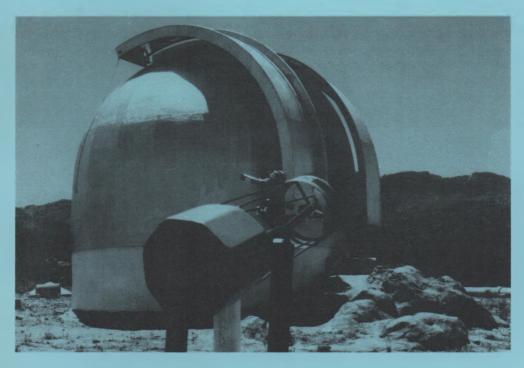
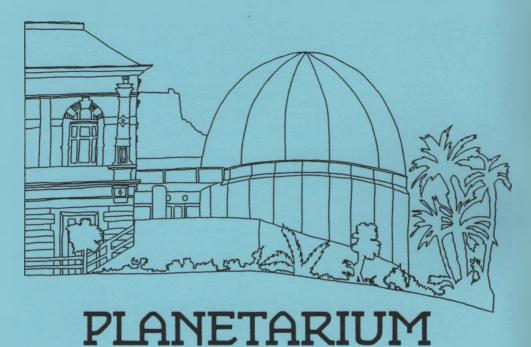
ASTRONOMICAL HANDBOOK FOR SOUTHERN AFRICA

1991



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ASTRONOMICAL HANDBOOK FOR SOUTHERN AFRICA 1991

This booklet is intended both as an introduction to observational astronomy for the interested layman - even if his interest is only a passing one - and as a handbook for the established amateur or professional astronomer.

Front cover
The dome of the Cederberg Observatory.
The Cederberg Observatory is 250 km north, by road, of Cape Town. This observatory, founded by Dr. P. Mack, is being operated and further developed by a group of 5 amateur astronomers. In the foreground is a 225 cm amateur made reflector telescope.
Photograph courtesy of Mr C. Forder.

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NOTE

All times are SAST unless otherwise stated. Right Ascension and Declination are given for equinox of date unless otherwise stated.

This handbook is produced for the Astronomical Society of Southern Africa. The data it contains has been adapted for Southern Africa from data obtained from The Astronomical Almanac for 1991, the Handbook of the British Astronomical Association, the International Lunar Occultation Centre, Tokyo. The star charts on pages 30, 32, 34 and 36 are from "A Beginner's Guide to the Southern Stars" by J.S. Bondietti, published by the South African Museum. The Minor Planet Occultations were provided by Edwin Goffin, who wishes to thank Dr. Josef de Kerfo, General Manager of Agfa-Gevaert IVV (Mortsel, Belguim) for making the computing facilities available.

Assistance in the compilation of this booklet was received from the Directors of the observing sections of the ASSA and Mrs. P. Kramer.

Further copies of this booklet are available at R5,00 per copy from The Business Manager, Astronomical Society of Southern Africa, P 0 Box 9, Observatory, 7935. All other correspondence concerning this booklet should be addressed to the Handbook Editor, Astronomical Society of Southern Africa, 10 Bristol Rd., Observatory, 7925.

Although every care has been taken in the compilation of the Handbook, it is sold and distributed on the explicit understanding that neither the Society nor its members accept any responsibility for errors therein.

Dit spyt ons dat as gevolg van beperkte fondse en produksie fasiliteite dit nie moontlik is om die handboek in Afrikaans te laat druk nie.

ASTRONOMY IN SOUTHERN AFRICA

As one of the few parts of the Earth having both access to the rich southern skies and a suitable climate, Southern Africa holds a favoured position in astronomy. Consequently, it has seen the establishment of a number of professional observatories engaged in research while many individuals have become enthusiastic amateur astronomers. Planetaria and visiting nights at observatories convey to the general public much of what goes on in this field.

OBSERVATORIES

The South African Astronomical Observatory (SAAO) was established in 1972, as a joint venture between the Council for Scientific and Industrial Research (CSIR) of SA and the Science Research Council of the UK,combining the facilities of the former Royal and Republic Observatories, and is directed by Prof M W Feast. On the closure of the Radcliffe Observatory in 1974, the CSIR acquired the 1.9-m telescope, and moved it to Sutherland in the Karoo, where there are also 1.0-m, 0.75-m and 0.5-m telescopes. The headquarters in Cape Town also carrys out a limited amount of observing. Research is undertaken in many areas, with considerable effort being put into the study of variable stars, the Galactic Centre, the Magellanic Clouds and sources detected by satellites. These studies involve the use of spectroscopic, photometric and infrared techniques. Besides providing research facilities for its own staff, SAAO observing time is allocated to astronomers from South African universities and elsewhere in the world.

Boyden Observatory, situated at Mazelspoort, 25 km from Bloemfontein, is owned by the Dept of Physics and Astronomy of the University of the Orange Free State. Observing facilities include a 1.52-m and two 0.41-m telescopes, as well as the 0.25-m Metcalf camera, a 0.33-m refractor and a 0.20-m solar installation. The main research areas include flare stars, short period variable stars, and atomic emissions from nebulae, the Sun and interplanetary space.

The Hartebeesthoek Radio Astronomy Observatory, 30 km NW of Krugersdorp, is a national facility managed by the Foundation for Research Development. The Director is Dr G D Nicolson. The 26 m telescope operates at 18, 13, 6, 3.6 and 2.5 cm wavelengths and is used for observations of interstellar and circumstellar molecules, pulsars, x-ray sources as well as quasars and active galaxies. The observatory provides research facilities for astronomers in South African universities as well as its own staff and frequently collaborates in global networks of telescopes using the technique of very long baseline interferometry.

The Nooitgedacht Gammaa Ray Telescope, established in 1985 in the Vredefort area south of Potchefstroom, is operated as a facility of the FRD/PU Cosmic Ray Research unit of the Potchefstroom University, under the leadership of Prof B C Raubenheimer. It consists of twelve parabolic mirrors with a total reflecting area of 21 square metres. The weak blue Cerenkov light emitted by high energy gamma rays in the atmosphere is detected by fast coincidence techniques. Radio pulsars, X-ray binaries, Supernova Remnants and Cataclysmic Variables are some of the objects studied.

In addition to the professional observatories mentioned above, South Africa and Zimbabwe have numerous private observatories, built and operated by amateur astronomers.

OBSERVATORIES OPEN TO THE PUBLIC

SAAO headquarters in Observatory, Cape Town is open to visitors on the second Saturday of each month at 20h00. It is not necessary to make a booking, unless there are more than ten persons in a party. Day visits are possible to the SAAO observing site near Sutherland, and enquiries should be made to Sutherland prior to the intended visit.

Enquiries as to visiting nights at Boyden Observatory, Bloemfontein should be made to the Dept of Physics and Astronomy of the University of the Orange Free State. The Nooitgedacht Telescope, Potchefstroom. Interested individuals or groups are welcome to contact Prof. B C Raubenheimer to arrange visits.

The Port Elisabeth Peoples Observatory Society. The Observatory, situated on the corner of Westview Drive and MacFarlane Road, is open to the public on the 1st and 3rd Wednesdays of every month and on every Wednesday during December and January. Admission is free. Donations are accepted to help with running costs. Viewing evenings are arranged for groups at other times during the month. Society meetings are held bi-monthly on the 3rd Monday. Secretarial address: P. O. Box 7988, Newton Park, Port Elizabeth. 6055.

PLANETAR IA

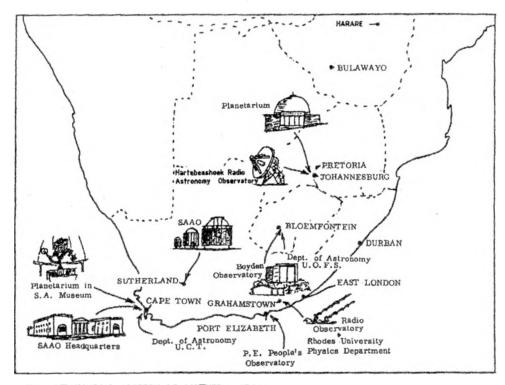
A planetarium is located within the South African Museum in Cape Town. This new planetarium, containing a Minolta Series 4 projector and seating 120, represents one of the most advanced of its kind in the world today.

A planetarium is situated in the grounds of the University of the Witwatersrand in Johannesburg (entrance in Yale Road, alongside the M1). It is equipped with a Zeiss projector and seats over 400 persons.

Regular shows are given at both of these planetaria, from which details may be obtained.

UNIVERSITIES

Several universities undertake research in astronomy and offer teaching courses. The chair of astronomy at UCT is occupied by Brian Warner, whose department uses the SAAO observing facilities at Sutherland. The Dept of Applied Mathematics, UCT has a group carrying out research in theoretical cosmology lead by Profs G F R Ellis and D R Matravers. The University of OFS has a Dept of Physics and Astronomy, headed by Prof. P E Viljoen, incorporated with the Boyden Observatory. The Dept of Physics and Electronics at Rhodes University, under Prof E E Baart, specialises in radio astronomy, and has its own observatory outside Grahamstown. The Dept of Mathematics, Applied Mathematics and Astronomy at UNISA offers a number of courses in astronomy and astrophysics. Prof W F Wargau is the head of Astronomy at UNISA. Courses in Gamma Ray Astronomy and General Astrophysics form part of the regular honors and masters courses of the Department of Physics at Potchefstroom.



THE ASTRONOMICAL SOCIETY OF SOUTHERN AFRICA

The Astronomical Society of Southern Africa is a body consisting of both amateur and professional astronomers. Membership is open to all interested persons, regardless of knowledge or experience. In addition to this Handbook, the Society issues the "Monthly Notes of the Astronomical Society of Southern Africa" (MNASSA). Members also receive the popular monthly magazine "Sky and Telescope" published in the USA, which provides information on professional and amateur activities, together with news of space research and other related subjects. The Society's annual

subscription is R75.00 and there is an entrance fee of R10.00. A prospectus and application form may be obtained from the Honorary Secretary, Astronomical Society of Southern Africa, c/o S A Astronomical Observatory, P O Box 9, Observatory 7935.

LOCAL CENTRES OF THE SOCIETY

Autonomous local centres of the Society hold regular meetings in Cape Town, Durban, Johannesburg, Bloemfontein, Pietermaritzburg, Pretoria and Harare. Visitors are very welcome at meetings and may, if they wish, join a Centre without becoming a full member of the Society. Centre members receive neither Society publications, nor "Sky and Telescope".

CAPE CENTRE (Cape Town): Formal meetings are held on the second Wednesday of the month (except in January and December) when professional and prominent local amateur astronomers present lectures on the latest topics in Astronomy. Informal meetings are held on most other Wednesdays except during January and December. At these informal meetings discussion groups and observing sessions are held. Meetings are held at the SAAO. Observatory Road, Observatory at 20h00. Two or three out-of-town weekend observing sessions are held annually in areas where dark skies are available. The Centre publishes a quarterly journal, the "Cape Observer" which gives information on meetings, current activities and objects to observe. Further information may be obtained from the Secretary, P.O.Box 13018, Mowbray, 7705, or telephone (021) 725897.

TRANSVAAL CENTRE (Johannesburg): General meetings are held on the second Wednesday of each month, excluding December, in the Sir Herbert Baker building in the grounds of the former Republic Observatory, 18A Gill Street, Observatory, Johannesburg at 20h00. The meetings consist of lectures, films or observing evenings. There are two small observatories on the site. One contains the Jacobs telescope, a 30cm F8 Newtonian and in the Papadopoulos Dome is housed a combined instrument comprising a 18cm F16 refractor, a 15cm refractor and a 30cm F16 Cassegrain reflector. Informal observing evenings are held every Friday night. The Centre publishes a monthly newsletter "Canopus", which carries information on meetings and the Centre's activities. Secretarial address: P O Box 93145, Yeoville 2143, telephone (011) 8865602.

NATAL CENTRE (Durban): Regular monthly meetings are held at 19h45 on the second Wednesday of each month at Marist Brothers School, South Ridge Road, Durban. The Centre publishes a monthly magazine "Ndaba" which contains news, views and current information on astronomical and related topics. Secretarial address: P O Box 5330, Durban, 4000, or telephone (031) 3072093 / 7011104 / 7013178 / (0323) 51647.

NATAL MIDLANDS CENTRE (Pietermaritzburg): Regular monthly meetings on the second Wednesday of each month at 19h45. Information on activitys and membership is available from the Secretary, P O Box 2106, Pietermaritzburg, 3200 or by phoning (0331) 33646.

BLOEMFONTEIN CENTRE: Meetings are held every fourth Friday of the month. For information contact the Secretary, Mrs S Rabe, P O Box 1238, Bloemfontein, 9300 or telephone (051) 224977.

PRETORIA CENTRE: Meetings are held on the fourth Wednesday of each month (except December) at 19h00 at the Christian Brothers' College, Silverton Road, where the Centre's observatory containing a 30cm reflecting telescope is situated. For further information contact the Secretary Mr N Young at 201 Kritzinger St., Meyers Park, Pretoria, 0184 telephone 012-833765.

HARARE CENTRE: The Centre holds a meeting on the last Wednesday of each month (except December). These are usually held at 17h30 at the Queen Victoria Museum and consist of lectures, films or general discussions. Informal observing sessions are also held at the homes of members. Secretarial address P O Box UA 428, Union Avenue, Harare, Zimbabwe.

OBSERVING SECTIONS OF THE SOCIETY

These sections exist to co-ordinate constructive observing programmes. More information on a section is given in the appropriate parts of this handbook.

DIARY OF PHENOMENA

	DIARY O	F PHENC	DMENA
d h. Jan 1 17 Venus 1'-2 S. ol 1 18 Mars stationary 3 1 Jupiter 2 M. ol 3 5 Earth at perihe 3 21 Mercury station 5 5 Weptune in conj 5 10 Vesta stationar 7 21 LAST QUARTER 12 5 Antares 0'-6S. 12 13 Moon at apogee 13 22 Hercury 4 N. o 14 11 Mercury greates 16 2 NEW MOOM 17 20 Venus 3' S. of 18 10 Saturn in conju 23 16 FIRST QUARTER 23 19 Mercury 0'-4 N. 23 23 Pallas etationa	f Moon lion ary unction with Sun y of Moon f Moon t elong. W.(24*) Moon nction with Sun of Uranus		d h 8 23 Saturn 1'-4 S. of Moon 10 2 Pallas stationary 14 22 MEW MOON 14 23 Hercury in inferior conjunction 17 18 Ceres at opposition 17 18 Venus 2' S. of Moon 17 19 Moon at perigee 18 14 Uranus stationary 19 1 Neptune stationary 19 1 Neptune stationary 20 2 Mars 0'-6 M. of Moon 21 15 FIRST QUAPTER 21 21 Jupiter 1'-9 M. of Moon 22 6 Venus 7' N. of Aldebaran 27 8 Mercury stationary 28 23 FULL MOON 3 17 Moon at apogee 4 10 Uranus 0'-3 M. of Moon
23 23 Pallas stationa 25 17 Mars 2' S. of M 26 16 Mercury 1'-1 S. 28 11 Moon at perigee 29 2 Jupiter at oppo 30 7 Jupiter 1'-8 M. 30 8 FULL MOOM	oon of Meptune sition of Moon Penumbral Eclips	e	4 16 Neptune 1 · 2 N. of Moon 6 9 Saturn 1 · 8 S. of Moon 7 3 LAST QUARTER 10 5 Pluto at opposition 12 15 Mercury 9 S. of Moon 12 10 Mercury greatest elong. W. (26) 14 7 NEW MOON
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Apr 4 5 Antares 1°-1 S 4 19 Mercury station 5 23 Moon at apogee 7 2 Oranus 0°-6 N. 7 8 Beptune 1°-5 M 7 9 LAST QUARTER	nary	cn.	8 2 Meptune at opposition 9 21 Vesta in conjunction with Sun 11 10 Venus 1 · 0 S. of Regulus 11 12 Moon at perigee 11 21 TEW MOON 13 16 Mercury 3 N. of Moon

