids	18	32	+35	
Jun	17	-	Jun	92	Ophiuchids	17	20	-20	
Jun	56	í	Jun	59	Cetids (new)	02	00	-15	
Jul	0	- 1	Aug	5	Capricornids	2	00	-15	
Jul	15	-	Aug	15	Delta Aquarids	22	36	(-17	
Jul	5	-	Aug	20	Pisces Australids	22	40	-30	
Jul	15	1	Aug	52	Alpha Capricornids	20	36	-10	
Jul	Jul 15	1	Aug	24	Iota Aquarids	(22)	32	-15	
Oct	91	-	Oct	27	Orionids	90	24	+15	
Oct	0_	1	Dec	2	Taurids	(03	44	+14	
Nov	14 - Nov 20		101	50	Leonids	10	80	+22	
Dec	<u>د</u> ا	ä	Dec 5	5	Phoenicids	0	00	-55	
Dec	- 1	ă	Dec 15	15	Geminids	07	28	+32	
Dec	5	J	Jan 7		Velaids	09	99	-51	

* Uncertain

Comets and Meteors 1984

Conditions at Maximum	Unfavourable	Favourable	Favourable	Unfavourable	Unfavourable	Favourable	Favourable	Favourable		Favourable	Favourable	Favourable	Unfavourable	Favourable	Unfavourable	Favourable	Favourable	Favourable	Favourable
Recommended Time of	ı	20h - 24h	22h - dawn	1	1	19h - 23h	22h - dawn	20h - 02h		22h - dawn	22h - dawn	00h - dawn	ı	20h - 02h	1	20h - 24h	2h30- dawn	00h - dawn	02h - dawn
Transit of Radiant (xonqqs)	04h45m	04 15	07 30	00 30	00 10	23 30	07 40	00 20		07 70	02 10	00 00	(01 10	04 30	(00 50	06 30	20 10	05 00	03 00
an Maximum Ylnuch StsR	16 5	22 15	5 18	14 ?	16 8	20 8	28 ?	25 8	6	29 35	31 11	2 10	6 12	21 35	4 16	17 10	5.	14 55	2 62
Date	Mar	Apr	May	Jun	Jun	Jun	Jun	Jul		200	Jul	Aug	Aug	Oct	Nov	Nov	Dec	Dec 14	Dec 29

THE STARS

CONSTELLATIONS

Apart from our Sun all the stars that we see are so incredibly distant that, despite their high speed velocities, their apparent positions change by only minute amounts each year. Consequently the patterns that they form appear unchanged. The Greeks and other ancient civilisations identified these patters, or constellations, with various mythological characters and creatures, and most of the names they gave are still used today.

In all there are 88 constellations, roughly one half of which would be above the horizon at any one time. Some contain distinctive patterns of bright stars and are relatively easy to find; others are difficult to locate, even with suitable maps. The Southern Cross and Centaurus, Orion and Taurus, Scorpius and Sagittarius, are featured later in this section. Detailed information on other constellations is beyond the scope of this handbook and interested observers are advised to obtain a suitable star atlas.

STAR NAMES

Within each constellation, the brightest star is generally labelled a(Alpha), the next B(Beta) and so on through the Greek alphabet. Most of the brightest stars also have their own names – usually of arabic origin. For example, α Canis Majoris, otherwise known as Sirius, is the brightest star in the constellation Canis Major.

STELLAR MAGNITUDES

The apparent brightness of a star - which depends both on its true luminosity and its distance - is indicated by its magnitude. Equal intervals of magnitude represent equal ratios in light intensity. A star of magnitude 1,0 (typical of the brightest stars in the night sky) would be exactly one hundred times more luminous than a star of magnitude 6,0 (about the limit of visibility to the naked eye). The maps in this section show stars down to magnitude 4,5.

STELLAR DISTANCES

Distances are often expressed in units of light years - the distance light would travel in a year (equal to $9.5 \times 10^{12} km$).

DOUBLE STARS

It now appears that single stars such as our Sun are the exception, the majority of stars being double or multiple - two or more suns in orbit around one another.

STAR CLUSTERS

These are of two completely different sorts. Galactic clusters, having of the order of 100 stars, are found close to the plane of the Milky Way. The ones we can see are relatively nearby. Globular clusters are much larger and far more distant. They contain of the order of 100 000 stars each and are seen above and below the Milky Way on that side of the sky towards the centre of our galaxy. So great is their distance that small telescopes fail to resolve individual stars - instead they appear as fuzzy balls.

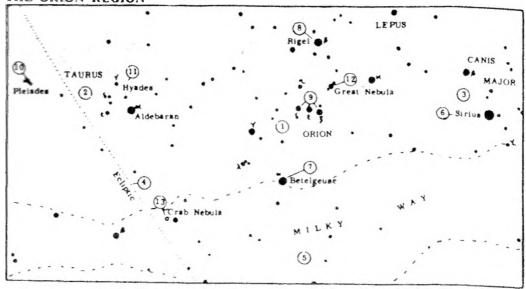
NEBULAE

Possibly one third of the matter in our region of the Galaxy is in the form of gas and dust (the remainder being contained in stars, plus a tiny amount in planets). Condensations of this material are called nebulae, some of which are illuminated by nearby stars while others are dark. They are usually referred to by their numbers in Messier's catalogue (M) or the New General Catalogue (NGC).

THREE POPULAR REGIONS

The dominating constellation of the summer skies is Orion, that of the winter skies is Scorpius, while the Southern Cross is conspicuous for most of the year. The regions around these constellations are also rich in interesting objects - visible either to the naked eye, or with the aid of binoculars or a small telescope - and are featured in the maps and text below. It may be necessary to rotate the maps to match the orientation of the constellations in the sky.

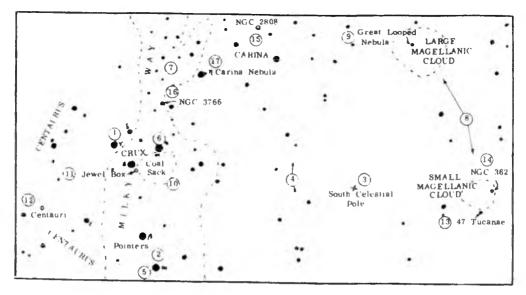
THE ORION REGION



- 1) The constellation of Orion. The figure of the legendary hunter of Greek mythology is unfortunately upside down when seen from Southern Africa. The faint stars by λ represent the head, α and γ the shoulders, $\delta \in \zeta$ the belt, and β and ζ the legs. Orion forms part of the "great hunting scene" in which he faces the poslaught of 2) Taurus, the bull. Only the forepart of the bull is depicted and, like Orion, it is upside down. Of and ξ are the eyes. γ the nose. Orion is accompanied by 3) Canis major, the large \log and the small dog (off map) while Lepus, the hare, crouches at his feet.
- 4) A section of the Ecliptic a line encircling the entire sky and representing the plane of the Earth's orbit. As the Earth revolves around the Sun, the Sun appears to move along the ecliptic through the constellations of the Zodiac, of which Taurus is one.
- A portion of the Milky Way (looking out towards the edge of our Galaxy).
- 6) Sirius the brightest star in the night sky. It is somewhat brighter than our Sun and relatively close by at a distance of 9 light years. It is a double star but the companion is a white dwarf (only slightly larger than the Earth, and with a mass comparable to our Sun) and is only visible through a large telescope.
- 7) Betelgeuse most famous of the red giant stars. Its diameter is of the order of the size of the Earth's orbit and its luminosity is nearly 10 000 times that of our Sun. Its red colour should be obvious to the eye. It is 520 light years distant.
- Rigel, despite being physically smaller than Betelgeuse, is more luminous (higher surface temperature - bluish colour) and more distant.
- 9) The stars in Orion's belt are distant hot blue stars.
- 10) The Pleiades or Seven Sisters form the best known nearby star cluster. Six or seven stars are visible to the naked eye, binoculars or a small telescope show more,
- 11) The Hyades is another nearby galactic cluster, but Aldebaran is not a member (it lies closer to us).
- 12) The Great Nebula in Orion, just visible to the naked eye, shows up as a fan shaped mass of luminous gas through binoculars or a telescope. A telescope will also show a tiny "Trapezium" of four stars in the centre.
- 13) The Crab Nebula, the remnant of a supernova recorded by the Chinese in 1054, requires a moderate sized telescope for observation. In its heart is located the extraodinary pulsar which emits a double flash of light 30 times every second. The current belief is that it is a rapidly rotating neutron star a star with the mass of our sun but with a diameter of only 10 km.