CONFIGURATIONS OF SUN, MOON AND PLANETS

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Jul 13 19 Jupiter 3° N. of Moon
                                                                                                                Oct 1 2 LAST QUARTER
       13 19 Jupiter 3 N. of Moon
14 17 Mars 5 N. of Moon
14 18 Mars 0 7 N. of Regulus
14 20 Venus 3 N. of Moon
15 10 Mercury 0 08 S. of Jupiter
17 1 Juno at opposition
17 7 Venus greatest brilliancy
18 17 FIRST QUARTER
22 8 Venus 4 S. of Mars
24 13 Moon at appose
                                                                                                                         2 20 Moon at perigee
                                                                                                                        3 19 Mercury in superior conjunction
4 17 Venus 0 · 2 N. of Moon
                                                                                                                         5 4 Saturn stationary
5 7 Jupiter 5 N. of Moon
6 2 Mars 3 N. of Spica
                                                                                                                        8 0 NEW HOON
                                                                                                                     8 6 Venus 3 S. of Regulus
14 20 Uranus 0 06 S. of Moon
15 4 Meptune 0 7 N. of Moon
        24 13 Moon at apogee
24 23 Uranus 0 -4 N. of Moon
       25 4 Mercury greatest elong. E.(27')
25 7 Weptume 1 ·2 N. of Moon
26 20 FCLL MOON Penumbral Eclipse
26 22 Saturn 1 ·9 S. of Moon
                                                                                                                      15 13 Moon at apogee
15 20 FIRST QUARTER
                                                                                                                     16 14 Saturn 2° S. of Moon
17 5 Venus 2° S. of Jupiter
                                                                                                                      23 13 FULL MOON
27 18 Moon at perigee
30 9 LAST QUARTER
        27 2 Saturn at opposition
27 3 Mercury 2 S. of Regulus
        30 6 Venus stationary
Aug 2 21 Pluto stationary
                                                                                                              Nov 1 22 Jupiter 6' N. of Moon
         3 13 LAST QUARTER
                                                                                                                         2 11 Venus greatest elong. W.(47°)
2 23 Venus 6 N. of Moon
         7 6 Mercury stationary
7 8 Mercury 2 N. of Venus
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         8 20 Moon at perigee
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8 7 Mercury 0°-7 N. of Moon
      8 20 Moon at perigee
10 4 MFW MOOM
11 9 Venus 3 S. of Moon
11 10 Mercurv 0 6 S. of Moon
12 10 Mars 6 N. of Moon
17 2 Venus 9 S. of Regulus
17 7 FIRST QUARTER
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11 5 Uranus 0 4 S. of Moon
11 9 Mercury 2 N. of Antares
11 13 Meptune 0 4 N. of Moon
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                                                                                                                    12 10 Moon at apogee
13 0 Saturn 2'S. of Moon
13 6 Pluto in conjunction with Sun
14 16 FIRST QUARTER
       18 O Jupiter in conjunction with Sun
       20 4 Moon at apogee
21 4 Uranus 0 4 M. of Moon
21 12 Neptune 1 2 N. of Moon
                                                                                                                     19 4 Mercury greatest elong. E.(22°)
22 1 FULL MOON
       21 23 Mercury in inferior conjunction
       22 22 Venus in inferior conjunction
23 1 Saturn 1 .8 S. of Moon
                                                                                                                      24 4 Moon at perigee
28 17 LAST QUARTER
       25 11 FULL MOON
29 7 Mercury 6 N. of Venus
                                                                                                                      28 20 Mercury stationary
29 10 Jupiter 6 N. of Moon
       30 22 Mercury stationary
                                                                                                                      29 11 Venus 4' W. of Spica
Sep 1 20 LAST QUARTER
                                                                                                               Dec 2 16 Venus 8° N. of Moon
        5 21 Moon at perigee
6 19 Venus 5 S. of Moon
                                                                                                                        6 6 NEW HOON
                                                                                                                    6 6 NEW MOOM
8 16 Cranus O'-6 S. of Moon
8 17 Mercury in inferior conjunction
8 22 Meptune O'-2 N. of Moon
10 4 Moon at apogee
10 12 Saturn J' S. of Moon
11 17 Mars 4 N. of Antares
11 21 Ceres in conjunction with Sun
13 18 Mercury J' N. of Mars
        6 20 Juno stationary
        7 7 Mercury 3 N. of Moon
7 13 Jupiter 5 N. of Moon
        7 20 Mercury greatest elong. W(18°)
8 13 NEW MOON
      10 4 Mars 6 N. of Moon
10 10 Jupiter 0 4 N. of Regulus
10 12 Mercury 0 07 S. of Jupiter
10 12 Mercury 0 3 N. of Regulus
                                                                                                                     14 12 FIRST QUARTER
                                                                                                                      16 O Mercury & N. of Antares
      12 4 Venus stationary
                                                                                                                    18 14 Mercury stationary
21 8 Mercury 7 N. of Antares
     16 O FIRST QUARTER
17 11 Uranus O -2 N. of Moon
17 17 Moon at apogee
                                                                                                                      21 12 FULL MOON
                                                                                                                    22 11 Moon at perigee
22 11 Solstice
      17 19 Weptune 1 • 0 W. of Moon
19 6 Saturn 1 • 8 S. of Moon
                                                                                                                    26 19 Jupiter 7° N. of Moon
      19 11 Oranus stationary
                                                                                                                    27 23 Mercury greatest elong. W.(22')
28 4 LAST QUARTER
      23 15 Equinox
      24 1 FULL WOON
26 7 Meptune stationary
                                                                                                                    31 14 Jupiter stationary
      29 1 Venus greatest brilliancy
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THE SUN

BASIC DATA:

Diameter: 1 392 000 km (109 times Earth diameter)

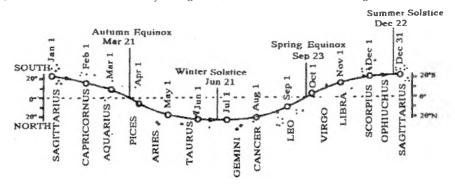
Mass: 1,99 x 10 10 kg (330 000 times Earth mass)

Surface Temperature: Approximately 6 000 C

Temperature at centre: Approximately 10 million C

The Sun is our nearest star. It is composed chiefly of hydrogen and is in a gaseous state throughout. So hot and dense is its interior that nuclear reactions occur there - thus producing the energy that is eventually radiated from its surface. At times its surface is disturbed by sunspots (which may persist for some weeks) and short-lived flares.

The Earth's orbit round the Sun is not quite circular. In 1991 we will be closest to the Sun on January 3 (perihelion - approximate distance 147 million km) and furthest from the Sun on July 6 (aphelion - approximately 152 million km). During the year, the Sun appears to us to make a complete circuit of the sky (i.e. relative to the starry background) as indicated in the diagram.



Permanent damage to the eye can be caused by looking directly at the Sun. The diagram below shows how a small telescope (or half a binocular) may be used to project an image of the solar disc onto a piece of white card. It may also be advisable to stop down the telescope aperture so that the eyepiece is not damaged by the intense light passing through it. Tiny black sunspots are generally visible on the otherwise white solar disc - if monitored over a period of a week or so, the rotation of the Sun should be apparent.

					Sunlight	-		<u>_</u>		Cardboard Telesco		Adh	yept	-	O Image				
THE	SUN'	S DI	CL	NAT	ION AT	02 F	HOUR!	S :											
Jan	1	23	4	S	Apr	11	8	3	'N	Jul	20	20°	8'	N	0ct	28	12°	'53'	S
	11	21	55	S	-	21	11	37	N		30	18	41	N	Nov	7	16	4	S
	21	20	4	S	May	1	14	51	N	Aug	9	16	4	N		17	18	48	S
	31	17	35	S		11	17	42	N		19	13	1	N		27	20	60	S
Feb	10	14	35	S		21	20	2	N		29	9	37	N	Dec	7	22	31	S
	20	11	11	S		31	21	49	N	Sep	8	5	58	N		17	23	20	S
Mar	2	7	29	S	Jun	10	22	57	N	•	18	2	9	N		27	23	21	S
	12	3	36	S		20	23	26	N		28	1	44	S		31	23	9	S
	22	0	21	N		30	23	13	N	Oct	8	5	36	S					
Apr	1	4	16	N	Jul	10	22	20	N		18	9	21	S					

TIMES OF SUNRISE AND SUNSET FOR THE MAIN CITIES OF SOUTHERN AFRICA

	CAPETOWN	ī	I	OURBA	٩N		BL	DEME	ONTI	EIN	JOI	IANN	ESBI	URG		HAR	ARE	
	sunrise sur	iset :	sunri	ise s	sur	iset	suni	cise	sui	nset	suni	rise	sui	nset	suni	cise	sui	iset
	h m h	m O 1		m	h	OIL .	h	m	h	m	h	m.	h	m O t	h		h	m
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Feb l	06 07 19	52	05 2	24 1	18	55	05	46	19	13	05	42	19	00	05	42	18	36
11	06 17 19	44	05 3	32 1	8	48	05	54	19	06,	05	49	18	55	05	47	18	32
21	06 26 19	33	05 4	+1 1	18	39	06	02	18	57	05	54	18	47	05	52	18	27
Mar 1	06 33 19	23	05 4	6 1	8	30	06	08	18	48	06	00	18	39	05	55	18	21
11	06 41 19	11	05 5	53 1	8	19	06	13	18	38	06	04	18	29	05	57		15
21	06 49 18	58	05 5	9 1	. 8	06	06	18	18	27	06	11	18	19	06	00	18	06
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11	07 04 18	30	06 1		.7	43	06	30	18	03	06	21	17	56	06	04		50
21	07 13 18	17	06 1	17]	.7	31	06	35	17	52	06	25	17	47	06	07	17	43
May 1	07 20 18	06	06 2	04 1	7	22	06	42	17	44	06	31	17	38	06	10	17	37
11	07 28 17	-	06 3			14	06		17	36	06		17	31	06			32
21	07 34 17	50	06 3	16 1	.7	06	06	54	17	30	06	41	17	26	06	16	17	29
Jun 1	07 43 17	45	06 4	.3 1	7	04	07	01	17	27	06	47	17	23	06	20	17	28
11	07 48 17	44	06 4	8 1	7	03	07		17	26	06	52	17	22	06	23	17	
21	07 51 17	44	06 5	1 1	.7	04	07	06	17	27	06	55	17	24	06	26	17	29
Jul 1	07 53 17	48	06 5	3 1	.7	07	07	10	17	30	06	57	17	27	06	27	17	32
11	07 51 17	52	06 5	1 1	.7	11	07	06	17	34	06	55	17	30	06	27	17	
21	07 47 17	58	06 4	8 1	.7	16	07	05	17	39	06	53	17	35	06	26	17	40
Aug 1	07 39 18	06	06 4	2 1	.7	22	07	00	17	45	06	48	17	41	06	23	17	42
11	07 30 18		06 3	4 1	7	29	06	53	17	51	06	41	17	46	06	18	17	
21	07 19 18	20	06 2	24 1	.7	35	06	42	17	55	06	32	17	50	06	11	17	48
Sep 1	07 06 18	27	06 1	2 1	7	40	06	31	18	01	06	21	17	54	06	04	17	49
11	06 52 18		06 0			46	06	19	18	06	06	11	17	59	05	55	17	-
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11	05 38 19		04 5	5 1	8.	26	05	17	18	44	05	13	18	32	05	14	18	
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ECLIPSES OF THE SUN

The annular eclipse of January 15 and the total eclipse of July 11 are not visible from Southern Africa.

COMPUTING AND SOLAR SECTIONS

COMPUTING SECTION

This section invites all those interested in Astronomical Computing in any form to share their expertise and any Software they may have with other members in the Society.

The objectives of the Computing Section were outlined in MNASSA. Vol 46 Nos. 5 & 6 June 1987 Page 66. Please refer to this write up for detailed information. Persons interested in the activities of the Computing Section are urged to contact the Director of the Societies Computing Section:

Mr Tony Hilton, P O Box 68846, Bryanston, 2021. Phone (w) (011) 53 8714 (h) (011) 465 2257.

Mr Hilton has compiled a comprehensive DATA BASE of all interested person's, equipment, available software etc. This report will be available to all interested persons. If you wish to become a subscriber to this DATA BASE list please contact Mr Hilton for the relevant quuestionnaire.

Furthermore, if you are embarking on any Computer Projects, Mr Hilton would like to hear from you, and would make himself or any other competent individuals available to supply expert advice or additional information where necessary.

SOLAR SECTION

The work undertaken by this section covers a broad range of techniques to observe activity on the sun's disk. Members of the section note their observations on appropriate forms, which are then forwarded to various organisations in the United States of America, the United Kingdom, the Federal Republic of Germany and South Africa. The results we provide are further reduced by these organisations and incorporated with the information provided by other world-wide groups of Solar Observers, such as ours. The data produced is then fed to over 450 scientific institutions all over the world, where it is used by a very wide range of scientific disciplines.

Observational techniques employed include the visual observation of the sun's disk (using suitable filters or by projecting the image onto an appropriate screen) to determine sun spots and active areas, the monitoring of solar flares by very low frequency radio waves and monitoring changes in the earth's magnetic field caused by solar activity. Other activities such as photographing and the drawing of visible solar features are also undertaken.

A weekly update of solar activity is broadcast country wide on Sunday mornings at 10h00 and again on Monday evenings at 20h30 by the Amateur Radio League, enabling amateur astronomers and amateur radio enthusiasts alike to keep abreast of the latest developments as well as with all other interesting solar information. Short lectures are also given at this time.

Towards the end of 1986, the Sun entered the new 11 year Solar cycle and this cycle is characterised by a steady climb over approximately 4½ years, followed by a slower decline to minimum lasting approximately 6½ years. 1991, will therefore experience a marked increase in Solar activity and provide the observer with much to look at.

A word of caution - NEVER observe the sun directly without adequate filtration as permanent eye damage can occur, and do not use the screw-in filters provided with some commercial telescopes, as they are inclined to shatter! Large instruments are not a prerequisite! Any telescope from 50mm, reflector or refractor can be used and provides an ideal opportunity for owners of small instruments to contribute immediately to Science. If one has no filter, then the only safe method is to project the image on to a white card. Image quality will be enhanced if the card is kept in the shade, or enclosed in a screen. Details of suitable filters are best sought from experienced solar observers or from the Director of the Solar Section.

Persons interested in observing the sun,or requiring information are invited to contact The Director of the Solar Section:

Jim Knight, at Mars Street, Atlasville, Boksburg, 1459 or at 011-9731380.

The Moon

BASIC DATA

Diameter: 3 480 km (0,27 of Earth)
Mass: 7,35 x 10²² x kg (1/81 of Earth)

Surface Gravity: 0,16 of Earth

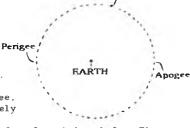
Average distance from Earth: 384 000 km

THE SURFACE OF THE MOON

In common with the bodies of our solar system, the Moon's surface suffered bombardment by numerous minor bodies during the period 4,5 to 3,0 billion years ago. This has produced the heavily cratered topography now visible. Some particularly large impacts, caused large circular depressions, which were flooded by molten lava from the Moon's interior. These are the maria basins which appear smoother and darker then the rest of the surface (the latin words mare and maria come from older times when the basins were mistaken for seas). The maria surfaces, being younger have fewer large craters, but the entire surface is peppered with tiny craters produced by tiny bodies which have also served to plough up the ground thus forming the regolith - a layer of loose material a metre or so deep.

THE MOON'S ORBIT

As a result of its motion around the Earth, the Moon appears to make a complete circuit of the heavens in just under a month. The Moon's orbit around the Earth is slightly elliptical; the Earth is situated at one of the foci of the ellipse. Thus the Earth-Moon distance varies slightly during the course of a revolution. Dates of Apogee, when the moon is furthest from the Earth (approximately 407 000 km) and of Perigee, when the Moon is closest to the Earth (approximately 357 000 km) are given on the next page.



SCALE DRAWING

The Moon is best positioned for evening viewing from 3 to 4 days before First Quarter to shortly after Full Moon.

ECLIPSES OF THE MOON

The three penumbral eclipses of the moon will be wholly or partly visible from Southern Africa. The Moon's entry of penumbra will be during morning twilight on January 30 and the middle of the eclipse will be close to sunrise and moonset. All of the eclipse of June 27 will be visible but that of July 26 will be in progress as the moon rises in the evening twilight. The partial eclipse of December 21 will not be visible.