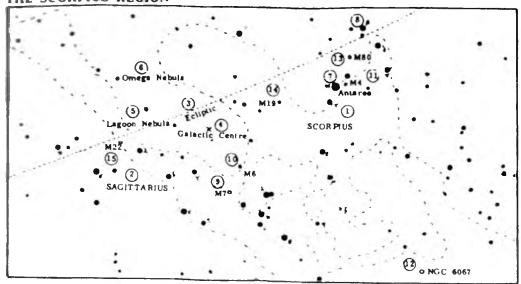
The Stars

THE SOUTHERN CROSS REGION



- 1) Crux, the Southern Cross, is one of the most compact patterns of bright stars to be found in the sky. It lies on the border of that region of the sky which never sets as seen from Southern Africa.
- 2) The two "Pointer" stars lie close to the Cross. (A similar pattern to the Southern Cross - called the False Cross - lies just outside and above the map, but has no accompanying pointer stars).
- 3) The South Celestial Pole: This is one of two opposite points in space towards which the Earth's axis of rotation is directed. As the Earth rotates so the sky appears to pivot about this point. It always lies above the south point on the horizon, elevated by an angle equal to the observer's southern latitude. (The north celestial pole lies below the northern horizon and can never be seen from the Earth's southern hemisphere).
- 4) The intersection of a line extended through the major axis of the Cross and the perpendicular bisector to the Pointers indicates the approximate position of the South Celestial Pole.
- 5) Ω Centauri has the distinction of being 2 the closest star to our solar system at a distance of approximately 40 \times 10 km or 4,3 light years. A small telescope readily shows that it is a double star the two components take 80 years to resolve about one another. A much fainter third star also belongs to the system.
- Crucis can also be resolved as a double star by a small telescope (separation 5 sec of arc).
- 7) The region indicated is one of the brightest sections of the entire Milky Way.
- 8) The Large and Small Magellanic Clouds are the nearest of the external galaxies (see also next section). They can be seen with the naked eye provided the sky is reasonably dark.
- 9) The Great Looped Mebula possibly the remnant of a supernova explosion in the Large Magellanic Cloud. (Naked eye or binoculars).
- 10) The "Coal Sack" a dark mass of gas and dust obscuring a part of the Milky Way. (Naked eye or binoculars).
- ll) Herschel's "Jewel Box" a galactic cluster containing stars of different colours. (Small telescope or binoculars).
- 12) ω .Centauri and 13) 47 Tucanae are perhaps the best known globular clusters. Binoculars will show their fuzzy appearance. 14) NGC 362 and 15) NGC 2808 are fainter globular clusters.
- 16) NGC 3760 a fine galactic cluster. (Binoculars or small telescope).
- 17) The $\,$ $\,$ $\,$ Carinae nobula a site of a slow supernova that brightened to magnitude 0.8 in 1843 and is now of magnitude 6.4.

THE SCORPIUS REGION



- 1) The constallation of Scorpius. The creature is depicted with α in the centre of the body and β and π the claws. The distinctive tail $\epsilon=\zeta=\rho$ curls round to the stin. λ
- 2) Sagittarius the figure of the centaur archer is very difficult to make out.
- 3) A section of the Ecliptic. Like Taurus, Scorpius and Sagittarius are constellations of the Zodiac,
- 4) The direction of the centre of our Galaxy the Milky Way is that part of our Galaxy visible to us. Unfortunately the central nucleus is obscured by foreground gaseous and dusty matter both dark and luminous neme the irregular shape of the Milky Way in this region. Luminous nebulae include 5) the Lagoon nebula and 6) the Omega nebula. These are best seen with the aid of binoculars.
- 7). Antares a distant rod giant, several hundred times the diameter of our Sunis so named because its red colour givals that of the planet Mars.
- 8) β . Scorpin can be resolven as a double star (separation 16 sec of arc) with a small telescope. In fact the brighter component is in itself a triple star, and the fainter component a double of α .

This region includes a number of galactic clusters including 9) M7, 10) M8, 11) M4 and 12) NGC 6067, (use binoculars or a small telescope).

Further from the plane of the Milky way are some globular clusters: 13) M80, 14) M19 and 15) M22,

NOVA SEARCHING

On rare occasions a star may undergo a nova outburst, its light increasing tremendously. The result is that a "new" star appears where previously no star was visible to the naked eye, or even with a small telescope. The light of the nova may fluctuate for a time, then gradually fades over a period of days, weeks or months.

Even observers having no telescopes can perform a useful task in keeping a watch for such novae in an allocated area of the sky. A good knowledge of the constellations is a recommendation, since part of the procedure is to scan the whole sky for bright novae before the more detailed search in the allocated area is begun. However, anyone can be given training in star recognition.

Interested persons should contact the Director of the Nova Search Section, Mr J C Bennett, 90 Malan Street, Riviera, Pretoria, 0064.

VARIABLE STAR OBSERVING

The "General Catalogue of Variable Stars" by Kukarkin and Parenago lists some 20 000 stars. Professional observatories cannot possibly monitor all of these, and this makes the observation of variable stars a field in which amateurs can make a real contribution to astronomical knowledge.

Of the 20 000 stars, at least 2000 are suitable for visual monitoring in the southern hemisphere. However, the number of active observers in this part of the world remains woefully small, and scarcely 200 variables are at present being observed from South Africa.

The Variable Star Section of the A.S.S.A. exists for the purpose of encouraging observers and of acting as a medium of communication. The Section disseminates incoming information amongst observers, and will forward (on request) the observations of individuals to various variable star bodies. These include the American Association of the Variable Star Observers (AAVSO) and the Variable Star Section of the Royal Astronomical Society of New Zealand. These bodies combine the South African light estimates with those from other parts of the world. The resulting "light curves" and tables are made available to a large number of professional observatories where astronomers are interested in investigating certain of the stars more fully.

Visual estimates of magnitude are made by comparing the variable with two or more comparison stars, respectively brighter and fainter than the unknown variable. Suitable comparison stars are shown on special charts, which have been prepared for each variable, mainly by the two variable star organisations mentioned above. The use of these charts is essential for accurate, standardized observations, and intending new observers are therefore advised to obtain the necessary data by contacting the Director of the Variable Star Section, Mr J Hers, P O Box 48, Sedgefield 6573, Telephone (04455) 736. They will then be sent charts of a few easy objects, and data on stars which may be observed with the equipment at their disposal.

Prospective observers should, when writing, give brief details of their equipment. Larger, more powerful telescopes will naturally greatly increase the number of stars which may be measured, but many variables are bright enough to be observed through most of their cycles with quite modest equipment, e.g. binoculars. Some stars, such as 07104 L2 Puppis, are so bright that they may be observed without optical aid whatever.

Variable stars are designated in two ways. The first of these, the Harvard designation, consists of six figures which give the position for 1900; the first four figures give hours and minutes of R.A., the last two give the declination in degrees, underlined for southern declinations. The second name consists of one or two letters (or letter V plus a number) and the name of the constellation.

Variables can be divided into three main classes: pulsating, eruptive, and eclipsing

binary stars.

Most suitable for beginners are the <u>long period variables</u> (or Mira variables, named after the typical representative Mira = $\frac{1}{2}$ 0 Leti) which belongs to the class of pulsating stars. They are giant stars which vary through a range of brightness of 2,5 to 5 magnitudes or more, and which have well-defined periodicities, ranging from 80 to 1000 days. In most cases one observation per observer every 10 days will suffice.

Typical examples include:

Among the eruptive variables, two groups are of special importance:
U Geminorum type. These are dwarf novae which have long periods of apparent quiescence at minimum, with sudden rises to maximum. A typical representative in the southern hemisphere is 040971 VW Hydri.

R Coronae Borealis type. These are high luminosity variables with slow, non-periodic drops in brightness. A typical representative is 191033 RY Sagittarii.

Eclipsing Binary Stars have orbital planes which lie close to the line of sight of the observer. The two components periodically eclipse each other, thus causing variations in the apparent brightness of the system. Periods are generally short, of the order of hours, so that observational programmes need very careful planning. Monitoring these interesting stars is therefore for experienced observers only.

MINOR PLANET OCCULTATIONS:

A number of A.S.S.A. members and professional observatories form part of a world wide network which observes the above events. The observations are very useful to astronomers who study the Solar System.

Often an amateur is located on or near an occultation path, and the observation which he or she can make may be of considerable value. The equipment requirements are modest. A 50 mm telescope and means to record the times of multiple events will suffice in most instances. The timing equipment can comprise a portable cape recorder and a radio tuned to a continuous time signal such as ZUO or WWV. If a continuous time signal cannot be received reliably, than an assistant can read off time intervals of say ten seconds from a quartz watch synchronised with the S.A.B.C. "six pips" time signal. The commentary of the observer and timekeeper is thus recorded for later analysis.

Stars occulted by minor planets are not always easy to identify, but occultations notices contain hints on how to find the stars with a minimum of fuss.

If you are in touch with one of the A.S.S.A. Centres and would like to participate then you are invited to contact one of the conveners listed below. If you do not live near a Centre then please contact M.D. Overbeek, P. O. Box 212, EDENVALE. 1610

OCCULTATIONS BY MINOR PLANETS

The following predictions for occultations by Minor planets indicate possible occurrences during 1984. These include the occultations which may be observable from the African continent and details near the time for local possibilities can be obtained from Mr Overbeek. Times are in S A S T. Star numbers are from AGK3 or SAO catalogues.

			Deauleina	Rody		Star		Mag change at	Max
_			Occulting						
D	ATE		Name	Mag		Nο	Mag	Occultation	Duration
Jan	8.02	4	Vesta	7.0	+	19°0410	9.3	0.1	(seconds)
	29.27	46	Hestia	13.1	+	3°1471	8.6	4.5	16
Feb	6.94		Dembowska	10.9	+	15°1128	9.6	1.6	13
	27.85	566	Stereoskopi	a 13.6	+	24°0996	11.4	2.3	14
May	2.15		Nemisis	12.4	-	2°0809	7.0	5.4	10
Jul	14.13	139	Juewa	12.4		209985	8.3	4.1	15
Jul	16.96	211	Isolda	12.7		164173	9.1	3.6	15
Aug	8.82	87	Sylvia	11.9		211985	10.0	2.1	11
Sep	17.86		Flora	8.6	+	0°0197	9.4	0.4	40
Dec	4.01	40	Harmonia	10.3	+	22°0798	8.7	1.8	14
	26.08	747	Winchester	10.4	+	4°0801	9.3	1.4	19
	30.88	111	Ate	11.0	+	26°0711	9.5	1.7	14

CAPE TOWN Cliff Turk, 20 Nerina Avenue, PINELANDS, 7405

PIETERMARITZBURG C.S. Lake, 23 Munroe Ave., Northern Park, PMBG, 3201

BULAWAYO Arthur G.F. Morrisby, Dept. of Surveyor General. P. O. Box 1580, Bulawayo, ZIMBABWE.

BLOEMFONTEIN G.N. Walker, 39 Vilonel St., Dan Pienaar, BLOEMFONTEIN. 9301

WITWATERSRAND M.D. Overbeek, P. O. Box 212, EDENVALE, 1610

HARARE R.W. Fleet, P.O. Box 1435, Harare, ZIMBABWE.

DURBAN R.K. Field, 303 Wakesleigh Road, BELLAJR. 4094

PRETORIA J.C. Bennett, 90 Malan Street, Riveria, PRETORIA. 0084

PORT ELIZABETH Mr D Jesson, Busaf, P.O. Box 4008, KORSTEN, 6014

ORDINARY OCCULTATIONS

This Section and that following concern a specialised branch of observational astronomy in which both professional and amateur participate. The tables of predictions must necessarily occupy a number of pages as this handbook is the sole published source for Southern Africa.

An occultation occurs when the disk of the Moon moves in front of a star. Timings of occultations, to a precision of one-tenth of a second if possible are very valuable for studies of the Moon's shape and motion. Since only very modest equipment is required, amateurs can make important contributions in this field. Persons interested in making and reporting occultation observations are urged to contact the Director of the Society's Occultation Section, Mr A GF Morrisby, (c/o Dept of Surveyor General, P O Box 1580, Bulawayo, Zimbabwe).

Predictions of occultations of all stars brighter than magnitude 7,5 supplied by Hydrographic Dept, Tokyo are given below. The main set of tables give predictions for three stations, namely,

	Longitude	Latitude
Cape Town	- 18°. 475	- 33°. 933
Johannesburg	- 28°. 075	- 26°. 182

This does not restrict its use to observers to those centres. The approximate time of an occultation at a place λ degrees west and ϕ degrees north of one of the standard stations given above may be found from

Approximate time = predicted time + $\partial \cdot \Delta \lambda$ + b. $\Delta \phi$

where a and b, in minutes of time, are given in the tables. Alternatively, rough times for intermediate stations can usually be estimated direct from the tables.

Occulted stars have been identified by their Z.C. numbers, that is their numbers in the "Catalogue of 3539 Zodiacal Stars for the Equinox 1950.0" by James Robertson (U.S. Naval Observatory, 1939).

Note: That the times of these occultations are given in U.T.

Explanations of Abbreviations used in Tables

the number of the star in the Zodiacal Catalogue. An "m" following the number indicates the star is not single.

Mag. the visual magnitude

the Phase: D = Disappearance, R = Reappearance the time of the occultation in U T

h.m.

parameters in minutes for predicting times other than at standard a,b

stations (explained above in text)

The Position Angle on the Moon's limb measured eastward from the P.A. north point

EAPE TOWN E1845 5 3349 Sate 2.0 Jac b All o 9 19 41.6 1.0 21 32.9 0128 7.3 1.2 19 12.3 0355 7.5 166 1.6 19 57.4 1251 25 08 43.4 2023 2.2 60 31.8 2259 2542 29 02 19.6 7 7 Feb 10 1.1 1.2 1.0 2 0 2.6 2.1 72 22 22 22 2.3 01 35.3 24 2376 02 00.9 2376 7.6 26 01 54.0 2659 Mar 13 21 26.2 1270 2.1 21 25.4 2302