The eclipse data is as follows:

	a	n	110.			a	n	D
Moon enters penumbra	Jan 30	5	57.8	Jun	27	3	46.	3
Middle of eclipse		7	58.6			5	14.	7
Moon leaves penumbra		9	59.4			6	43.	2

Contacts of Penumbra Position Angles with Limb of Moon from the North Point 67°.8 to East 25°.6 to West First Last Penumbral Penumbral

Magnitude of the Eclipse: 0.906

Jul 26 18 47.5 Moon enters penumbra 20 07.8 Middle of eclipse Moon leaves penumbra 21 28.3

Contacts of Penumbra Position Angles with Limb of Moon from the North Point First 135°.8 to East 173°.6 to West Last Penumbral Magnitude of the Eclipse: 0.280 Contacts of Penumbra Position Angles with Limb of Moon from the North Point 20°.2 to East 35°.6 to West First Last

Magnitude of the Eclipse: 0.339

TERMINATOR AND LIBRATION

During the changing phases, the terminator (the boundary between illuminated and dark portions) progresses from left to right in the diagram on page 10. Since the moon does not follow a perfectly circular orbit and its axis is not parallel to the Earth's axis, it is sometimes possible to see a slightly greater proportion of one limb than the opposite one. This effect is known as libration.

PHASES and VISIBILITY

May 15 19

Oct 2 20

NEW MOON А h h Jan 16 01 50 May 14 06 36 Sep 8 13 01 Feb 14 19 32 Jun 12 14 06 Oct 7 23 39 Mar 16 10 10 Jul 11 21 06 Nov 13 11 6 14 Apr 21 38 04 Aug 10 28 Dec 6 05 OF THE MOON'S ORBIT SCHEMATIC DIAGRAM Moon too close FIRST QUARTER LAST QUARTER to Sun to be visible h d m h d Jan 23 16 20 35 21 7 Jan Feb 22 00 52 58 15 Feb Mar 23 08 03 32 12 Noon Mar 8 Apr 21 14 39 45 08 7 Apr 20 21 May 46 46 7 02 May 19 Jun 06 19 17 30 5 Jun EARTH Jul 18 17 11 50 04 5 Jul 17 07 Aug 01 13 25 3 Aug Midnight 00 Sep 16 01 16 20 Sep 1 Oct 15 19 33 02 30 Moon visible in Moon visible in 1 Oct Nov 14 16 10 evening sky 09 30 morning sky Oct Dec 14 11 21 17 Nov 28 55 03 Dec 28 FULL MOON d m d h đ h 08 Jan 30 08 10 28 13 40 Mav 37 Sep 24 00 Feb 28 20 25 Jun 27 13 08 04 58 0ct 23 Mar 30 09 17 Jul 26 20 24 Nov 22 00 56 28 Apr 22 58 Aug 25 11 Dec 21 12 23 MOON at PERIGEE MOON at APOGEE A h d h d d Jan 28 11 Oct 15 13 Jun 13 Oct 27 18 Jan 12 13 May 31 5 Feb 25 Jul 11 12 Nov 24 4 Nov 12 10 9 Feb 6 Jun 27 9 Mar 22 Aug 8 20 Dec 22 11 Dec 10 4 Mar 9 3 Jul 24 13 Sep Apr 17 19 5 21

Apr

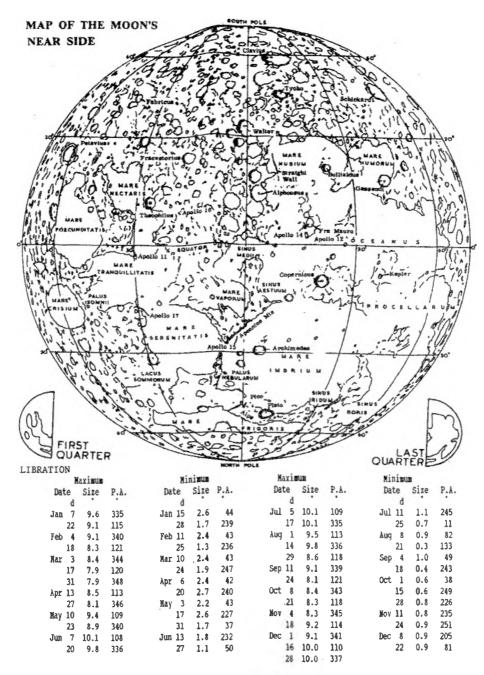
May

5 23

3 17

Aug 20

Sep 17 17



NOTE: Size of libration is given as an angle measured at the centre of the Moon. Position Angle (P.A.) is measured through East on the face of the Moon from the North point of the disk.

1991 TIMES OF MOON RISE AND SET CAPE FOWN

For PORT ELIZABETH subtract 28 MINUTES

							Subtract	EO MINUI	EO			
	JAM Rise	UARY Set	FEBRU Rise	UARY Set	Rise	ARCH Set	API Rise	RIL Set	Rise	\if Set	Ji Rise	UNE Set
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6 7 8 9	23 38 00 07 00 37 01 10	11 46 12 45 13 44 14 43 15 41	23 45 00 25 01 11 02 02	13 31 14 29 15 24 16 16 17 03	22 21 23 05 23 54 00 47	12 16 13 13 14 07 14 56 15 40	23 33 00 30 01 29 02 28	13 34 14 14 14 50 15 23 15 53	00 16 01 14 02 13 03 14	13 21 13 51 14 20 14 48 15 17	01 01 02 01 03 04 04 11 05 22	13 16 13 45 14 18 14 55 15 40
11 12 13 14 15	01 47 02 30 03 17 04 10 05 07	16 38 17 32 18 22 19 07 19 47	02 57 03 56 04 55 05 55 06 56	17 45 18 22 18 55 19 26 19 54	01 44 02 43 03 42 04 43 05 44	16 19 16 53 17 25 17 55 18 23	03 28 04 29 05 32 06 38 07 47	16 22 16 50 17 21 17 54 18 32	04 18 05 24 06 35 07 47 08 59	15 49 16 24 17 06 17 55 18 54	06 34 07 44 08 48 09 43 10 28	16 35 17 38 18 49 20 04 21 16
16 17 18 19 20	06 05 07 05 08 05 09 04 10 04	20 22 20 54 21 23 21 51 22 18	07 56 08 58 10 01 11 06 12 14	20 22 20 51 21 22 21 56 22 36	06 46 07 50 08 56 10 04 11 13	18 52 19 23 19 57 20 36 21 22	08 58 10 09 11 17 12 18 13 11	19 16 20 08 21 09 22 15 23 24	10 05 11 03 11 52 12 33 13 08	20 01 21 12 22 23 23 32	11 07 11 40 12 11 12 40 13 09	22 26 23 32 00 36 01 37
21 22 23 24 25	11 05 12 08 13 14 14 22 15 32	22 47 23 18 23 54 00 37	13 22 14 30 15 33 16 29 17 17	23 23 00 19 01 23 02 32	12 22 13 26 14 24 15 13 15 56	22 15 23 16 00 23 01 32	13 55 14 33 15 06 15 37 16 05	00 33 01 40 02 45 03 47	13 39 14 09 14 37 15 07 15 38	00 38 01 42 02 43 03 43 04 43	13 40 14 14 14 51 15 33 16 21	02 37 03 37 04 36 05 34 06 29
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16 17 18 19 20	05 42 06 38 07 34 08 29 09 24	19 28 20 04 20 37 21 09 21 41	07 18 08 15 09 14 10 14 11 17	19 43 20 16 20 52 21 30 22 14	06 05 07 04 08 05 09 09 10 15	18 16 18 51 19 30 20 13 21 02	08 00 09 08 10 15 11 17 12 12	18 55 19 49 20 50 21 56 23 02	09 04 10 04 10 56 11 41 12 21	19 42 20 50 21 58 23 03	10 17 10 56 11 31 12 04 12 38	21 54 22 55 23 54 00 36 00 51
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26 27 28 29 30 31	15 38 16 42 17 39 18 29 19 13 19 52	01 10 02 11 03 19 04 29 05 38 06 44	17 05 17 45 18 22	03 19 04 25 05 28	15 42 16 19 16 54 17 28 18 01 18 37	02 13 03 16 04 16 05 15 06 13 07 10	16 01 16 35 17 12 17 51 18 34	04 05 05 01 05 58 06 55 07 51	15 50 16 31 17 17 18 06 18 58 19 52	04 48 05 44 06 39 07 33 08 23 09 08	16 53 17 47 18 41 19 35 20 28	06 19 07 06 07 49 08 28 09 04
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1991 TIMES OF MOON RISE AND SET DURBAN

For BLOEMFONTEIN add 19 MINUTES

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21 22 23 24 25	10 09 2 11 10 2 12 13 2 13 20 2 14 28	21 59 22 32 23 10 23 55	12 18 13 24 14 27 15 25 16 15	22 42 23 38 00 42 01 49	11 17 12 21 13 20 14 11 14 55	21 34 22 35 23 41 00 48	12 55 13 35 14 11 14 43 15 15	23 48 00 52 01 55 02 55	12 45 13 17 13 48 14 20 14 53	00 50 01 48 02 46 03 44	12 55 13 30 14 09 14 52 15 40	01 39 02 37 03 34 04 30 05 25
26 27 28 29 30 31	16 39 0 17 35 0 18 24 0	00 47 01 48 02 56 04 08 05 19 06 27	16 58 17 37 18 11	02 58 04 07 05 12	15 34 16 09 16 42 17 14 17 45 18 19	01 55 02 59 04 02 05 03 06 03 07 02	15 46 16 18 16 53 17 31 18 13	03 54 04 52 05 50 06 49 07 47	15 30 16 10 16 55 17 44 18 36 19 31	04 42 05 39 06 36 07 29 08 19 09 04	16 31 17 25 18 20 19 16 20 11	06 16 07 03 07 45 08 22 08 56
1 2 3 4 5	h m 21 06 0 22 01 0 22 57 1 23 55 1	Set h m 99 27 99 57 10 26 10 56 11 29	Rise h m 22 47 23 49 00 53 02 00	Set b m 09 31 10 05 10 43 11 28 12 21	SEPTI Rise h m 00 54 01 56 02 52 03 42	Set h m 10 15 11 12 12 15 13 23 14 33	OCTO Rise h m 00 46 01 37 02 22 03 02 03 38	Set h m 11 14 12 21 13 28 14 34 15 38	Rise h m 01 38 02 12 02 45 03 19 03 54	Set b m 13 29 14 31 15 32 16 33 17 34	DECI Rise h m 01 20 01 54 02 30 03 10 03 54	2/BER Set h m 14 26 15 26 16 25 17 24 18 20
6 7 8 9	04 16 1	12 05 12 47 13 37 14 35 15 42	03 05 04 07 05 02 05 51 06 33	13 22 14 30 15 41 16 53 18 03	04 26 05 05 05 41 06 15 06 49	15 42 16 49 17 54 18 58 20 00	04 12 04 46 05 20 05 57 06 36	16 41 17 44 18 46 19 47 20 47	04 32 05 13 05 59 06 48 07 41	18 34 19 33 20 29 21 20 22 07	04 41 05 33 06 27 07 21 08 16	19 13 20 02 20 45 21 24 21 58
11 12 13 14 15	07 16 1 08 01 1	16 54 18 07 19 17 20 25 21 29	07 11 07 46 08 20 08 53 09 29	19 10 20 14 21 16 22 17 23 17	07 25 08 02 08 42 09 26 10 14	21 02 22 03 23 01 23 57	07 20 08 06 08 57 09 50 10 45	21 45 22 39 23 29 00 13	08 35 09 30 10 25 11 19 12 13	22 48 23 26 23 59 00 30	09 10 10 04 10 58 11 52 12 48	22 30 23 00 23 29 23 59
16 17 18 19 20	10 22 2 10 55 11 30 0	22 31 23 31 00 30 01 28	10 06 10 47 11 32 12 21 13 13	00 16 01 13 02 06 02 56	11 06 11 59 12 54 13 50 14 46	00 49 01 36 02 18 02 57 03 31	11 40 12 35 13 30 14 25 15 22	00 53 01 28 02 01 02 32 03 03	13 08 14 04 15 03 16 05 17 10	01 00 01 30 02 01 02 35 03 13	13 47 14 49 15 54 17 02 18 08	00 30 01 05 01 45 02 32 03 27
21 22 23 24 25	13 36 0 14 26 0 15 20 0	02 25 03 20 04 12 05 00 05 44	14 08 15 03 15 59 16 55 17 50	03 41 04 22 04 59 05 32 06 03	15 41 16 37 17 35 18 34 19 35	04 03 04 34 05 04 05 35 06 08	16 20 17 21 18 25 19 31 20 37	03 33 04 06 04 42 05 22 06 08	18 18 19 26 20 30 21 27 22 17	03 57 04 49 05 48 06 53 08 03	19 11 20 06 20 54 21 36 22 14	04 31 05 40 06 52 08 03 09 11
26 27 28 29 30	18 06 0	06 23 06 58 07 30	18 46 19 43 20 42	06 33 07 03 07 34	20 38 21 43 22 48 23 50	06 45 07 26 08 13 09 08	21 42 22 41 23 34	07 02 08 02 09 07	23 01 23 39	09 12 10 19 11 23 12 26	22 48 23 22 23 56	10 17 11 19 12 20 13 20

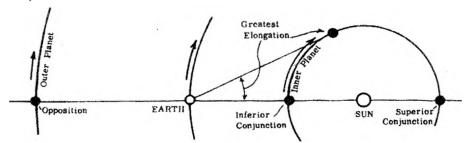
THE PLANETS

BAS	IC	DATA

	Dist from Sun 10 ⁵ km	Period of Revolution years	Nass (Earth = 1)	Equatorial Diameter 10 ³ km	Rotation Period	Inclination of Equator to orbit	No. of known satellites
Mercury	58	0.24	0.055	4.98	58.65d	0,	0
Venus	108	0.62	0.815	12.10	243d R	178°	0
Earth	150	1.00	1.000	12.76	23h56≡	23°27′	1
Kars	228	1.88	0.107	6.79	24h37m	23 59 '	2
Jupiter	778	11.9	318.867	142.80	091:51≡	03°04'	16
Saturn	1 426	29.5	95.142	120.00	10h14m	26 44'	17
Uranus	2 868	84.0	14.559	52.00	17.2h	97*52'	15
Neptune	4 494	164.8	17.207	48.40	17.8h	29*347	8
Pluto	5 896	247.6	0.002	3.00	6.39d	118*?	1

GENERAL

Apart from Uranus, Neptune and Pluto, the planets of our solar system are amongst the brightest objects in the night sky. Their apparent brightness is measured in magnitudes. A planet of magnitude 1.0, that of the brightest stars, will be 100 times brighter than one of magnitude 6.0, the limit of visibility to the naked eye in the total absence of artificial lighting. Unlike the distant stars, the relative positions of the planets do not remain fixed, but continually change as, like the Earth, they orbit around the Sun. Their apparent movements against the starry background are complicated as they result from a combination of their own motion and the Earth's motion. Their brightnesses also vary considerably, as both their distances from the Earth and the visible portions of their sunlit hemispheres change. Since the period of a planet increases with increasing distance from the Sun, so we find that the inner planets - Mercury and Venus - appear to "overtake" the Earth in their orbits, while the Earth in turn "overtakes" the outer planets - Mars, Jupiter and Saturn. The terms given in astronomy to the various Sun-Earth-Planet configurations are illustrated in the accompanying diagram. Dates of such configurations occuring in 1991 are listed chronologically in the Diary (pages 4 and 5) and are also mentioned in the text below.

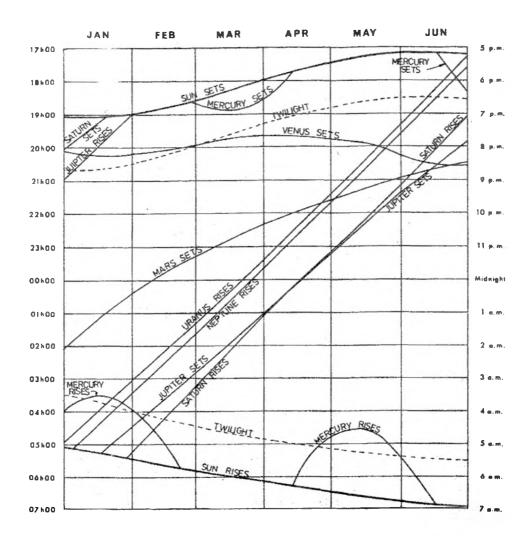


OBSERVING THE PLANETS

To the naked eye, planets appear as virtually point sources of light. However, their disks can be readily resolved with the aid of a small telescope. Even so, their angular diameters are of the order of 10 seconds of arc - roughly 1/200 of the Moon's angular diameter - so it is not always possible to distinguish details on their disks. The disks of Mercury and Venus are only seen fully illuminated when they are furthest from us - as they draw closer, their disks grow larger but the phase changes to a crescent as we see more of their dark hemispheres. In contrast, the disks of the outer planets are always seen fully or near fully illuminated.