2 c Hag Egl PA Nο #ag Elg PA 0020 6.8 72 31 -0.5 2.5 2334 5.6 173 77 -0.8 1.6 2854 4.7 153 279 0.1-1.2 2879 6.6 152 300 -0.8-2.3 5.9 174 238 -0.9-0.1 3164 4.7 127 210 -1.4 3.0 7.9 92 290 -0.8-1.6 22 02 38.9 6.7 67 295 0.1-1.5 23 03 02.9 0083 6.9 83 239 -1.4 0.3 0202 7.0 71 193 -0.3 2.2 -0.2-1.0 43 265 Lind 3 17 07.1 1610 8.7 61 73 -2.7 1.9
20 51.7 0568 8.1 98 94 -1.2 1.0 5 18 37.9 1648 8.0 88 164 -0.6-2.2
19 18.1 0709 4.3 110 76 -2.2 0.9 6 22 40.6 1985 7.1 103 98 -0.4 0.8
20 42.4 0867 6.9 123 65 -2.1 1.3 8 17 55.9 2211 7.2 127 97 .2.2-0.8
21 17.3 1739 6.5 151 339 -0.7-1.8 10 00 16.1 Uranus 5.8 142 47 -0.3 3.3
21 24.4 1856 6.6 137 262 -2.9 0.3 10 19 29.4 2510 6.3 153 65 -2.8 0.8
23 42.0 1976 6.9 124 264 -1.8-0.8 11 17 07.2 2652 6.4 164 143 0.2-2.6
00 16.9 1978 6.6 124 268 -2.0-0.8 11 17 07.2 2652 6.4 164 143 0.2-2.6
00 16.9 1978 6.6 124 268 -2.0-0.8 11 17 07.2 2652 6.4 164 143 0.2-2.6
00 20.2 2097 7.1 111 333 -0.4-2.5 15 00 44.9 3106 5.4 157 264 -2.6 0.4
10 08.4 Saturn 0.4 107 103 0.1 0.6 15 02 51.0 3116 6.7 157 257 -1.7 1.4
22 11.1 2213 5.9 99 300 0.2-1.5 17 02 22.4 3358 7.2 135 240 -2.1 1.5
23 14.8 2218 5.6 99 302 -0.1-1.7 20 03 04.0 0157 8.1 101 238 -2.0 0.9
02 44.9 2232 7.2 97 294 -1.7-1.5 22 03 05.5 0380 7.4 79 324 Graze 3 17 07.1 1610 8.7 61 73 -2.7 1.9 4.6 85 184 1.8-5.4 4.6 85 226 -3.8 2.2 6.4 62 265 -0.6-1.0 Aug 3 19 04.0 2060 6.3 86 127 -1.4-0.6 3 19 43-0 2064 6.5 86 87 -1.4 1.2 5 20 35.9 2327 6.7 112 172 -1.8-5.7 5 20 29.4 2330 6.3 112 88 -1.8 1.1 5 21 57.2 2337 6.4 112 99 -1.1 0.8 7 00 18.2 2490 5.4 126 118 -0.6 0.3 2.1 22 10.1 2302 2490 5.4 126 118 -0.6 0.3 22 10.6 2.1 2303 7 00 19.6 2491 6.7 126 69 -0.0 1.7 22 56.2 22 2450 7 20 43.9 2622 6.3 136 52 -2.2 2.7 7 22 04.5 2627 6.9 137 121 -2.3-0.7 9 02 00.8 2809 4.9 150 74 -0.1 1.5 2622 2 4 03 27.2 92 206 -4.5 8.7 26 03 51.5 2622 6.3 4.9 150 74 -0.1 1.5 12 19 22.2 3304 6.4 166 260 -0.8-0.8 12 19 22.2 3304 6.4 166 260 -0.8-0.8
0607 8.8 46 96 -1.2 1.0
0900 4.9 71 140 -1.9-1.3
16 27 48.6 0219 5.1 121 228 -0.7 0.4
1206 5.9 97 93 -2.3-0.1
17 03 52.4 0235 6.9 120 330
1211 6.2 97 150 -1.2-1.5
1349 8.1 110 70 -2.8 0.7
2114 5.8 163 248 -2.5 2.3
2376 4.6 140 273 0.1-1.0
2703 7.4 113 245 -1.5-0.1
2864 4.7 100 317 -0.7-3.5
3000 8.1 88 234 -2.3 1.3
3116 6.7 79 295 0.3-1.9
3116 6.7 79 295 0.3-1.9
3243 7.4 67 303 0.2-2.7 Apr 5 17 25.4 7 16 55.4 9 17 35.1 q 18 27.2 18 37.2 1.0 17 01 45.2 19 16.3 1.8 20 23 34.3 01 13.8 22 2.3 03 44.7 2.1 23 54.8 25 00 56.9 3243 7.4 67 303 0.2-2.7 18 02 36.8 0809 8.0 93 232 -1.4 0.4 7.6 64 95 -0.1 0.9 20 36.1 Hay 29 2 3 8 8 6 q 2852 7.4 100 109 -2,4-0.1 10 3116 6.7 124 350 2.0 7.1 1.4 3265 6.6 137 109 -0.6 0.7 15 3478 6.5 156 30 -1.1 1.7 15 0258 6.6 170 274 -2.2-1.4 15 0621 6.2 136 237 -1.6 0.4 15 0625 7.0 135 202 -1.0 1.8 1.6 1075 8.0 100 227 -1.0 0.3 1.7 2804 5.9 70 80 -0.4 1.3 6.7 156 110 6.3 145 237 -1.6 U.. 6.4 143 284 -2.2 0.0 7 110 228 -0.4 0.2 2 1 1.5 17 2963 8.1 82 72 -0.3 1.5 21 06.6 1.7 2622 18 03 06.2 2652 Hov 21 33.2 7.8 115 114 -4.6-2.2 2.0 3062 7.5 110 228 -0.4 0.2 2 18 45.3 3323 03 59.5 13 01 32.7 22 3214 1042 6.6 130 200 -1.9 3.5 00 25.5 16 01 47.9 18 02 16.7 1432 7.0 9.3 296 2.3 3323 7.8 86 260 -0.6-0.9 66 336 -0.7-2.5 1672 8.4 Jen 27 -0.3 2.7 29 19 49.6 3284 7.1 8.4 16 48.3 3 1270 6.1 50 134 -0.9-0.4 30 22 28.5 1609 7.0 9.6 33 0.0 2.2
 18
 11.0
 1274
 5.7
 51
 92
 -1.0
 1.0

 17
 31.9
 1401
 7.9
 64
 65
 -2.8
 2.1

 19
 00.3
 1533
 7.2
 78
 103
 -1.4
 0.5

 22
 31.6
 1773
 5.1
 106
 78
 -0.8
 1.8

 22
 56.3
 2016
 6.5
 133
 113
 -1.2
 0.1

 00
 18.3
 2022
 5.5
 133
 62
 -0.6
 2.9

 00
 02.7
 2275
 5.9
 159
 45
 -1.2
 1.2
 1.2
 3 Dec 4 1 21 34.4 3521 8.0 107 35 -0.6 2.4 ٩ 6 20 28.2 0301 6.8 139 355 0.3 3.7 7 5 23 39.9 7 00 57.8 14 00 25.3 g 3423 5.4 152 6 -1.2 3.7 6.0 163 0534 15 -1.8 3.7 1.0 12 1517 7.7 111 309 -1.4-1.9 77 20 27.7 1156 5.9 64 41 0.0 2.1 1.2 16 24.1 Uranus 5.8 169 149 0.5-2.1 1.3 02 11.2 2434 5.6 173 77 -0.8 1.6

			LUN	AR OC	CULTA	TIONS 1984	1		DAT		z c	MAG	РН	ELG	TIME	A B	р
				JOH	ANNES	BURG			JOF	15	3089	5.3	R D	159 157	19 51.3	-1.0 -1.9	
DATE	1 6	2 €	MAG	PH	ELG	TIME	A 8	Ρ		15	3116	6.7	R D	157	03 15.7	-1.3 1.3	
JAN	21	1651	6.8	D D R D	72 133	20 04.2	-0.1 2.7	22 337		15	3214	6.6	R D	148	18 49.0	-0.6 0.6	
	27	2259	6.7	R D	67	00 17.2	0.2 -2.0	322		17	3358	7.2	R D	135	02 54.8	-1.9 1.8	
FEB	11	0709	4.3	D D	110	19 48.8	-2.1 i.6	62		29 29	1435	6.6	ם ם ם ם	16	15 34.0	0.5 -1.8	
	12	0867	6.9	D D	123	21 19.6	-2.6 3.2	15	AUG	2	1933	7.0	D D	72	18 47.6	-0.3 -4.5	
	13	1046	6.9	םם	137	23 59.1	-0.5 -0.1	120		2	1941	4.8	D D	73	20 08.9	-0.4 -1.9	160
	13	1484	6.6	ם ס	137	23 54.1	-0.0 -0.5	133		2	1941	4.8	R 8	73	20 48.9	0.1 1.9	248
	20	1856	6.6	R D	137	01 42.9	-1.8 -1.2	305		3	2060	6.3	מם	86	19 18.1	-1.3 0.8	
	20	1976	6.9	R D	124	23 48.8	-1.6 -1.6	299		5	2307	6.5	D D	86 110	15 39.1	-0.5 8.1	
	21	1978	6.6	R D	124	00 24.6	-1.6 -1.7	305		5	2310	4.6	ם ם	110	16 16.8	-1.5 -2.5	
	22	2213	5.9	R D	99	21 55.6 22 59.8	0.1 -2.4	325 332		5	2337	6.4	D D	112	22 12.9	-0.5 1.3	80
	23	2232	7.2	R D	99	02 42.1	-1.0 -3.1	335		7	2490	5.4	D D	126	00 24.7	-0.2 0.4	
	24	2376	4.6	D D	85	02 23.6	-2.1 -1.1	279		7	2491 2622	6.7	D D	126	00 30.8	0.5 1.7	57
	25	2507	6.7	ם ם	74	00 02.6	2.1 -5.0	100		7	2627	6.9	ם ם	137	22 22.8	-1.8 0.3	
	25	2507	6.7	R D	74	00 23.0	-2.7 2.5	219		8	2652	6.4	0 0	138	02 01.0	-0.1 0.0	
	26	2659 2672	2.9	R D	62	01 47.9	-0.4 -1.8	301 251		9	2809	4.9	D D	150	02 10.9	0.3 1.3	68
MAR	6	0283	7.0	D D	43	17 52.6	-0.6 0.0	103		1.1	3175	4.8	N D	175	19 32.9	-1.9 -0.6	
	9	0634	5.3	ם ם	77	16 53.0	-2.6 0.3	74		12	3304	6.4 5.2	R D	166	19 20.7	-1.3 -2.5	
	21	2 3 0 2	2.9	DB	1 [9	21 08.3	0.0 -1.6	124		14	3446	1.2	8 D	153	02 34.0	-3.3 -0.2	
	21	2 30 2	2.9	R D	119	22 09.2	-0.7 -1.4	289		16	0219	5.1	R B	121	23 01.4	-2.5 4.0	
APR	16	2330	6.3	R D	116	04 01.2	~3.0 4.7 -0.6 -0.9	228		17	0235	6.9	D D	120	04 21.6		334
67.6		Saturn		D B	151	23 16.8	-1.5 -6.3	188		21	0716	6.2	R D	75	00 50.0	-0.8 -0.5	
	17	Saturn	-1.1	R D	151	23 16.8	-5.0 3.6	233	SEPT	31	4006 2267	0.6	RB	63	07 41.0	-1.0 4.5	
	19	2523	4.9	R D	127	19 27.0	1.2 -2.8	344	3671	2	2430	7.0	DD	94	20 09.0	-0.4 3.3	
	20	2703	7.4	R b	113	23 41.8	-1.4 -1.4	284		16	0660	4.4	R D	106	23 38.6	-0.9 0.3	
MAY	28	1169	5.4	RD	34 67	02 43.2	-0.3 -1.4	278 98		1.7	0664	5.4	R D	106	00 22.6	-1.5 0.1	247
MAL	10	1689	5.5	D D	121	20 42.0	-1.4 -0.6	122		17	0676	7.1	R D	105	02 36.5	-4.5 -2.4	
	1.2	1923	7.1	D D	148	20 10.3	-3,3 0.4	84		18	0809 1263	8.0 7.1	R D	93 56	02 58.6	-2.2 0.9	
	13	2032	7.3	D D	161	16 54.1	0.2 -3.0	172		30	2500	3.4	D D	74	12 02.0	-1.2 -1.7	
	14	2060	6.3	D D	164	00 15.5	-1.0 -2.5	161		30	2500	3.5	R B	74	13 23.1	-2.1 -1.1	285
	15	2302	2.9	D 8	172	17 08.3	-0.3 -0.8	88	OCT	1	2676	6.5	ם ם	88	17 12.2	-3.2 -2.7	
	15	2 30 2	2.9	R D	172	17 59.5	-0.1 -2.1	323		2	2852	7.4	D D	100	20 02.0	-1.8 0.5	
0.0	15	2303	5.1	R D	172	17 59.3	0.1 -2.1	324		6	3116	6.7	0 0	124	21 27.3	3.3 7.4	
	15	2330	6.3	R D	170	23 51.9	-2.5 -0.2	281		7	3478	6.5	0 0	156	18 51.8	-0.9 3.2	
	16	2337	5.4	R D	169	01 28.4	-1.8 0.4 -1.8 0.1	278		1.4	0621	6.2	R D	136	00 22.4	-2.3 0.9	238
	17	2491	6.7	N D	156	02 08.2	-1.a U.I	355		14	0625	7.0	R D	135	01 20.8	-1.7 2.6	
	17	2622	6.3	R D	145	21 15.5	-1.3 1.2	279		17	1067 2750	7.2	R D	100	00 11.7	-1.8 -1.3	
	17	2627	6.9	RD	145	22 21.9	-2.8 1.0	241		29	2750	2.1	D D R B	66 66	10 01.1	-0.9 -1.9	
	18	2672	2.9	D B	141	06 57.2	0.3 1.3	66		29	2804	5.9	D D	70	20 25.2	0.0 1.2	
	19	1202	5.9	R D	132	00 49.3	-2.8 1.8 -1.7 3.0	236		31	3089	5.3	D D	94	23 13.2	0.2 0.7	96
	22		6.6	R D	96	04 33.0	-2.4 1.6	242	NOV	6	0219	5.1	D D	158	16 40.6	-1.8 -2.3	
JUN.	2		7.2	מפ	37		-3.2 5.5	34		10	0714	6.2	R D	156	18 46.9	-1,2 -1.9	
		1263		ם מ		16 08.4	-6.0 6.6	38		13	1042	6.6	R D	130		-3.1 1.9	
	3			яв			3.2 -6.7	9		16	1431		R D			-1.9 -2.0	
		1270	7.4		50 64	16 59.6	0.2 -1.1	96		21	2033	4.3	A D	25	02 43.4	0.0 -1.6	310
	9		6.5				-0.7 -2.6	149		24	2500		D D			-0.9 -1.2	
	9		6.5	D D	133		-1.0 1.5	79		24	4001	-0.1	D D	22	15 12.1	Me	reury
	12		5.8	ם ם	169	16 07.9	-0.0 -1.4	118		29	3284		D D	84		0.0 2.5	26
		2434	5.6	D D	173		-0.1 2.0	60	DEC	30	3409 0301		D D	96	22 44.1	-0.2 4.1	
-	15	2864	4.7	R D R D	153	18 24.8	-1.3 2.7	317 216	DEC	5	0413		םם	150		-1.8 1.3	
	23	0202	7.0	R D	71	03 24.9	-0.8 2.2	202		7	0664		D D	173		-2.2 -2.2	
		0709	4.3	D B	24		-1.1 -1.5	104		18	1985		R D	58	00 52.0	0.0 -1.8	316
	21	0709	4.3	R D	24	04 54.2	0.1 1.4			19	2114		R D		02 59.3	0.5 -3.1	
JUL	7	1006	0.4	D D	114	16 49.0	-1.2 -2.5	150 urn		23	2750		D D			-0.3 -1.0	
6	7	4006	0.4	A B	114	L8 Q6.7	-2.6 -0.5			23	2750 3228		R B D D			-0.8 -1.2	
							Sat	urn		27	3356		0 D	64	20 41.5	0.4 2.0	
92		2490	5.4	D D	152	16 05.3	1.6 -4.8			29	0036		D D	85		-1.1 2.5	
4	11	2652	6.4		164		-0.3 -1.6			31	0257	4.5	D D	107	20 06.9	-G.7 3.8	6
2	12		4.9				-1.6 1.7	37	37								
0.0																	