TIMES OF RISING AND SETTING

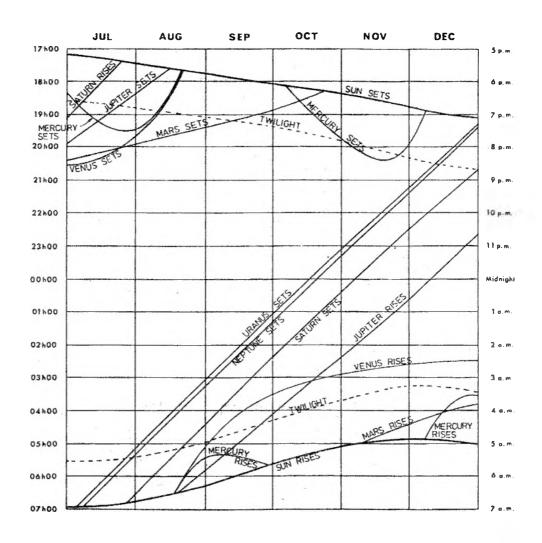
The times of rising and setting given by the diagram are accurate for position 30° East, 30° South and approximately correct for other places in Southern Africa. Strictly speaking, corrections for latitude and longitude should be applied, but the latitude correction is in general sufficiently small to be ignored and in no case will exceed 15 minutes.



CORRECTION FOR PLACES NOT ON THE 30° E MERIDIAN

Approximate longitude corrections from the 30° East meridian are:

Bloemfontein	+15 ^m	East London	+8 ^m	Port Elizabeth	+18 ^m
Bulawayo	+6 ^m	Grahamstown	+14 ^m	Pretoria	
Cape Town	+46 ^m	Johannesburg	+8 ^m	Harare	-4 ^m
Durban	-4 ^m	Kimberley	+21 ^m	Windhoek	+52 ^m



MERCURY

The planet may be seen low in the east before sunrise between the following approximate dates:

January 1 (at mag.+1.0) to February 18 (at mag.-0.8), April 23 (at mag.+3.0) to June 10 (at mag.-1.4), August 30 (at mag.+2.0) to September 23 (at mag.-1.3) and December 14 (at mag.+1.8) to 31 (at mag.-0.3).

The best conditions for viewing will occur in January and early February when Mercury is in Sagitarius and from the end of April until the end of May when it is in Pices and later in Aries.

Mercury may also be seen low in the west after sunset between the following approximate dates:

March 12 (at mag.-1.4) to April 6 (at mag.+2.2), June 25 (at mag.-1.3) to August 15 (at mag.+2.9) and October 18 (at mag.-0.7) to December 3 (at mag.+1.8)

The best conditions for viewing will be from early July until just before mid-August, when Mercury passes through Cancer and into Leo, and again from late October, when Mercury is in Libra, to late November as it moves through Scorpio to Ophiuchus.

		d	h			đ	h			d	h		(1	h	
Superior																
Conjunction					Mar	2	. 2		Jun	17	7		Oct	3	19	
Greatest																
Elongation East					Mar	27	17	(19°)	Jul	25	4	(27°)	Nov	19	4	(22°)
Stationary					Apr	4	19		Aug	7	6		Nov	28	20	
Inferior																
Conjunction					Apr	14	23		Aug	21	23		Dec	8	17	
Stationary	Jan	3	21		Apr	27	8		Aug	30	22		Dec	18	14	
Greatest									_							
Elongation West	Jan	14	11	(24°)	May	12	20	(26°)	Sep	7	20	(18°)	Dec	27	23	(22°)

VENUS

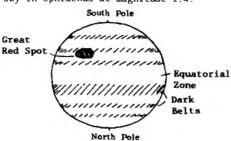
Venus will be in the evening sky from January (at mag. -3.9) to mid-August (at mag.-4.2). It will reach greatest brilliancy on July 17 at magnitude -4.5. It will be in the morning sky from the end of August (at mag.-4.0) untill the end of the year (at mag.-4.1). It will reach greatest brilliancy on September 29 at magnitude -4.6.

MARS

Mars, visible for most of the night, begins the year in the constellation of Taurus (at mag.-1.0) passing in early April (at mag. 1.0) to Gemini and in late May (at mag. 1.6) to Cancer having become an evening sky object. It passes into Leo towards the end of June and into Virgo at the end of August (at mag. 1.8) where by the end of September it will be too close to the Sun to be seen. Late December will see its return to the morning sky in Ophiuchus at magnitude 1.4.

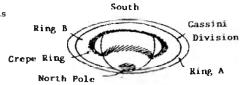
JUPITER

Jupiter begins the year in Cancer (at mag. -2.5) where it is visible for most of the night. It will be an evening sky object by June (at mag. -1.8). It passes to Leo in early July becoming too close to the Sun in early August. Jupiter re-appears in the morning sky at the beginning of September (at mag. -1.7) still in Leo where it remains for the rest of the year. It can be seen for more than half the night by mid-December (at mag. -2.2)



SATURN

Saturn will be seen in Sagitarius as an the evening sky object on 1 January but will shortly there after not be seen until early February when it will be in the morning sky having passed into Capricornus where it remains for

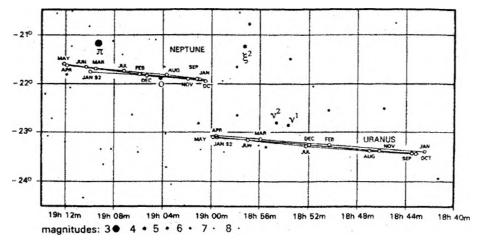


the rest of the year. It will be an all night object by July and an evening sky object from late October. Its magnitude will be +0.6 at the beginning of February increasing to +0.1 in late July and decreasing to +0.7 by the end of the year.

URANUS AND NEPTUNE

Uranus and Neptune are both in Sagitarius all year. Uranus is at magnitude 5.6 at opposition on July 4 and Neptune is at magnitude +7.9. at opposition on July 8. The paths of Uranus and Neptune.

Neptune passes south of omicron SGR on January 20, 26 July and 22 November.



PLITTO

Pluto at magnitude +14 in Serpens is an object for a large telescope.

EVENTS OF INTEREST

Eveni	ng	Sky	:
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1 Jan. Venus (the brighter object) in conjunction with Saturn.

mid-Mar. Mercury and Venus visible. Jun.

Venus, Mars and Jupiter visible.

14 Jun. Mars and Jupiter (the brighter object) in conjunction.

18 Jun. Venus and Jupiter in conjunction.

23 Jun. Venus and Mars in conjunction.

Mercury, Venus, Mars and Jupiter visible. Jul. to early Aug

15 Jul. Jupiter and Mercury in conjunction.

22 Jul. Venus and Mars in conjunction.

early to mid-Aug. Mercury, Venus and Mars visible.

7 Aug. Venus and Mercury in conjunction.

Morning Sky:

early Feb. Mercury (the brighter object) and Saturn visible.

5 Feb. Mercury and Saturn in conjunction.

end Aug. to mid.Sep. Mercury, Venus and Jupiter visible. 29 Aug.

Mercury and Venus in conjunction. 10 Sep. Mercury and Jupiter (the brighter object) in conjunction.

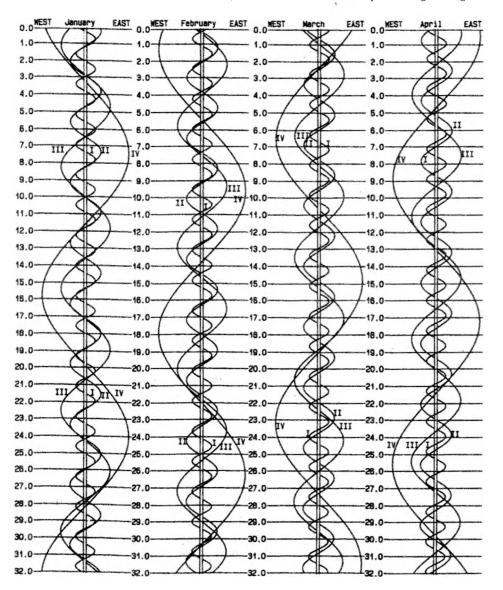
Oct. Venus and Jupiter visible.

17 Oct. Venus and Jupiter in conjunction. mid to end Dec. Mercury, Venus and Mars visible.

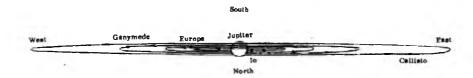
APPARENT	Mercury	Venus	Mars	Jupiter
	RA DEC	RA DEC	RA DEC	RA DEC
Jan 1 Jan 11 Jan 21 Jan 31 Feb 10 Feb 20	h m ° ', 17 35.8 -20 11 17 47.0 -21 21 18 33.1 -22 43 19 32.2 -22 34 20 36.6 -20 22 21 43.4 -15 55	h m ° , 19 47.6 -22 27 20 40.2 -19 53 21 30.8 -16 22 22 19:2 -12 7 23 57 -7 19 23 51.0 -2 12	h m 6 7 6 7 6 7 7 8 7 9 7 9 7 9 7 9 7 9 7 9 7 9 7 9 7	h m ' ' ' ' ' ' ' ' ' ' ' ' ' ' ' ' ' '
Mar 2	22 51.9 -9 11	0 35.7 3 1	4 50.0 24 39	8 29.5 19 50
Mar 12	0 1.3 1 30	1 20.4 8 8	5 10.6 25 3	8 26.5 20 0
Mar 22	1 4.3 8 21	2 5.7 12 57	5 32.4 25 18	8 24.7 20 6
Apr 1	1 40.2 13 41	2 52.3 17 15	5 55.3 25 22	8 24.2 20 7
Apr 11	1 36.1 13 2	3 40.2 20 52	6 18.8 25 14	8 25.1 20 4
Apr 21	1 13.9 8 17	4 29.5 23 36	6 43.0 24 51	8 27.1 19 56
May 1	1 10.4 5 18	5 19.6 25 19	7 7.4 24 15	8 30.4 19 45
May 11	1 33.1 6 21	6 9.6 25 55	7 32.0 23 24	8 34.7 19 29
May 21	2 15.4 10 22	6 58.6 25 55	7 56.6 22 19	8 39.9 19 9
May 31	3 14.8 16 2	7 44.7 23 55	8 21.1 21 0	8 45.9 18 46
Jun 10	4 33.3 21 43	8 27.4 21 32	8 45.4 19 28	8 52.6 18 19
Jun 20	6 7.3 24 47	9 5.7 18 30	9 9.5 17 44	8 59.8 17 50
Jun 30	7 37.3 23 30	9 38.6 15 4	9 33.3 15 49	9 7.4 17 17
Jul 10	8 49.3 19 10	10 5.0 .11 29	9 57.0 13 43	9 15.5 16 42
Jul 20	9 42.2 13 42	10 23.0 8 4	10 20.4 11 29	9 23.7 16 4
Jul 30	10 15.6 8 37	10 29.9 5 15	10 43.7 9 8	9 32.1 15 24
Aug 9	10 24.7 5 38	10 22.9 3 35	11 6.9 6 40	9 40.7 14 43
Aug 19	10 4.6 6 47	10 3.3 3 34	11 30.1 4 7	9 49.2 14 0
Aug 29	9 39.7 11 7	9 39.9 5 5	11 53.5 1 31	9 57.7 13 16
Sep 8	9 56.6 12 45	9 26.0 7 6	12 17.0 -1 8	10 6.0 12 32
Sep 18	10 54.2 8 51	9 27.3 8 37	12 40.8 -3 47	10 14 2 11 48
Sep 28	12 1.7 1 38	9 42.4 9 10	13 5.0 -6 25	10 22:0 11 5
Oct 8	13 5.6 -6 5	10 7.4 8 39	13 29.8 -8 59	10 29 5 10 23
Oct 18	14 6.1 -13 5	10 38.9 7 8	13 55.2 -11 29	10 36.6 9 43
Oct 28	15 5.4 -18 52	11 14.4 4 46	14 21.3 -13 52	10 43.1 9 7
Nov 7	16 4.3 -23 6	11 52.6 1 43	14 48.3 -16 6	10 48.9 8 33
Nov 17	16 59.1 -25 21	12 32.8 -1 51	15 16.2 -18 8	10 54.1 8 4
Nov 27	17 33.3 -25 8	13 14.5 -5 41	15 45.0 -19 56	10 58.4 7 40
Dec 7	17 9.2 -21 58	13 58.0 -9 36	16 14.8 -21 27	11 1.7 7 23
Dec 17	16 27.3 -18 48	14 43.2 -13 21	16 45.4 -22 39	11 3.9 7 11
Dec 27	16 45.7 -20 12	15 30.5 -16 42	17 16.9 -23 29	11 5.1 7 7
	Saturn RA DEC	Uranus RA DEC	Neptune RA DEC	RA DEC
Jan 1 Jan 15 Feb 1 Feb 15	h m '9 19 50.7 -21 9 19 57.6 -20 51 20 6.1 -20 28 20 12.8 -20 9	h m 18 42.4 -23 24 18 46.1 -23 20 18 50.2 -23 16 18 53.3 -23 12	h m 19 1.0 -21 54 19 3.2 -21 51 19 5.9 -21 47 19 7.8 -21 44	15 24.6 -3 13 15 26.0 -3 13 15 27.2 -3 11 15 27.8 -3 6
Mar 1	20 19.1 -19 50	18 55.9 -23 9	19 9.5 -21 41	15 27.9 -3 0
Mar 15	20 24.7 -19 32	18 58.0 -23 7	19 10.9 -21 39	15 27.6 -2 52
Apr 1	20 30.3 -19 14	18 59.6 -23 5	19 11.9 -21 36	15 26.7 -2 43
Apr 15	20 33.8 -19 3	19 0.2 -23 5	19 12.3 -21 36	15 25.6 -2 35
May 1	20 36.4 -18 56	18 59.9 -23 5	19 12.1 -21 36	15 24.0 -2 27
May 15	20 37.3 -18 54	18 58.9 -23 7	19 11.5 -21 36	15 22.6 -2 21
Jun 1	20 36.6 -18 58	18 57.0 -23 10	19 10.3 -21 38	15 20.8 -2 17
Jun 15	20 34.7 -19 7	18 54.9 -23 13	19 8.9 -21 41	15 19.5 -2 16
Jul 1	20 31.2 -19 22	18 52.2 -23 17	19 7.2 -21 43	15 18.3 -2 18
Jul 15	20 27.4 -19 37	18 49.7 -23 20	19 5.6 -21 46	15 17.6 -2 23
Aug 1	20 22.2 -19 56	18 47.0 -23 23	19 3.7 -21 49	15 17.2 -2 31
Aug 15	20 18.1 -20 10	18 45.1 -23 25	19 2.3 -21 52	15 17.4 -2 40
Sep 1	20 13.8 -20 25	18 43.5 -23 27	19 1.1 -21 54	15 18.1 -2 52
Sep 15	20 11.4 -20 33	18 43.0 -23 27	19 0.6 -21 55	15 19.2 -3 4
Oct 1	20 10.1 -20 38	18 43.2 -23 27	19 0.4 -21 56	15 20.8 -3 18
Oct 15	20 10.4 -20 38	18 44.1 -23 26	19 0.8 -21 56	15 22.5 -3 29
Nov 1	20 12.5 -20 31	18 46.2 -23 23	19 1.9 -21 55	15 24.9 -3 43
Nov 15	20 15.7 -20 21	18 48.6 -23 20	19 3.3 -21 53	15 27.0 -3 52
Dec 1	20 20.7 -20 6	18 52.0 -23 16	19 5.3 -21 51	15 29.4 -4 1
Dec 15	20 26.1 -19 48	18 55.3 -23 12	19 7.3 -21 48	15 31.4 -4 7
Dec 31	20 33.1 -19 24	18 59.3 -23 6	19 9.8 -21 44	15 33.5 -4 11

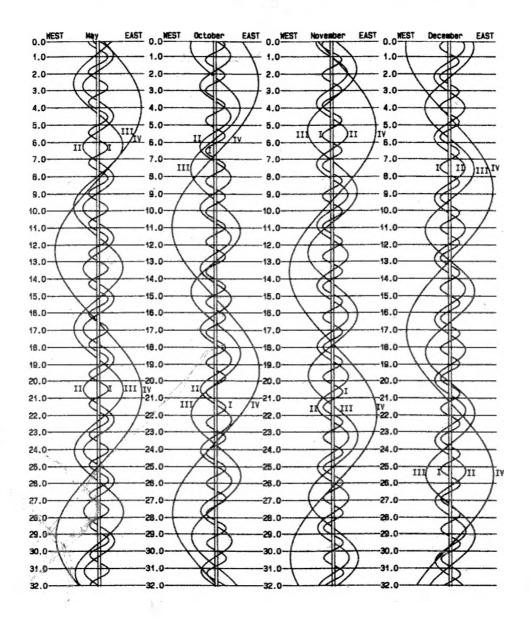
THE MOONS OF JUPITER

One of the most popular sights for an observer with a small telescope is Jupiter and its moons. Four of the sixteen - Io, Europa, Ganymede and Callisto - are generally clearly visible - they would just be visible to the naked eye were it not for the glare from the mother planet. As the diagram on page 22 indicates, the system is seen almost edge-on so the moons always lie close to a straight line extending from the planet's equator. As they orbit, so they appear to oscillate from one side to the other, alternately passing in front and behind the planet. This motion is represented in the following diagrams which show how their positions along such a straight line change during the 8 months when Jupiter is prominent. For each month, time increases downward; the disc of Jupiter is stretched to make the central column, and horizontal lines representing midnight



(0 am SAST), are shown for every day of the month. The wavy lines show how the moons appear to ossilate from each side of the planet to the other.