GRAZING OCCULTATIONS

When a star moves tangentially to the limb of the Moon, and is occulted for a very short period only - a few minutes, or even seconds - a grazing occultation is said to occur. Because the limb, as seen from the Earth, is in fact the outline of numerous mountains and valleys, there may be several disappearances and reappearances, which are not only fascinating, to observe. but which may be accurately timed to yield valuable data on the relative positions of star and Moon, in both right ascension and declination, as well as on the shape of the Moon. Some of these data cannot readily be obtained in any other way.

The maps on the following pages have been prepared by Hydrographic Dept, Tokyo to show the tracks of stars brighter than 7,5 magnitude which will graze the limb of the Moon when it is at a favourable elongation from the Sun and at least 10° above the observer's horizon (2° in the case of bright stars). Each track starts in the West at some arbitrary time given in the key and ends beyond the area of interest, except where the letters "A", "B" or "S" are given. "A" denotes that the Moon is at a low altitude, "B" that the bright limb interferes, and "S" that sunlight interferes. The tick marks along the tracks denote 5 minute intervals (at whole minutes), give the approximate time of the graze at places along the tracks.

The tracks as shown on the maps are approximate only. Since the observer's location is very critical, successful observations call for very accurate predictions. Such predictions, which include graphical representations of the probable profile of the Moon, are computed annually for a number of centres in Southern Africa. By plotting the predicted graze track on a reliable survey map (e.g. South African 1:50 000 series) it is usually possible to select a convenient site from where the graze may be observed. Ideally a team of observers would be stationed at intervals along a line running at right angles to the graze track - say, along a main road - each with his own telescope and timing equipment. Each observer will see a different sequence of events, the combined results forming an accurate picture of the limb of the Moon.

The equipment needed is similar to that used for ordinary (or "total") occultations, but must, of course, be portable. A 75 mm refractor is ideal for average events, but better instruments with a larger aperture have often shown their superiority under difficult conditions. Timing is best carried out with a portable tape recorder and radio receiver tuned to ZUO or other time signal station.

It will be seen from the maps that many grazing occultations occur in regions which are rather far removed from the main cities, and which cannot easily be reached by teams of observers from one of the ASSA centres. It is worth remembering, however, that a team of many observers, while ideal, is by no means essential; that a single good observer is worth more than many unsuccessful ones, and that one good observation is worth infinitely more than no observation at all

Interested observers - especially those living in the more distant regions who wish to be informed of favourable grazes occurring in their neighbourhood. are therefore invited to contact the co-ordinator for grazing occultations.

M.D. Overbeek, P.O. Box 212, EDENVALE, 1610 TEL: (011) 535442 EXPLANATION OF COLUMN HEADINGS IN TABLES:

Nο. the number of the track on the map. An asterisk denotes that the same is double - notes are given below.

Z.C. the number of the star in the Zodiacal Catalogue.

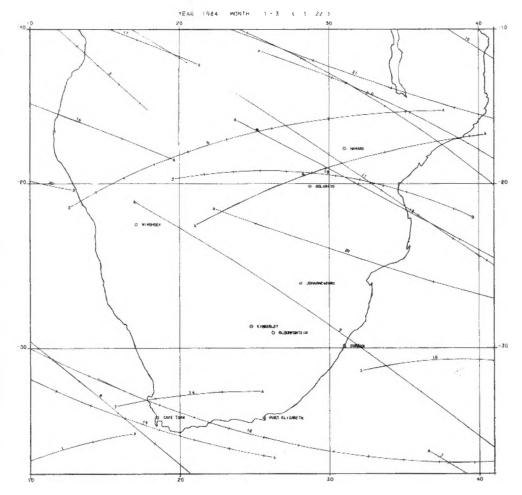
Date Beginning -

an arbitrary time (U T) of the beginning of the track in the west.

the percentage of the Moon's disk lit by the Sun.

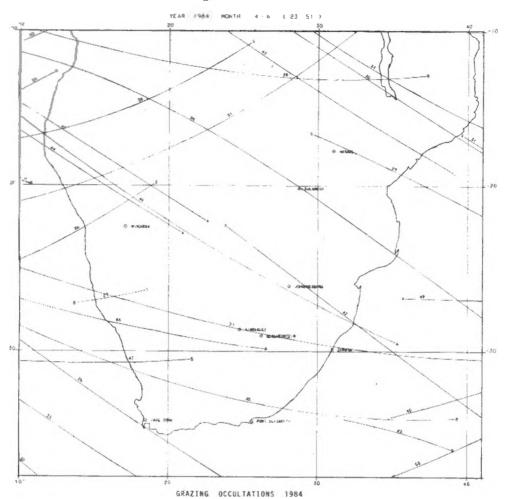
Surlit Limit northern limit (a complete occultation takes place south of

S = southern limit (complete occultation north of the track).

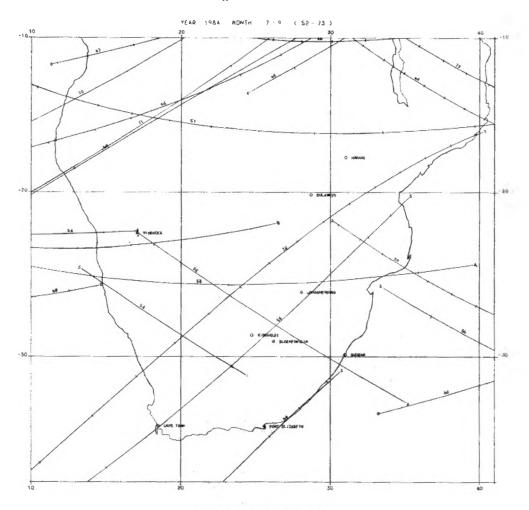


GRAZING OCCULTATIONS 1984

SEQ	NZC NO	MAG	MON	DAY	н	н	S	SUNLIT &	LIMIT
1	128	7.29	1	10	21	30	19.62	44.64	S
2	2020	6.56	1	24	23	44	17.27	-51.59	S
3	2032	7.29	1	25	3	50	25.13	-50.16	S
4	83	6.92	2	6	1.7	43	48.76	18.66	S
5	202	6.96	2	7	18	15	9.27	26.87	S
6	2088	6.24	2	21	21	21	55.74	-68.57	S
7	2209	5.92	2	22	20	54	15.30	-58.40	S
8	2376	4.57	2	24	1	34	3.80	-45.74	S
9	2507	6.72	2	25	0	5	56.02	-36.12	S
10	2510	6.26	2	25	0	46	18.30	-35.91	S
11	2513	4.28	2	25	1	6	21.93	-35.69	S
12	2672	2,94	2	26	4	46	52.60	-25.44	Š
13	2809	4.93	2	27	1	6	52.05	-18.34	Š
14	283	7.05	3	6	17	51	17.75	13.48	S
15	631	5.56	3	9	16	40	10.88	38.67	Š
16	656	4.36	3	9	20	57	3.99	39.97	N
17	657	5.42	3	9	21	4	18.95	39.95	N
18	1251	5.87	3	13	1.7	20	9.64	80.85	S
19	2622	6.27	3	24	3	21	14.26	-51.56	Š
20	2627	6.89	3	24	4	44	16.46	-51.22	5
21	3037	7.26	3	27	1	46	10.12	-23.82	S
22	3164	4.72	3	28	1	34	32.56	-16.30	S

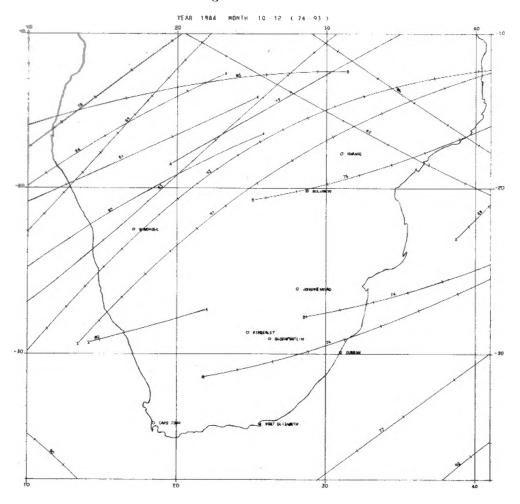


			KET	V T0	МАР	2			
SEQ	NZC NO	MAG	HON	DAY	Н	М	S	SUNLIT %	LIMIT
23 24 25 26 27 28 29 30	742 902 1206 1363 2879 3130 852 1169	5.99 6.61 5.88 5.22 6.60 5.47 5.00 5.40	4 4 4 4 4 5 5	6 7 9 10 22 24 4	16 17 17 22 4 2 15	6 44 47 19 26 0 51 56	3.47 29.66 36.50 20.02 17.69 37.47 16.33 3.49	23.89 34.07 55.97 68.41 - 58.09 · 39.05 12.18 30.25	\$ N N N S
31 32 33 34 35 36 37 38 39 40	1424 1432 1570 1689 2804 3089 3202 3214 128 237	6.80 7.00 5.64 5.47 5.86 5.27 6.14 6.61 7.29 7.14	សិតិសិតិសិសិសិសិសិសិសិសិសិសិសិសិសិសិសិស	8 9 10 18 21 21 22 26 27	15 17 21 21 23 4 23 4 23	22 14 40 4 38 23 28 4 56 3	20.87 36.98 47.17 43.99 38.39 42.62 23.22 19.72 18.75 58.28	51.62 52.46 65.32 75.77 - 83.31 - 64.73 - 56.41 - 55.40 - 18.79 - 12.45	N N N S N S N S N
41 42 43 44 45 46 47 48 49 50 51	1131 1263 1270 1274 1533 1773 2822 3364 3506 76 83	7.15 7.10 6.06 5.73 7.17 5.10 5.53 4.72 6.27 5.93 6.92	6666666666	2 3 3 5 7 10 18 20 22 22	17 16 17 18 19 22 0 2 21 0	24 11 24 33 48 31 10 59 55 55	40.01 25.09 40.14 6.67 38.56 19.98 51.53 4.18 4.95 11.95 24.38	10.20 17.66 17.96 18.37 39.16 63.51 64.15 - 79.96 - 54.35 - 43.85 - 43.48	N 性 N N M M M M



GRAZING	OCCULTATIONS	1984
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			K	EY TO	HA	P 3			
SEQ	MZC NO	MAG	MON	DAY	Н	Н	S	SUNLIT 1	LIMIT
52	1363	5.22	7	1	16	28	44.85	8,20	N
53	1365	6.09	7	1	16	54	9.81	8.31	N
54	1985	7.05	7	6	23	7	10.66	61.42	H
55	380	7.40	7	22	2	58	14.85	-40.12	N
56	1689	5.47	7	31	15	30	3.22	13.55	N
57	2060	6.30	8	3	19	13	57.45	46.09	N
58	2064	6.48	8	3	20	- 1	18.73	46.31	R
59	2170	6.80	8	4	16	3	52.41	56.27	N
60	2307	4,13	8	5	16	2	26.55	67.09	N
61	2330	6,31	8	5	20	52	28.17	68.45	N
62	2327	6.70	8	5	21	8	35.56	68.34	S
63	219	5.12	8	16	22	14	9.27	-75.81	N
64	586	6.95	8	20	3	42	54.31	-46.36	N
65	716	6.17	8	21	0	18	18.73	-37,21	N
66	2267	5.06	9	1	17	45	50.80	41.96	N
67	2282	5.87	9	1	20	24	39.92	42.89	S
68	2430	7.04	9	2	20	4 /	26.55	53,77	N
69	2750	2.14	9	4	23	10	46.62	74.58	S
70	660	4.40	9	16	22	56	37.03	-63.61	H
71	664	5.41	9	16	23	31	40.48	-63.42	N
72	676	7.08	9	17	1	39	17.95	-62.84	N
73	2500	3.37	9	30	12	32	16.87	35.89	N



GRAZING OCCULTATIONS 1984

KEY TO MAP 4

SEQ	N2C NO	MAG	MON	DAY	н	М	S	SUNLIT %	LIMIT
74 75 76 77 78 79 80	2669 2673 2676 2852 2985 1067 2159	6.23 6.28 6.46 7.36 6.88 7.19	10 10 10 10 10	1 1 2 3 16	16 17 17 20 21 23	54 3 25 20 44 27	53.46 12.06 18.57 27.50 36.13 6.76	47.63 47.70 47.86 59.00 69.05 -58.91	S S S S
81 82 83 84 85 86	2459 2750 3202 3214 1432 2500	5.28 7.22 2.14 6.14 6.61 7.00 3.27	10 10 10 11 11 11	25 27 29 1 1 16 24	17 18 10 18 22 0	29 32 12 58 58 40 37	43.88 40.87 28.89 43.00 9.94 55.88 3.69	1.96 13.70 29.69 61.99 62.98 -52.28 2.73	N S N S S N N
97 88 89 90	3164 3506 76 1755	4.72 6.27 5.93 6.85	11 12 12 12	28 1 2 16	21 16 18 2	11 22 51 53	35.01 16.83 0.07 38.94	35.47 62.80 72.57 -44.30	\$ \$ \$ \$
91 92 93	3478 37 258	6.51 7.47 6.62	12 12 12	28 29 31	18 18 19	35 50 0	27.23 31.74 6.24	36.15 45.51 64.62	\$ \$ \$

TIME SYSTEMS AND TELESCOPE SETTINGS

This section is intended to serve established amateurs and professional astronomers - i.e. those having some knowledge of time and coordinate systems. Space in the booklet does not permit full explanation, which in any case would appear complicated to the layman.

TIME SIGNALS FROM RADIO STATION ZUO

Radio signals of mean solar time are generated by the Precise Physical Measurements Division of the National Physical Research Laboratory in Pretoria. They are broadcast by the Post Office, the 2.5 and 5 MHz signals from Olifantsfontein, and the 100 MHz signals from Johannesburg.