When the moons pass in front and behind the planet, transits, occultations and eclipses occur. Details of such phenomena, occurring between the end of astronomical twilight in the evening and its commencement in the morning when the planet is above the horizon in Southern Africa, are given in the table below.

EXPLANATION OF THE TABLE.

20 23 II.Sh.I.

- Date and predicted times are given; these are for mid-phenomean and are not instantaneous.
- The moon concerned are I Io, II Europa, III Ganymede and IV Gallisto.
- Phenomena the abbreviations used are D Disappearance; Ec Eclipse ie.the satellite passes through the shadow of Jupiter; R Reappearance;
 Oc Occultation ie.the satellite is obscured by the disc of Jupiter;
 I Ingress; Sh Shadow Transit ie.the shadow of the satellite transits the

disc; E - Egress; Tr - Transit ie the satellite crosses the disc of Jupiter.

20 47 I.Tr.E.

14 19 24 I.Sh.E.

23 19 IV.Oc.D.

	4 6 -	4 > -	4 5 -	4.5.
Man	d h m	d h m	d B B	d h n
MAL	14 20 24 II.Oc.D			Nov 21 04 22 I.Tr.E.
	15 01 12 II.Ec.R		30 19 22 I.Ec.R.	04 56 III.Ec.D.
	16 20 13 II.Sh.E			24 03 26 II.fr.I.
	18 19 20 III.Tr.I		04 19 07 II.Ec.R.	03 47 II.Sh.E.
	22 54 III.Tr.E		05 18 44 III.Tr.E.	25 03 19 III.Tr.E.
	23.26 IV.Tr.I		19 24 III.Sh.I.	26 02 43 IV.Tr.I.
	23 27 III.Sh.I		20 36 I.Tr.I.	28 02 48 I.Sh.I.
	19 02 22 I.Oc.D.		06 21 17 I.Ec.R.	04 01 I.Tr.I.
	23 32 I.Tr.I		07 18 25 I.Sh.E.	05 04 I.Sh.E.
	20 00 33 I.Sh.I.		12 19 26 III.Tr.I.	29 03 26 I.Oc.R.
	01 48 I.Tr.E	-	13 19 56 I.Oc.D.	Dec 01 03 33 II.Sh.I.
	20 49 I.Oc.D.		14 19 23 I.Tr.E.	02 02 23 III.Sh.E.
	21 00 11 I.Ec.R	. 23 02 III.Sh.E.	20 20 I.Sh.E.	03 52 III.Tr.I.
	19 02 I.Sh.I.		18 19 35 II.Oc.D.	03 03 58 II.Oc.R.
	20 16 I.Tr.E	. 20 01 II.Tr.E.	20 19 22 II.Sh.E.	04 03 56 IV.Ec.R.
	21 19 I Sh.E.	22 35 II.Sb.E.	21 19 05 I.Tr.I.	05 04 41 I.Sb.I.
	22 48 II.Oc.D	. 27 21 50 I.Tr.I.	19 57 I.Sb.I.	06 01 48 I.Ec.D.
	23 19 57 II.Sb.I.	23 06 I.Sh.I.	22 19 35 I.Ec.R.	07 01 26 I.Sh.E.
	20 38 II.Tr.E	. 28 19 09 I.Oc.D.	27 18 31 IV.Tr.E.	02 39 I.Tr.E.
	22 50 II.Sh.E.	22 45 I.Ec.R.	19 05 II.Sh.I.	09 02 48 III.Sh.I.
	25 22 58 III.Tr.I	. 29 18 35 I.Tr.E.	20 23 II.Tr.E.	10 01 15 II.Ec.D.
	27 01 22 I.fr.I.	. 19 51 I.Sh.E.	29 18 26 I.Oc.D.	12 00 38 II.Tr.E.
	21 49 IV.Ec.R	21 38 IV.Oc.R.		13 00 29 IV.Tr.E.
	22 39 I.Oc.D	. 30 21 52 III.Tr.E.	18 39 I.Sh.E.	01 21 III.Oc.R.
	28 19 49 I.Tr.I		Jul 06 18 53 II.Ec.R.	
	20 57 I.Sb.I			14 01 03 I.Sh.I.
	22 06 I.Tr.E	•	11 18 58 III.Sh.E.	02 16 I. Tr .I.
	23 13 I.Sh.E		Sep 27 04 51 I.Tr.I.	03 19 I.Sh.E.
	29 01 13 II.Oc.D	•	28 05 46 II.Tr.I.	04 31 I.Tr.E.
	20 35 I.Ec.R		05 48 IV-Oc.D.	
	21 15 III.Ec.R		Oct 09 05 10 III.Ec.D.	
	30 20 14 II.Tr.I		12 05 11 I.Ec.D.	19 00 26 II.Tr.I.
	22 34 II.Sh.I		13 04 47 I.Sh.E.	00 46 II.Sh.E.
	23 06 II.Tr.E		05 38 I.Tr.E.	03 10 II.Tr.E.
	31 01 27 II.Sh.E	_	14 04 33 II.Ec.D.	20 00 21 III.Ec.R.
Anr	01 19 41 II.Ec.R.		20 04 23 I.Sh.I.	01 43 III.Oc.D.
uhr	04 00 31 I.Oc.D			05 11 III.0c.R.
	20 30 IV.Tr.E		21 04 47 I.Oc.R.	21 02 56 I.Sh.I.
	21 41 I.Tr.I		23 03 23 II.Tr.I.	04 07 I.Tr.I.
	22 52 I.Sb.I		04 15 II.Sh.E.	05 12 I.Sh.E.
	23 57 I.Tr.E		28 03 26 I.Ec.D.	05 24 IV.Oc.D. V
	05 01 08 I.Sh.E		29 03 02 I.Sh.E.	22 00 02 I.Ec.D.
	18 59 I.Oc.D		04 04 I.Tr.E.	03 31 I.Oc.R.
	20 18 III.Oc.R		30 04 00 II.Sh.I.	23 00 50 I.Tr.E.
	21 36 III.Ec.D		Nov 01 04 03 II.Oc.R.	26 00 33 II.Sh.I.
	22 30 I.Ec.R		04 05 19 I.Ec.D.	02 56 II.Tr.I.
	06 19 37 I.Sh.E			03 21 II.Sh.E.
				27 00 43 III.Ec.D.
	22 44 II.Tr.I 08 22 16 II.Ec.R		03 45 I.Tr.I. 04 56 I.Sh.E.	04 18 III.Ec.R.
	11 23 33 I.fr.I		06 03 11 I.Oc.R.	28 00 46 II.Oc.R.
	12 00 47 I.Sh.I		07 05 08 III.Oc.R.	04 49 I.Sh.I.
	20 31 III.0c.D			
	20 52 I.Oc.D		13 05 07 I.Oc.R.	02 58 IV.Sb.I.
	23 01 IV.Oc.D			05 22 I.Oc.R.
	13 00 08 III.Oc.R		04 35 III.Ec.R.	23 17 I.Sh.I.
	00 25 I.Ec.R		15 04 13 II.Ec.D.	30 00 25 I.Tr.I.
	19 15 I.Sh.I		17 03 33 II.Tr.E.	01 33 I.Sh.E.
	20 18 I.Tr.E		20 03 33 I.Ec.D.	02 40 I.Tr.E.
	21 32 I.Sb.E		21 02 07 I.Tr.I.	23 49 I.Qc.R.
	14 18 54 I.Ec.R	. 19 43 I.Sh.I.	03 11 I.Sh.E.	

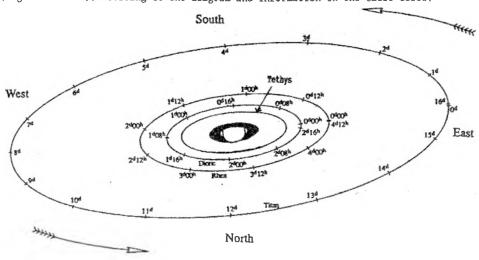
MUTUAL PHENOMENA OF JUPITER'S SATILITES

The table below gives predictions of observable mutual eclipses and occultations of the Galilean Satilites occuring in 1991 conforming to same conditions as stated on page 23 and taking place more than 2.0 Jovian radii from the planet's centre.

Date	Begin	End	Event	J Dist
d	hш	h m		radii
Jan 02	00 45.3	01 01.3	II Ec. III	10.5
Jan 02	03 30.2	03 50.2	II Oc. III	9.3
Jan 05	02 09.5	02 28.1	II Oc. I	5.5
Jan 09	04 52.3	05 14.0	II Ec. III	10.2
Jan 12	03 40.1	04 05.1	II Ec. I	5.4
Jan 12	04 58.1	05 11.4	II Oc. I	5.8
Jan 22	03 32.4	04 01.7	II Oc. I	5.0
Jan 22	20 14.7	20 29.4	II Ec. I	5.9
Jan 22	20 37.2	20 48.2	II Oc. I	5.9
Jan 23	17 20.9	20 13.9	II Oc. III	6.9
Jan 23	21 49.7	22 50.4	II Ec. III	5,2
Jan 29	22 54.5	23 04.5	II Oc. I	5.9
Jan 29	22 56.2	23 08,5	II Ec. I	5.9
Feb 06	01 07.9	01 17.3	II Oc. I	5.9
Feb 06	01 30.3	01 40.6	II Ec. I	5.9
Feb 13	03 19.1	03 27.7	II Oc. I	5.8
Feb 13	03 59.6	04 08.9	II Ec. I	5.7
Feb 20	05 29.4	05 37.4	II Oc. I	5.8
Feb 21	03 05.4	05 27.0	III Oc. IV	8.7
Feb 23	19 38.0	19 46.3	II Ec. I	5.5
Mar 02	20 44.5	20 51.5	II Oc. I	5.7
Mar 02	22 00.8	22 08.8	II Ec. I	5.2
Mar 09	22 54.6	23 00.9	II Oc. I	5.5
Mar 10	00 22.1	00 29.4	II Ec. I	5.0
Mar 17	01 05.7	01 11.1	II Oc. I	5.4
Mar 17	23 49.7	04 14.4	IV Oc. III	9.7
Mar 18	21 37.7	21 47.4	IV Ec. II	6.7
Mar 25	21 38.8	21 45.2	III Ec. II	5.4
Apr 02	00 52.5	01 00.1	III Ec. II	5.9
Apr 03	20 28.3	20 34.3	II Ec. I	4.0
Apr 10	20 51.1	20 54.4	II Oc. I	4.9
Apr 10	22 45.6	22 51.2	II Ec. I	3.8
Apr 13	23 11.1	23 18.4	II Ec. IV	8.8
Apr 21	20 48.3	20 56.3	IV Ec. I	2.1
Apr 29	23 08.8	23 14.1	I Ec. II	2.0
Apr 30	19 37.8	19 43.8	III Ec. I	3.9
May 07	22 25.4	22 32.4	III Ec. I	4.5
May 12	19 06.6	19 08.9	II Oc. I	4.0
May 14	20 30.5	20 37.5	III Ec. II	7.8
May 19	21 25.2	21 27.9	II Oc. I	3.8
	19 02.7	19 08.3	I Ec. II	2.9
,		21 24.8	I Ec. II	3.2
,	21 19.1 19 25.1	19 33.4	II Ec. III	6.5
Jun 09				5.3
Jun 11	19 01.4	19 11.7 20 59.7	IV Ec. I	4.2
Jun 19	19 37.7			2.6
Jun 20	19 55.5	19 59.5		5.5
Jun 26	19 30.1	19 53.5		4.8
Jun 27	18 38.6	18 48.6		5.2
Jul 02	18 31.4	18 36.4		
Jul 02	19 42.1	19 46.8	I Ec. II	4.5

THE MOONS OF SATURN

Saturn's moons are considerably fainter than the 4 Galilean moons of Jupiter. The diagram shows the orbits of 4 of Saturn's moons. The easiest to find is Titan (magnitude +8.5), according to the diagram and information in the table below.



TITAN											
Eastern	rn Elongation Infe		Inferior	Conjunction		Western	Elongation		Superior	Conjunction	
	đ	h		d	h		d	h		d	h
Jan	15	00.0		19	01.6		23	06.6		27	05.2
	31	00.7	Feb	4	02.5	Feb	8	07.4	Feb	12	05.8
Feb	16	01.3		20	03.2		24	08.0		28	06.2
Mar	4	01.8	Mar	8	03.7	Mar	12	08.4	Mar	16	06.5
	20	02.0		24	04.0		28	08.6	Apr	1	06.4
Apr	5	01.9	Apr	9	03.9	Apr	13	08.3		17	06.0
	21	01.4		25	03.4		29	07.7	May	3	05.2
May	7	00.6	May	11	02.5	May	15	06.7	Ť	19	04.1
	22	23.3		27	01.1		31	05.2	Jun	4	02.5
Jun	7	21.6	Jun	11	23.3	Jun	16	03.3		20	00.5
	23	19.6		27	21.1	Jul	2	01.0	Jul	5	22.3
Jul	9	17.2	Jul	13	18.6		17	22.5		21	19.9
	25	14.8		29	16.0	Aug	2	20.0	Aug	6	17.4
Aug	10	12.2	Aug	14	13,4	_	18	17.5	_	22	15.0
	26	09.9		30	11.0	Sep	3	15.2	Sep	7	12.9
Sep	11	07.8	Sep	15	09.0		19	13.2		23	11.1
	27	06.1	0ct	1	07.3	0ct	5	11.7	0ct	9	09.7
Oct	13	04.8		17	06.2		21	10.7		25	08.8
	29	04.0	Nov	2	05.5	Nov	6	10.1	Nov	10	08.3
Nov	14	03.6		18	05.3		22	10.0		26	08.1
	30	03.5	Dec	4	05.4	Dec	8	10.2	Dec	12	08.2
Dec	16	03.8		20	05.9		24	10.6		28	08.6

COMETS AND METEORS

COMETS

A typical comet consists of a solid nucleous surrounded by a very large envelope of gas and dust, called the coma; in some instances, comets might develop a tail. Depending on the length of their periods around the Sun, comets are classed into 2 groups: Short-period comets, with an average of a 7 year period, an orbital inclination of about 13 degrees and small orbital eccentricities, (0.2 to 0.9) and usually travel in a direct motion; Long-period comets have a period greater than 200 years, random orbital inclinations and about 0.9999 eccentricity and random motion

About a dozen comets are observed each year and of these, 5 or 6 are new discoveries. As a comet approaches or moves away from the Sun its visual appearance changes drastically and it might become bright enough to be observed with modest equipment such as a good pair of binoculars or a telescope with an aperture not less than 75mm; on rare occasions a very bright comet may become visible to the naked-eye.

Interested observers can contribute with valuable information by reporting on their visual appearance such as coma magnitude, apparent diameter and degree of condensation; if the comet sports a tail, its length in degrees and its p.a. (position angle) should also be reported, as well as the type of instrument and magnification used.

Another area much in need of observers, is that of regularly conducting visual searches of possible new comets, namely in the southern celestial hemisphere. Interested members are asked to contact the Director of the Comet and Meteor Section:

Jose Campos, 19 Fiskaal Place, Woodhaven, Durban 4001.

METEORS

Orbiting within the solar system, METEOROIDS often collide with the Earth's atmosphere and as a result, they heat to incandescence by friction with the atmosphere's gaseous molecules and an emission of light takes place due to ionization; while in flight through the atmosphere they are called METEORS and they disintegrate completely. There are two types of meteors: Sporadic ones which may be seen at any time of the night, anywhere in the sky; the shower meteors are associated with known meteor showers that are active during certain times of the year (See the Table on the next page). Bright meteors of visual magnitude equal or greater than that of Venus (-4.0 mag.), are classed as Fireballs and if they explode while in flight, they are termed Bolides. Fireballs of visual magnitude equal or brighter than that of the Full Moon are known to produce debris that when found on the ground are called METEORITES. Interested persons are asked to contact the Director:

Jose Campos, 19 Fiskaal Place, Woodhaven, Durban 4001. In the event of Fireballs, please phone 031-423684 at any time.

				Recommended watch											
	dicted		Neteor	Radia		Da	te	Trans.					Nar.		٧
Li	∎its		Shower	(195	0)	8	t	Radia	int	at	I.H.R.	Beginni			
				R.A.				SAST	Alt	Maximum			t s	SAST Alt	
Jan	06-28		Alpha Crucids	12h48				04 h 56	57	Very good	?	2 2h3 0 2	25 (Mh00 55	50
Jan	23-Mar	12	Theta Centaurids	14h00	-40	?eb	08	05ხ05	76	Pavourable	?	00b00 3	1 0	14h00 70	60
Mar	01-12	?	Pyxids(new)	09b00	-35	Mar	06?	22b 02	85	Favourable	6?	20b00 6	3 (3h15 26	
Mar	13-18		Corona Australids	16b20	-48	Mar	16	04b43	72	New moon	5	23b15 2	6 0	4 b 30 71	
Mar	21-Apr	80	Delta Pavonids	20b10	-65	Mar	29?	07 b43	55	Unfavourable	7	02b00 2	8 0	Mh00 41	
λpr	16-25		April Lyrids	18h05	+34	Apr	22	04b06	28	Good	15	03h00 2	16 C	5b00 26	49
Apr	11-Kay	12	Alpha Scorpids	16b00	-22	Hay	03	01h16	82	Pavourable	7	21h00 3	3 (Mh30 46	
Apr	20-Jul	30	Sco-Sgr System	18h00	-30	Jun	14	00ћ29	90	Good	10	21h00 4	5 0	4b00 44	30
Apr	19-May	28	Eta Aquarids	22h24	-01	Nay	03	07h23	60	Good	30	04b00 3	1 0	5b00 43	
May	25 -Ju n	20	Chi Scorpids	16h28	-13	Jun	05	23b30	73	Good	6?	19h00 2	5 0	Mh10 25	
Jun	08-16		Sagittarids	20h16	-35	Jun	11	02h59	85	Very good	4?	21145 2	6 0)5h10 62	
Jun	08-16		Theta Ophiuchids	17h48	-28	Jun	13	00b23	9.8	Very good	5?	20b00 3	4 0	5b00 30	
Jun	10-21		June Lyrids	18h32	+35	Jun	16	00b53	25	Pavourable	5	23h30 2	2 0	1h30 24	31
Jun	17-26		Ophiuchids	17020	-20	Jun	20	23 <u>h</u> 22	80	Favourable	10	19b00 3	0 0	2b00 53	
Jun	26-29		Cetids (new)	02h00	-15	Jun	28	07 b 35	75	Poor	2	03b00 2	5 0	/5 <u>h2</u> 0 56	
Jul	10-Aug	05	Capricornids	21h00	-15	Jul	26	00b47	75	Full moon	8	-		-	
			North Delta Aquarids			Aug	12	01 b 13	65	Good	5	21h00 2	6 0	4h30 37	42
			Alpha Capricornids	20h36	-10	Jul	30	23141	70	Poor	8	20b00 3	3 0	4 b 00 29	23
Jul	09-Aug	17	Piscis Australids	22h40	-30	<u>jaj</u>	29	0 2b 03	89	Poor	12	21b30 3	2 0	5h00 51	35
			South Iota Aquarids	22h20	-15	Aug	05	01h 29	75	Favourable	3	22b00 4		4b00 52	34
			North Iota Aquarids	21h48		Àug	20	23 b 51		Unfavourable		20b00 3	1 0	4h00 28	31
Jul	08-Aug	19	South Delta Aquarids	22h12	-17	Jul	29	02b07	77	Unfavourable	20	22h00 3	3 0	5b00 47	41
			Southern Taurids	03 <u>h2</u> 2		Ноч		00h31	46	Very good	12	21 h 30 2		3b00 34	27
			Northern Taurids	03 b 53	+22	MOA	13	00h27	38	Favourable	8	23h30 3	6 0	1h00 37	29
0ct	0 2-No v	07	Orionids	06b24	+15	0ct	22	04 <u>h</u> 23	45	Poor	20	02000 3	3 0	3b20 42	66
Nov	14-21		Leonids	10b08	+22	MOA	17	06 <u>b21</u>	38	Poor	10	03b00 1	8 0	3h30 23	71
Mov	28-Dec	09	December Phoenicids	01h00	-55	Dec	05	20b07	65	Very good	5?	20130 б	4 0	1b00 33	18
Dec	04-16		Geminids	07h28		Dec		0 1h5 5	28	Favourable	50	23h30 1	9 0	3b00 26	35
Dec	05-Jan	07	Velaids	09b56	-51	Dec	29	03 h2 5	69	Pavourable	15?	23b00 3	7 0	3 h 30 69	40

The times (SAST) and the altitudes (Alt) given, are for an observer stationed at E 30.5 S 29.5 (Durban), to be used as a guide for meteor observers elsewhere. The times for the recommended watch and the conditions, are based on the radiant altitude and on moonlight and twilight interference during the date of the predicted maximum meteor activity. The Zenithal Hourly Rate (ZHR) is the probable number of meteors expected to be seen by an experienced observer during 1 hour, when the radiant is at the zenith and with seeing conditions corresponding to a naked-eye limiting magnitude of +6.5 mag. near the zenith.

For meaningful visual work, the minimum altitude of the radiant should be 25 degrees above the horizon. Depending on the altitude (Alt) of the radiant at the time of observations, the following table gives the factor (F) by which the hourly observed rate should be multiplied to compute the ZHR: Alt (deg) 21 27 35 43 52 66 90 F 2.0 1.67 1.43 1.25 1.11 1.0

The perception coefficient (P-value) varies from observer to observer, hence the absolute necessity of keeping the count per hour (hourly rate) on an individual basis when working in group. A high P-value can mean that a specific observer gets a higher ZHR under the reported seeing conditions than the "average" observer and this can be explained by the fact that the observer may have really seen more meteors/hour or more frequently, that the limiting magnitude at the time of the observation, is not properly estimated. Hence, it is important that the faintest naked-eye star seen near the zenith, be carefully derived from reliable star atlases or catalogues.

The Stars

CONSTELLATIONS

Apart from our Sun all the stars that we see are so incredibly distant that, despite their high speed velocities, their apparent positions change by only minute amounts each year. Consequently the patterns that they form appear unchanged. The Greeks and other ancient civilisations identified these patterns, or constellations, with various mythological characters and creatures, and most of the names they gave are still used today.

In all there are 88 constellations, roughly one half of which would be above the horizon at any one time. Some contain distinctive patterns of bright stars and are relatively easy to find; others are difficult to locate, even with suitable maps. The Southern Croms and Centaurus, Orion and Taurus, Scorpius and Sagittarius, are featured later in this section. Detailed information on other constellations is beyond the scope of this handbook and interested observers are advised to obtain a suitable star atlas.

STAR NAMES

Within each constellation, the brightest star is generally labelled α (Alpha), the next β (Beta) and so on through the Greek alphabet. Most of the brightest stars also have their own names - usually of Arabic origin. For example α Canis Majoris, otherwise known as Sirius, is the brightest star in the constellation Canis Major.

STELLAR MAGNITUDES AND STELLAR DISTANCES

The apparent brightness of a star - which depends both on its true luminosity and its distance - is indicated by its magnitude. Equal intervals of magnitude represent equal ratios in light intensity.

Distances are often expressed in units of light years - the distance light would travel in a year (equal to 9.5 x 10^{12} km).

DOUBLE STARS

It now appears that single stars such as our Sun are the exception, the majority of stars being double or multiple - two or more suns in orbit around one another.

STAR CLUSTERS

These are two completely different sorts. Galactic clusters, having of the order of 100 stars, are found close to the plane of the Milky Way. The ones we can see are relatively nearby. Globular clusters are much larger and far more distant. They contain of the order of 100 000 stars each and are seen above and below the Milky Way on that side of the sky towards the centre of our galaxy. So great is their distance that small telescopes fail to resolve individual stars instead they appear as fuzzy balls.

NEBULAE

Possibly one third of the matter in our region of the Galaxy is in the form of gas and dust (the remainder being contained in stars, plus a tiny amount in planets). Condensations of this material are called nebulae, some of which are illuminated by nearby stars while others are dark. They are usually referred to by their numbers in Messier's catalogue (M) or the New General Catalogue (NGC).

THE STAR CHARTS

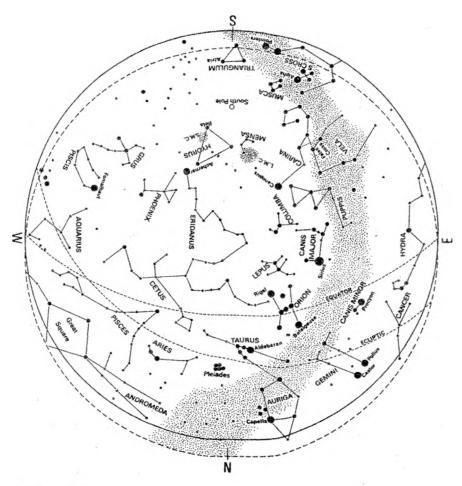
The star charts in this section show the night sky for each of the four seasons of the year. These seasonal charts depict stars down to magnitude 3.5 which is approximately what will be visible to the naked eye in city areas. Charts of 3 of the most interesting regions in the sky (showing stars down to magnitude 4.5) - the Orion region, visible in Summer; the Scorpius region, visible in Winter and the Southern Cross Region, visible all year round - are featured. They are rich in interesting objects visible to the naked eye, or with the aid of binoculars or a small telescope. To use them locate the constellations in the sky from the seasonal chart and rotate the regional chart to match the orientation of the constellations in the sky.

THE SUMMER SKY

The chart below represents the sky in Cape Town on December 1 at midnight, January 1 at $10~\rm pm$ and February 1 at 8 pm. Corrections for places other than Cape Town are

Bloemfontein and Port Elisabeth -30 minutes
Johannesburg -40 minutes
Durban -50 minutes
Harare -52 minutes

Correct times for places elsewhere may be found by subtracting 4 minutes for each degree of longitude east of Cape Town or adding 4 min for each degree of longitude west of Cape Town.



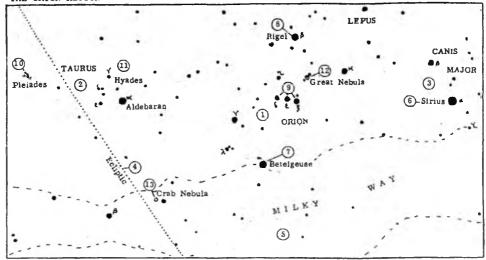
Courtesy of the

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SA MUSEUM

THE ORION REGION



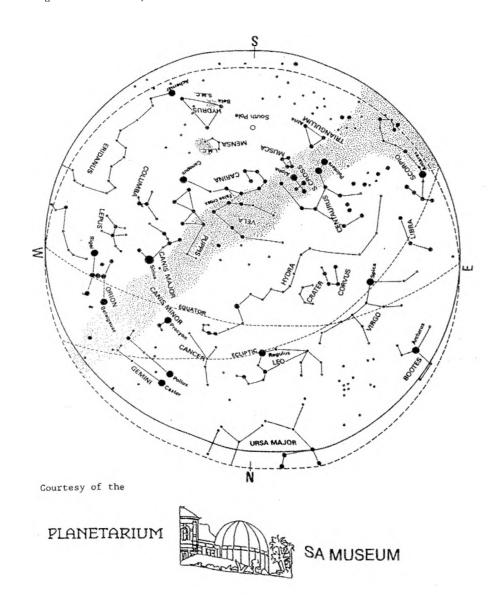
- The constellation of Orion. The figure of the legendary hunter of Greek mythology is unfortunately upside down when seen from Southern Africa. The faint stars by λ represent the head, α and Υ the shoulders, $\delta \epsilon \eta$ the belt, and δ and κ the legs. Orion forms part of the "great hunting scene" in which he faces the onslaught of 2 Taurus, the bull. Only the forepart of the bull is depicted and, like Orion, it is upside down. α and ϵ are the eyes, Υ the nose. Orion is accompanied by 3 Canis major, the large dog, and the small dog (off map) while Lepus, the hare, crouches at his feet.
- A section of the Ecliptic a line encircling the entire sky and representing the plane of the Earth's orbit. As the Earth revolves around the Sun, the Sun appears to move along the ecliptic through the constellations of the Zodiac, of which Taurus is one.
- (5) A portion of the Milky Way (looking out towards the edge of our Galaxy).
- (6) Sirius the brightest star in the night sky. It is somewhat brighter than our Sun and relatively close by at a distance of 9 light years. It is a double star but the companion is a white dwarf (only slightly larger than the Earth, and with a mass comparable to our Sun) and is only visible through a large telescope.
- Petelgeuse most famous of the red giant stars. Its diameter is of the order of the size of the Earth's orbit and its luminosity is nearly 10 000 times that of our Sun. Its red colour should be obvious to the eye. It is 520 light years distant.
- (8) Rigel, despite being physically smaller than Betelgeuse, is more luminous (higher surface temperature bluish colour) and more distant.
- (9) The stars in Orion's belt are distant bot blue stars.
- 10 The Pleiades or Seven Sisters form the best known nearby star cluster. Six or seven stars are visible to the naked eye, binoculars or a small telescope show more.
- (1) The Hyades is another nearby galactic cluster, but Aldebaran is not a member (it lies closer to us).
- The Great Nebula in Orion, just visible to the naked eye, shows up as a fan shaped mass of luminous gas through binoculars or a telescope. A telescope will also show a tiny "Trapezium" of four stars in the centre.
- The Crab Nebula, the remnant of a supernova recorded by the Chinese in 1054, requires a moderate sized telescope for observation. In its heart is located the extraodinary pulsar which emits a double flash of light 30 times every second. The current belief is that it is a rapidly rotating neutron star a star with the mass of our sum but with a diameter of only 10 km.

THE AUTUMN SKY

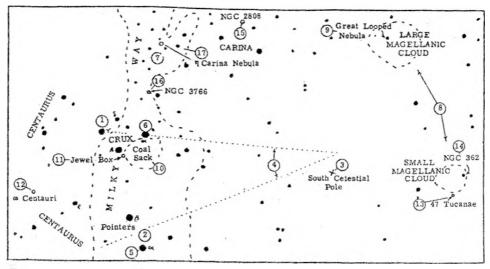
The chart below represents the sky in Cape Town on March 1 at midnight, April 1 at 10 pm and May 8 pm. Corrections for places other than Cape Town are

Bloemfontein and Port Elisabeth -30 minutes
Johannesburg -40 minutes
Durban -50 minutes
Harare -52 minutes

Correct times for places elsewhere may be found by subtracting 4 minutes for each degree of longitude east of Cape Town or adding 4 min for each degree of longitude west of Cape Town.