Carrier Frequency	Radiated Power	Time of Transmission
2,5 MHz	4 Kw	2000 - 0600 SAST
5 MHz	4 Kw	Continuous

The signals consist of one pulse per second, each pulse consisting of 5 cycles of 1000 Hz tone. The first pulse in every minute is lengthened to 500 milliseconds. Morse code announcements are made during the minute preceding every fifth minute. They consist of the call sign ZUO (repeated 3 times) and the Universal Time (formally known as Greenwich Mean Time) at the next minute. (A special coding indicating UTI minus UTC is also indicated in the first 15 seconds of the minute by slightly lengthened second pulses).

SOUTH AFRICAN STANDARD TIME

South African Standard Time (as in everyday use) is mean solar time and the 30° East meridian (which runs east of Johannesburg and just west of Durban) and is exactly 2 hours ahead of Universal Time.

TIME OF SUN'S TRANSIT OVER 30° MERIDIAN

The table below gives the SAST when the Sun transits the 30° meridian and a sundial on that meridian reads noon.

jan 1	12 h 03 m 04 s	May 11	11 h 56 21 s	Sep 18	11 h 54 m 11 5
1.1	12 07 33	2 1	11 56 31	2.8	11 50 42
2 i	12 11 02	3 1	11 57 35	Oct 8	11 47 36
3 1	12 13 19	Jun 10	11 59 20	18	11 47 30
Feb 10	12 14 16	2 0	12 01 26	2 8	11 43 49
2 0	12 13 54	3 0	12 03 33	Nav 7	11 43 40
Nar 2	12 12 15	Jul 10	12 05 16	1 7	11 44 55
1 2	12 09 52	2 0	12 06 16	2 7	11 47 34
2 2	12 06 59	3 0	12 06 23		
Apr 1	12 03 58	Aug 9	12 05 31	Dec 7	11 51 23
1 1	12 01 07	1 9	12 03 37	1 7	11 56 01
2 1	11 58 45	2 9	12 00 58	2 7	12 00 59
May 1	11 57 05	Sep 8	11 57 43	3 1	12 02 56

Time Systems

SIDE	REAL	TIME ON	THE 30°	MERIDIAN						
		At O hrs SAST	At 21 hrs SAST		At O hrs SAST	At 21 hrs SAST			At O hrs SAST	At 21 hrs SAST
Jan	1	6 ^h 39 ^m	3 ^h 43 ^m	May II	15 ^h 16 ^m	12 ^h 19 ^m	Sep	18	23 ^h 48 ^m	20 ^h 52 ^m
	11	7 18	4 22	21	15 55	12 58		28	0 29	21 31
	21	7 58	5 01	31	16 34	13 38	0ct	8	1 07	22 10
	31	8 37	5 41	Jun 10	17 14	14 17		18	1 46	22 50
Feb	10	9 17	6 20	20	17 53	14 57		28	2 26	23 29
	20	9 56	7 00	30	18 33	15 36	Nov	7	3 05	0 10
Mar	2	10 40	7 43	Jul 10	19 12	16 16		17	3 45	0 48
	12	11 19	8 22	20	19 51	16 55		27	4 24	1 27
	22	11 58	9 02	30	20 31	17 34	Dec	7	5 03	2 07
Apr	I	12 38	9 41	Aug 9	21 10	18 14		17	5 43	2 47
	11	13 17	10 21	19	21 50	18 53		27	6 22	3 26
	21	13 57	11 00	29	22 29	19 33		31	6 38	3 42
May	l	14 36	11 40	Sep 8	23 09	20 12				

CORRECTION FOR PLACES NOT ON THE 30° MERIDIAN

Approximate longitude corrections from the 30° East Meridian are provided below. To find time of Sun's transit over local meridian, apply the longitude corrections to the data in the table above.

To find the sidereal times at SAST 0 hrs and SAST 21 hrs apply the corrections with the sign reversed to the data in the table.

Bloemfontein	+15 ^m	East London	+ 8 ^m	Port Elizabeth	+18 ^m
Bulawayo	+ 6 ^m	Grahamstown	+14 ^m	Pretoria	+ 7 ^m
Cape Town	+46 ^m	Johannesburg	+ 8 ^m	Harare	- 4 ^M
Durban	- 4 ^m	Kimberley	+21 ^m	Windhoek	+52 ^m

TELESCOPE SETTING

When a telescope equipped with setting circles is aimed on the meridian, its R.A. circle should read the sidereal time. Thus once can calculate the sidereal time and then set the circle, but is is usually simpler to aim the telescope at one of the well known stars given below and then to adjust the R.A. circle.

A LIST OF BRIGHT STARS FOR CHECKING TELESCOPE CIRCLES

Star	R.A.	Dec.	Mag.	Sp.	Star	R.A.	Dec.	Mag.	Sp.
Achernar	1 ^h 37 ^m , 1	-57°19	0,6	B 5	Procyon	7 ^h 38, 4	+ 5°161	0,5	F5
Aldebaran	4 35,0	+16 29	1,1	K5	Regulus	10 07,5	+12 03	1,3	88
Rigel	5 13,7	- 8 13	0,3	88	Spica	13 24,3	-11,05	1,2	B2
Betelgeuse -	5 54,3	+ 7 24	0,4	MO	Arcturus	14 14,9	+19 16	0,2	KO
Canopus	6 23,6	-52,41	-0,9	F0	Antares	16 28,4	-26 24	1,2	MI
Sirius	6 44,4	-16,42	-1,6	AO	Altair	19 50.0	+ 8 49	0.9	A5

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1960	W H van den Bos	1970	J C Bennett		

JULIAN DATE AT 1400 HOURS - 1984

	Jan.	Feb.	Mar.	Apr.	May	June	July	Aug.	Sep.	Oct.	Nov. Dec	٥.
	2445	2445	2445	2445	2445	2445	2445	2445	2445	244	2446 244	16
1 2 3 4 5	701 702 703 704 705	732 733 734 735 736	761 762 763 764 765	792 793 794 795 796	822 823 824 825 826	853 854 855 856 857	883 884 885 886 887	914 915 916 917 918	945 946 947 948 949	5975 5976 5977 5978 5979	007 03 008 03 009 03	36 37 38 39 40
6 7 8 9 10	706 707 708 709 710	737 738 739 740 741	766 767 768 769 770	797 798 799 800 801	827 828 829 830 831	858 859 860 861 862	888 889 890 891 892	919 920 921 922 923	950 951 952 953 954	5980 5981 5982 5983 5 984	012 04 013 04 014 04	41 42 43 44 45
11 12 13 14 15	711 712 713 714 715	742 743 744 745 746	771 772 773 774 775	802 803 804 805 806	832 833 834 835 836	863 864 865 866 867	893 894 895 896 897	924 925 926 927 928	955 956 957 958 959	5985 5986 5987 5988 5989	017 04 018 04 019 04	46 47 48 49 50
16 17 18 19 20	716 717 718 719 720	747 748 749 750 751	776 777 778 779 780	807 808 809 810 811	837 838 839 840 841	868 869 870 871 872	898 899 900 901 902	929 930 931 932 933	960 961 962 963 964	5990 5991 5992 5993 5994	022 05 023 05 024 05	51 52 53 54 55
21 22 23 24 25	721 722 723 724 725	752 753 754 755 756	781 782 783 784 785	812 813 814 815 816	842 843 844 845 846	873 874 875 876 877	903 904 905 906 907	934 935 936 937 938	965 966 967 968 969	5995 5996 5997 5998 5999	027 05 028 05 029 05	56 57 58 59 60
26 27 28 29 30 31	726 727 728 729 730 732	757 758 759 760	786 787 788 789 790 791	817 818 819 820 821	847 848 849 850 851 852	878 879 880 881 882	908 909 910 911 912 913	939 940 941 942 943 944	970 971 972 973 974	6000 6001 6002 6003 6004 6005	032 00 033 00 034 00 035 00	61 62 63 64 65 66

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