THE SOUTHERN CROSS REGION



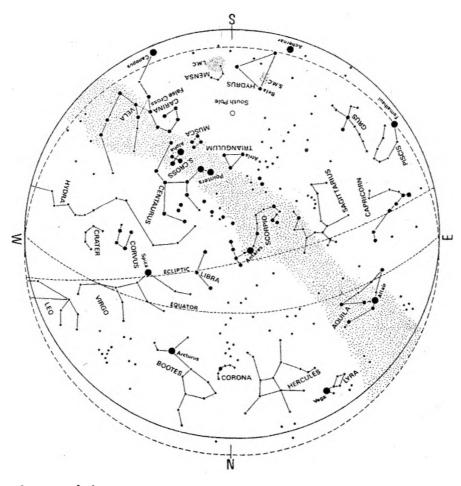
- (1) Crux, the Southern Cross, is one of the most compact patterns of bright stars to be found in the sky. It lies on the border of that region of the sky which never sets as seen from Southern Africa.
- (2) The two "Pointer" stars lie close to the Cross. (A similar pattern to the Southern Cross called the False Cross lies just outside and above the map, but has no accompanying pointer stars).
- 3 The South Celestial Pole: This is one of two opposite points in space towards which the Earth's axis of rotation is directed. As the Earth rotates so the sky appears to pivot about this point. It always lies above the south point on the horizon, elevated by an angle equal to the observer's southern latitude. (The north celestial pole lies below the northern horizon and can never be seen from the Earth's southern hemisphere)
- (4) The intersection of a line extended through the major axis of the Cross and the perpendicular bisector to the Pointers indicates the approximate position of the South Celestial Pole.
- (5) c Centauri has the distinction of being the closest star to our solar system at a distance of approximately 40 million million km or 4,3 light years. A small telescope readily shows that it is a double star the two components take 80 years to revolve about one another. A much fainter third star also belongs to the system.
- (6) α Crucis can also be resolved as a double star by a small telescope (separation 5 sec of arc).
- (7) The region indicated is one of the brightest sections of the entire Milky Way.
- (8) The Large and Small Magellanic Clouds are the nearest of the external galaxies (see also next section). They can be seen with the naked eye provided the sky is reasonably dark.
- (9) The Great Looped Nebula possibly the remnant of a supernova explosion in the Large Magellanic Cloud. (Naked eye or binoculars).
- 19 The "Coal Sack" a dark mass of gas and dust obscuring a part of the Milky Way. (Naked eye or bino-culars).
- Herschel's "Jewel Box" a galactic cluster containing stars of different colours. (Small telescope or binoculars).
- (12) ω Centauri and (13) 47 Tucanae are perhaps the best known globular clusters. Binoculars will show their fuzzy appearance. (14) NGC 362 and (15) NGC 2808 are fainter globular clusters.
- (16) NGC 3760 a fine galactic cluster. (Binoculars or small telescope).
- 17) The n Carinae nebula site of a slow supernova that brightened to magnitude -0,8 in 1843 and is now of magnitude 6,4.

THE WINTER SKY

The chart below represents the sky in Cape Town on June 1 at midnight, July 1 at 10 pm and August 1 at 8 pm. Corrections for places other than Cape Town are

Bloemfontein and Port Elisabeth -30 minutes Johannesburg -40 minutes Durban -50 minutes Harare -52 minutes

Correct times for places elsewhere may be found by subtracting 4 minutes for each degree of longitude east of Cape Town or adding 4 min for each degree of longitude west of Cape Town.



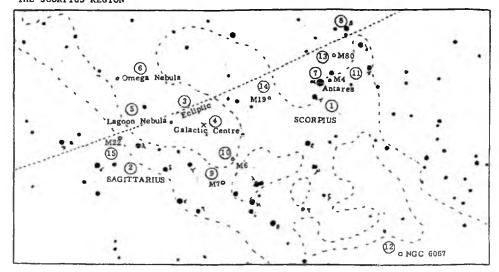
Courtesy of the

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THE SCORPIUS REGION



- 1) The constellation of Scorpius. The creature is depicted with α in the centre of the body and 8 and 1 the claws. The distinctive tail $\epsilon \zeta \ell$ curls round to the sting λ
- (2) Sagittarius the figure of the centaur archer is very difficult to make out.
- 3 A section of the Ecliptic. Like Taurus, Scorpius and Sagittarius are constellations of the Zodiac.
- The direction of the centre of our Galaxy the Milky Way is that part of our Galaxy visible to us. Unfortunately the central nucleus is obscured by foreground gaseous and dusty matter both dark and luminous hence the irregular shape of the Milky Way in this region. Luminous nebulae include (5) the Lagoon nebula and (6) the Omega nebula. These are best seen with the aid of binoculars.
- (7) Antares a distant red giant, several bundred times the diameter of our Sun is so named because its red colour rivals that of the planet Mars.
- 8 6 Scorpli can be resolved as a double star (separation 16 sec of arc) with a small telescope. In fact the brighter component is in itself a triple star, and the fainter component a double star:

This region includes a number of galactic clusters including (9 M7, (10 M8, (1) M4 and (12 NGC 6067. (Use binoculars or a small telescope).

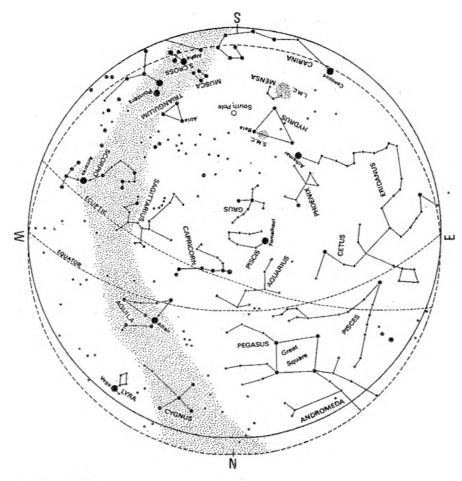
Further from the plane of the Milky Way are some globular clusters: 13 M80 (14) M19 and (15) M22.

THE SPRING SKY

The chart below represents the sky in Cape Town on September 1 at midnight, October 1 at 10 pm and November 1 at 8 pm. Corrections for places other than Cape Town are

Bloemfontein and Port Elisabeth -30 minutes
Johannesburg -40 minutes
Durban -50 minutes
Harare -52 minutes

Correct times for places elsewhere may be found by subtracting 4 minutes for each degree of longitude east of Cape Town or adding 4 min for each degree of longitude west of Cape Town.



Courtesy of the

PLANETARIUM



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VARIABLE STAR OBSERVING

The latest (1985) edition of the "General Catalogue of Variable Stars" lists more than 28 000 stars. Professional observatories cannot possibly monitor all of these, and this makes the observation of variable stars a field in which amateurs can make a real contribution to astronomical knowledge.

Of these 28 000 stars at least 2 000 are suitable for visual monitoring in the southern hemisphere. However, the number of active observers in this part of the world remains woefully small, and less than 400 variables are at present being observed from Southern Africa.

The Variable Star Section of the A.S.S.A. exists for the purpose of encouraging observers and of acting as a medium communication. The Section disseminates incoming information amongst observers, and will forward (on request) the observations of individuals to various variable star bodies. These include the American Association of the Variable Star Observers (AAVSO) and the Variable Star Section of the Royal Astronomical Society of New Zealand. These bodies combine the South African light estimates with those from other parts of the world. The resulting "light curves" and tables are made available to a large number of professional observatories where astronomers are interested in investigating certain of the stars more fully.

In recent years amateur observers have played an invaluable part by alerting the operators of orbiting satellite observatories whenever outburst occurred of certain eruptive variables.

Visual estimates of magnitude are made by comparing the variable with two or more comparison stars, respectively brighter and fainter than the unknown variable. Suitable comparison stars are shown on special charts, which have been prepared for each variable, mainly by the two variable star organisations mentioned above. The use of these charts is essential for accurate, standardized observations and intending new observers are therefore advised to obtain the necessary data by contacting the Director of the Variable Star Section,

Mr. J. Hers, P.O. Box 48, Sedgefield, 6573, Telephone (04455) 31736. They will then be sent charts of a few easy objects and data on stars which may be observed with the equipment at their disposal.

Prospective observers should, when writing, give brief details of their equipment. Larger, more powerful telescopes will naturally greatly increase the number of stars which may be measured, but many variables are bright enough to be observed through most of their cycles with quite modest equipment, e.g. binoculars. Some stars, such as 071044 L2 Puppis, are so bright that they may be observed without optical aid whatever.

Variable stars are designated in two ways. The first of these, the Harvard designation, consists of six figures which give the position for 1900; the first four figures give hours and minutes of R.A., the last two give the declination in degrees, underlined for southern declinations. The second name consists of one or two letters (or letter V plus a number) and the name of constellation.

Variables can be divided into three main classes: pulsating, eruptive and eclipsing binary stars.

Most suitable for beginners are the long period variables (or Mira variables, named after the typical representative Mira — o Ceti) which belongs to the class of pulsating stars. They are giant stars which vary through a range of brightness of 2.5 to 5 magnitudes or more, and which have well defined periodicities, ranging from 80 to 1000 days. In most cases one observation per observer every 10 days will suffice.

Typical examples include:

Approximate magnitude range

0214<u>03</u> o Ceti(Mira) 2.0-10.1

092962 R Carinae 3.9-10.0

1006<u>61</u> S Carinae 4.5-9.9

Among the eruptive variables, two groups are of special importance:

<u>U Geminorum type.</u> These are dwarf novae which have long periods of apparent quiescence at minimum, with sudden rises to maximum. A typical representative in the southern hemisphere is 040971 VW Hydri.

R. Coronae Borealis type. These are high luminosity variables with slow, non-periodic drops in brightness. A typical representative is 191033 RY Sagittarii.

Eclipsing Binary Stars have orbital planes which lie close to the line of sight of the observer. The two components periodically eclipse each other, thus causing variations in the apparent brightness of the system. Periods are generally short, of the order of hours, so that observational programmes need very careful planning. Monitoring these interesting stars is therefore for experienced observers only.

NOVA SEARCHING

On rare occasions a star may undergo a nova outburst, its light increasing tremendously. The result is that a "new" star appears where previously no star was visible to the naked eye, or even with a small telescope. The light of the nova may fluctuate for a time, then gradually fades over a period of days, weeks or months.

Even observers having no telescopes can perform a useful task in keeping a watch for such novae in an allocated area of the sky. A good knowledge of the constellations is a recommendation, since part of the procedure is to scan the whole sky for bright novae before the more detailed search in the allocated area is begun. However, anyone can be given training in star recognition.

ORDINARY OCCULTATIONS

These phenomena concern a specialised branch of observational astronony in which both professional and amateur participate. The tables of predictions must necessarily occupy a number of pages as this handbook is the sole published source for Southern Africa.

An occultation occurs when the disk of the Moon moves in front of a star. Timings of occultations, to a precision of one-tenth of a second if possible are very valuable for studies of the Moon's shape and motion. Since only very modest equipment is required, amateurs can make important contributions in this field. Persons interested in making and reporting occultation observations are urged to contact the Director of the Society's Occultation Section:

Mr. M.D. Overbeek, P.O. Box 212, Edenvale, 1610 Tel: (011) 453 6918

Predictions of occultations of all stars brighter than magnitude 7.5 supplied by Hydrographic Dept., Tokyo are given below. The main set of tables give predictions for three stations, namely:

	Longitude	Latitude
Cape Town	18°.475 E	33°.933 S
Johannesburg	28°.075 E	26°.182 S
Harare	31°.000 E	17°.800 S

This does not restrict its use to observers in those centres. The approximate time of an occultation at a place $\Delta\lambda$ degrees west and $\Delta\phi$ degrees north of one of the standard stations given above may be found from:

Approximate time = predicted time + $a.\Delta\lambda$ + $b.\Delta\varphi$ where a and b, in minutes of time, are given in the tables. Alternatively, rough times for intermediate stations can usually be estimated directly from the tables.

Occulted stars have been identified by their Z.C. numbers, that is their numbers in the "Catalogue of 3539 Zodiacal Stars for the Equinox 1950.0" by James Robertson (U S Naval Observatory, 1939).

Note: That the times of ordinary occultations are given in U.T.

EXPLANATIONS OF ABBREVIATIONS USED IN THE TABLES:

- Z.C. the number of the star in the Zodiacal Catalogue. An "m" following the number indicates the star is not single.
- Mag. the visual magnitude
- Ph the Phase: D Disappearance, R Reappearance
- h.m. the time of the occultation in UT
- a,b parameters in minutes for predicting times other than at standard stations (as explained above in the text)
- P.A. The Position Angle on the Moon's limb measured eastward from the north point

						CAPE	TOWN			JO	HANNESBU	IRG				HARA	RE	
UT						18.5	S 33		£	2	8.1 5	26.2			E :	31.0	S 17.	.8
DATE	I.C.	Mag.	Ph	ELG	TIME	a.	b.	P.A.	TIME	:	a.	b. I	A.	TI	Œ	a.	b.	P.A.
N D					p =		m	•	h	В	1	n		þ	1	1		•
Jan 2	1331	5.9		209	20 47.9	-1.2	-3.0	343										
Jan 3	1442	5.0		222												-0.5		280
Jan 4	1565≥	6.3	RD	236										22 3	36.9	-2.1	-0.3	262
Jam 10	2108≡	6.4	RD	296										1.4	16.3	-2.0	-0.1	257
Jan 10	2109	6.1		296										2 2	29.8	-2.7	+0.1	256
Jan 12	2371	4.9	DB	318	2 2.8	-1.9	+1.9	34										
Jan 12	2371	4.9		317	2 21.8													
Jan 19	3340	7.5		43	18 48.0	-0.1	+2.3	28										
Jan 20	3464	7.1		54												-1.0		59
Jan 21	42	5.6	DD	67					18 28	.1		1	36			-1.2		
Jan 24	443	7.4		105												-3.0		94
Jan 24	448	7.1	DD	105	18 46.6	-2.7	-1.0	129	19 5	. 4	-2.1 -0).1 1	11	19 1	3.7	-2.1	+0.6	88
T 25	co2	2.5	DD.	110	10.56.0				10.20		25.0		20					
Jan 25	603			119	18 56.2	-2.1	+1.0	55			-2.5 +2		37					
Feb 1	1611	5.7	ED	214	22 22 2	10.3		200			-1.2 -0			20	8.0	-1.1	-0.8	274
Feb 6	2183	5.7	RD	274	23 33.7				23 12	-1	+0.8 -3	i.U J	44					
Feb 7	2311	6.2	RD RD	285	23 28.8			333										
Feb 8	2317		ED	286	0 26.8	+1.2	-3.5	347										
Feb 9	2458	6.2	BD	298										2 5	3.2			211
Feb 21	E20	E 6	nn.	87										12.4		-1.0		114
	538	5.6		88					10.11				41			-2.0		
Feb 21 Feb 21	555		DD		20 E 2	2.0	12.0	22	19 13	. 6	-0.5 -1	. 3 1	41	19 1	1.7	-1.1	-0.1	103
	569	5.4	DD	88	20 5.3	-2.0	+2.9	33	10.16	,			7.4	12.5			2.1	120
Feb 24	1058	7.0	DD	127					18 16		10.0		74			-2.3		
Feb 24	1070	5.2	DD	129 130	23 14.6	-0.7	10.4	111	21 25	* I	+0.0 -2	.0 1	02	21 1	6.8	-1.1	-0.8	128
Feb 24	1080	6.9	ш	130	۵۵ 14.0	-0.7	10.4	TIT										
Mar 5	2134m	6.1	PD:	243	23 22.9	-1 1	-1 5	283	23 16	٥	-0.7 -2	4 3	10	22 5	1 7	+0.7	-4.1	353
Mar 7	2269	5.4	RD	255	2 20.2						-2.7 -1					-2.2		
Mar 14	3208	6.5		333	2 20.2	3.0	10.7	240	2 44	. 0	-2.7 -1		07			-0.5		
Nar 22	844	5.7	DD	83	17 31.4	-2.5	41 3	61						,	2.3	-0.5	14.1	223
Mar 22	849	6.5	DD	84	18 31.2				18 50	n	-1.7 +1	n	83	19 1	0.3	-2.8	42.1	45
Mar 23	1030	3.2		97	19 34.2			106			-2.2 +1		66 66	17 1	0.3	-2.0	73.1	40
IML 6-3	1030	312	-	,,	27 34.2	1.0	.0.5	200	17 50		2.2 ()		••					
Mar 23	1030	3.2	PR	97	20 46.2	-0.8	40.3	249										
Nar 24	1167		DD	110	18 48.5			77										
Mar 26	1405		DD	135	10 1013		10.0	* *						16.5	4.4	-1.4	-2 9	158
Mar 26	1413	6.7	DD	136	19 2.4	-1.5	-1.8	142	19 10	. 3	-2.4 -0	.8 1	19			-4.1		78
Apr 4	2371	4.9	RD	236	4 16.5				27 20	• -	2		~ -	-/-	,	***	.010	,,
Apr 4	2469	6.3	RD	244	21 7.5													
-T- '								,										
Apr 5	2622	6.3	RD	256										23	1.6	-1.9	+1.2	232
Apr 6		5.7	RD	266	22 45.3			344										
Apr 19	977	6.6	DD	67					18 40.	.9		1	69	18 2	9.7	-0.4	-0.6	129
Apr 20	1118	6.0	DD	79					16 25	.7	-1.8 -1					-2.5		
Apr 20	1128	6.9	DD	81										19 3	1.9	+1.0	-3.0	173
Apr 21	1275	5.6	DD	95	21 48.2	-0.5	t0.6	105										
-																		
Apr 22	1386	6.6												20 1	4.8	-0.1	-2.3	163
May 1	2311	6.2		205	1 40.2				1 47.				35					
May 3	2714		RD	236	22 9.6	-0.6	-1.4	276	21 57.	.7	-0.2 -3	.4 3	18					
May 6	2981m			259												-2.0		
May 6	2987m	5.0		260	2 31.9	-1.7	+2.6	208			-2.2 +2			3 2	9.6	-2.9	+1.3	247
May 6	2989	6.8	DD	260					2 55.	.5		3:	38					
 -			**	001														
May 7	3112	6.2		271	3 24.5	-1.9	+1.0	234										
May 10	3482		ED.	305										1.4	1.1	-0.1	+2.8	190
May 16	900	4.9		35							-1.6 +1		70	16.0		.1.0		
May 16	902	6.6	טט	35					16 26.	.3	-0.9 +0	.6	97	16 3	5.3	-1.5	+1.7	63

						CAPE	TOM			JO	BAJORES	SBURG			HARAR	E	
UT					E	18.5	\$ 33	3.9		E 2		S 26	6.2	E	31.0 S	17.	.8
DATE	I.C.	Mag.	Ph	ELG	TIME	a.		P.A.		LVCE	a.	b.	P.A.	TIME	a.	b.	P.A.
N D				7	h m	1			þ		3	n	•	h m	9		•
May 19	1360	7.5		77	19 53.2			83	20	26.4			70				
May 22	1688	6.3		115	20 0.3 17 29.1			118	20	26.4	-2.8	+2.6	70				
May 23 May 23	1778 1788m	7.1 6.7	DD	125 128	17 27.1	-2.4	70.8	93	22	20.7	-0.0	-2.1	156	22 13.1	-1.2	_^ 7	122
May 28	2270	5.4		174	1 42.3	-1.3	¥1 3	85	22	20.7	-0.9	-2.1	130	22 13.1	-1.2	-0.7	122
Nay 31	2692	5.7	RD	209	3 38.8				- 4	1.6	-0.8	¥1.5	251		1		
			_							2.0	***		572				
Nay 31	2694	6.4	RD	209					3	50.9			184				
Jun 1	2822	5.6	RD	218					0	12.7	-1.9	+4.3	203	0 44.5	-2.6	+2.1	232
Jun 1	2838	5.6	ED	220	4 9.2	+0.0	+3.8	196									
Jun 17	1546	7.2	DiD	73										18 56.8			
Jun 20	1858	6.5	DD	109										18 13.0	-0.8	-4.3	173
Jun 20	1872	7.3	DID	110	20 48.5	-1.6	+0.5	102	21	14.0	-1.2	+2.7	64				
Tun 21	1067		Pun	100									1.0				
Jun 21 Jun 21	1967 1970m	5.7	DD	120 120	16 38.7	-4.6	13.1	55	To	10.8	-1.2	-4.6	143	16 4.3	-2-1	-1.5	112
Jun 21	1993	6.8	DD	123	23 20.4				22	22 6	-0.6	_0 2	122	23 25.9	-0.3	10 E	95
Jun 22	2084	6.5		131	25 20.4	-1.0	1.2	140	23	٠.٠	-0.0	-0.2	126	17 13.9			
Jun 25	2371	4.9	DO	157	0 18.2	-1.8	-0.1	117	٥	34.7	-1.2	+0.4	103	0 43.7			80
Jul 2	3371		RD	244	21 55.4				•	4417			203	0 1317	***	-2.00	••
Jul 6	221m	3.7	RD	283										3 24.6			173
Jul 16	1723	7.1	DD	68					19	23.5	-0.5	-1.2	146	19 19.3	-0.5	-0.2	114
Jul 16	1726	6.9	DD	68	19 51.1			60									
Jul 18		5.7	DD	92	18 38.4	-1.3	-2.9	159	18	41.8	-1.8	-0.7	121	18 47.3	-2.0	+0.7	89
Jul 19	2066m	6.4	DD	104										21 11.7			183
Jul 20	2183	5.7	DD	115	19 45.1	-2.5	+2.9	59									
Jul 21	2299	6.4	DΒ	125					16	46.2	-0.9	-4.3	154	16 25,8	-2.1	-1 0	110
Jul 21	2311	6.2		126	18 54.9	-2 0	41 1	72		52.5	~0.9	-4.3	22	10 23,8	-2.1	-1.9	119
Jul 21	2317	6.6		126	20 35.1			83			-1.5	+2.4	59	21 42.2			19
Jul 21		6.4	DD	128	23 9.3			36		37.2	1.5		12	21 44.2			17
Jul 22	2452	6.7	DD	138	21 45.7	G		174		_	-3.0	-1.8	_	21 43.8	-2.3	ю.1	101
Jul 22	2455			138	21 52.6						-2.5			22 11.7			95
Jul 23	2469	6.3	DD	140	1 51.5	-0.2	+0.7	106									
Jul 31	42	5.6	RD	239	23 4.3	-1.1	-0.4	252	23	15.1	-1.9	-0.5	262	23 13.2	-3.8	-2.5	291
Aug 3	302	6.4	₽D	264	0 20.6			228	0	29.7	-1.0	+0.4	235	0 35.8	-1.7	Ю.1	255
Aug 3	320	5.9	2 D	267	5 6.1			296									
Aug 16	2134∎	6.1	DD	85	18 5.6			95			-1.5		65				
Aug 17	2257	6.7	DD	96	16 48.1	-2.6	-0.5	98	17	22.0	-2.9	+2.4	62				
Aug 17	2270	5.4	DD	97	20 34.2	_1 2	43.0	94	20	52.2	-0.6	41 2	79	21 6.9	40.1	L7 N	54
Aug 18	2411		DD	109	20 34.2	-1.3	11.0	71	20	JE.J	.0.0	.1.3	13	21 10.1			
Aug 20	2689	6.8	DD	130	20 51.6	-2.8	-0.3	107	21	17.5	-2.3	+0.5	97	21 30.5			7 7
Aug 20	2692	5.7	DD	130	21 18.7				_		-1.9		92	21 57.4			73
Aug 21	2706	5.8	DD	132	0 13.5			68			+0.1		64	0 38.4			46
Aug 21	2822	5.6	DD	140										18 15.7	-3.6	-1.9	111
Aug 21	2838	5.6		142										22 56.7			
Aug 29		5.7		235	22 56.7						-1.8			23 12.1			
Sep 1		3.8		261	0 39.2						+0.2			1 13.3			
Sep 1 Sep 1	536 539	5.4 4.4		261 261	1 5.0			234			-1.5 -2.3			1 27.1			
Sep 1	541	4.0	ED.	261	1 39.2	_					-1.8			2 7.8			253
·					1 37.2	***			•	-4.7			200			J.7	
Sep 1	542	5.9	RD	261	1 50.6	-1.9	-0.7	269	2	7.4	-2.6	-0.5	273	2 9.2	-3.5 -	1.4	293
Sep 1	543	6.5		261	1 55.1			261			-2.5			2 17.6			284
Sep 6	1337	5.6		329							-0.6						
Sep 11	1967	5.7	DD	42	17 53.9	-1.3	-3.7	171	17	49.3	-0.7	-0.9	138	17 47.7	-0.4 -	0.0	108

						CAPE	TOWN			JC	HANNE	SBURG			HARARI		
Uľ					E 1	8.5	S 33	.9		E 2	28.1	S 26	5.2	E	31.0 S	17.	8
DATE	I.C.	Mag.	Pb	ELG	TIME	a.		P.A.		ME	à.		P.A.	TIME	à.		P.A.
N D	1970m	6.2	DD.	43	h m 18 10.9	.0.1	14.5	42	h		ı	•		p .		1	
Sep 11 Sep 12	2095	7.2		55	19 59.4			49									
Sep 15	2483	7.1	DD	88	16 33.9				16	59.0	-2.7	+0.6	87	17 18.0	-2.2 +	2.1	59
Sep 15	2499	6.6	DD	89	20 11.8			53			+0.1		44	20 56.0			14
Sep 15	2500m	3.4	DD	89	20 15.9	-0.9	+1.5	78		_	-0.4		70	20 48.3	+0.2 +	1.9	50
Sep 15	2500m	3.4	RB	89	21 26.1	-0.3	+1.4	259									
Sep 16	2641	7.4	DD	101	21 41.0	+0.5	+3.4	22	21	59.6	+1.2	+3.6	14				
Sep 17	2769	6.3	DD	110	16 37.6	-2.3	+0.9	61	17	16.5	-2.2	+3.2	36	18 4.3			356
Sep 17	2777	7.0		111										19 22.3			
Sep 17	2785	6.8		111	20 43.2				21	4.6	-2.2	-0.3	116	21 11.5	-1.4 4	Ю.5	94
Sep 17	2802	6.4	DD	113	23 58.9 20 18.1			48									
Sep 18	2908	6.9	טע	122	20 18.1	11.3	10.8	354									
Sep 19	3019	5.9	DD	131	16 53.0	-1.6	+1.4	40	17	29.6	-1.3	+4.1	14				
Sep 19	3029	6.9		133	20 29.0			102	20	59.6	-3.1	+0.2	101	21 12.5	-2.2	Ю.9	82
Sep 20	3146	6.5		142	17 31.0			19	18	12.6			350				
Sep 25	221m	3.7	DD	204	20 11.9			47									
Sep 25	221≡ 245	3.7 6.1	RD RD	204	21 19.5				21	34.4	-2.0	+0.1	253	21 40.3	-3.0 -	0.5	2/2
Sep 26	240	0.1	ш	207	3 3.0	-1.5	11.8	236									
Sep 30	842	6.3	RD	258	2 45.9	-2.2	-0.4	269	3	6.1	-2.6	-0.5	284				
Oct 1	1015	6.4	ED	271	1 52.5	-1.5	-1.2	280			-2.2			1 54.4	-2.6 -	-2.2	314
Oct 1	1023	6.5	RD	271	2 59.2	-2.0	-1.2	292						2 57.5	-2.4 -	3.9	338
Oct 11	2305	5.9		47	18 56.4			41	19	12.1			24				
Oct 11	2314	5.8	DD	47	19 59.1	+0.4	+1.5	68									
Oct 12	2443	5.8	DD	58					19	21.7			164	19 10.6	-0.7 -	-0.5	121
Oct 12	2443	5.8	RD	58					19	29.2			178				
Oct 15	2859	6.7	DD	90	18 23.5	-6.1	-5.4	137	18	45.9			127	18 49.8	-2.4 +	Ю.2	99
Oct 15	2859	6.7	D₽	90	18 48.0	+2.7	÷8.5										
Oct 18	3216	6.6	DD	123	17 49.2			118	18	13.1	-4.7	-1.7	108	18 21.3			85
Oct 19	3326	6.4	DD	134	20. EE 4		2.5	120	22	20. 2			107	16 26.0			
Oct 19	3340	7.5	DĐ	136	20 55.4	-5.5	-2.5	122	21	28.2			127	21 32.1	-2.1 1	ru. 2	99
Oct 24	440m	4.6	RD	197					18	11.3	+0.4	÷1.3	206	18 18.7	-0.2 t	0.6	231
Oct 30	1259	5.9	RD	268	2 39.1	-1.8	-1.3	292	2	46.3	-1.9	-1.8	316	2 31.1	-1.5 -	-3.4	344
Nov 10	2675	7.1	DD	49	19 13.9	+0.5	+2.6	31									
Nov 14	3163	7.3	DD	91										16 42.9	-0.7 1	3.7	13
Nov 14	3184	7.1 5.3	DD	94 94	21 47.4			79 92	22	1.0	-0.1	41.1	83				
Nov 14	3185	2.3	טט	74	21 47.0	-0.0	71.2	92	22	1.7	-0.1	71.1	63				
Мот 15	3287m	5.9		103							-1.0			18 34.1			5
Nov 16	3417	6.8	DD	116	22 1.2			60			-0.5		51	22 41.7	-0.3 4	2.3	31
Mov 19	245	6.1		151	18 36.1	-0.4	+1.7	19			-0.7		17	19 26.1			352
Mov 20	266∎	5.7	DD	154	0 3.8	1.0		136			-0.9		107		-0.9 1		82
Nov 22	552 552	3.0	DD 20	177 184	0 6.5			121			-1.7 -1.3		98		-1.9 f		74 283
Nov 22	332	3.0	au.	104	1 9.1	-1.7	12.0	221	1	3/.4	-1.3	71.2	200	1 40.0	-0.8 1	ru. 2	203
Nov 25	1205			235	23 12.1								307		-1.9 -	3.0	
Mov 26	1336			248	22 32.0	-0.7	-1.9	304			-1.1	-2.3		22 0.6			358
Nov 28	1458 1688m	5.9 6.3		263	2 12 4	_1 1	-1.4	77^		45.2	_1 2	-1.0	241	1 54.9			
Mov 30 Dec 10	3015	5.3		50	2 13.4 19 4.1					21.5	-1.2	-1.9	306 5	1 3/.6	-0.8 -	2.3	330
Dec 16	191			119	1/ 4-1	1914	14.7	10	47	47.3			,	20 18.9	-2.3 -	-0.3	109
														44			
Dec 22	1129			199							-0.2			19 15.3			
Dec 22	1129	5.3	KU	199					19	10.5	-0.2	10.9	225	19 15.3	-0.8 1	J.U	201

GRAZING OCCULTATIONS

When a star moves tangentially to the limb of the Moon, and is occulted for a very short period only - a few minutes, or even seconds - a grazing occultation is said to occur. Because the limb, as seen from the Earth, is in fact the outline of numerous mountains and valleys, there may be several disappearances and reappearances, which are not only fascinating to observe, but which may be accurately timed to yield valuable data on the relative positions of star and Moon, as well as on the shape of the Moon. Some of this data cannot readily be obtained in any other way.

The maps on the following pages have been prepared by the Hydrographic Dept., Tokyo to show the tracks of stars brighter than 7.5 magnitude which will graze the limb of the Moon when it is at a favourable elongation from the Sun and at least 10° above the observer's horizon (2° in the case of bright stars). Each track starts in the west at some arbitrary time given in the key and ends beyond the area of interest, except where the Moon is at a low altitude, the bright limb or sunlight interferes.

The tracks as shown on the maps are approximate only. Since the observer's location is very critical, successful observations call for very accurate predictions. Such predictions, which include graphical representations of the probable profile of the Moon, are computed annually for a number of centres in Southern Africa. By plotting the predicted graze track on a reliable survey map (e.g. South African 1:50 000 series) it is usually possible to select a convenient site from where the graze may be observed. Ideally a team of observers would be stationed at intervals along a line running at right angles to the graze track - say, along a main road - each with his own telescope and timing equipment. Each observer will see a different sequence of events, the combined results forming an accurate picture of the limb of the Moon.

The equipment needed is similar to that used for ordinary or "total" occultations, but must, of course, be portable. A 75 mmm refractor is ideal for average events, but instruments with a larger aperture have often shown their superiority under difficult conditions. Timing is best carried out with a portable tape recorder and a small FM radio tuned to a pre-arranged transmission.

It will be seen from the maps that many grazing occultations occur in regions which are rather far removed from the main cities, and which cannot easily be reached by teams of observers from one of the ASSA centres. It is worth remembering, however, that a team of many observers, while ideal, is by no means essential, that a single good observer is worth more than many unsuccessful ones, and that one good observation is worth infinitely more than no observation at all.

Interested observers - especially those living in the more distant regions - who wish to be informed of favourable grazes occurring in their neighbourhood, are therefore invited to contact the co-ordinator for grazing occultations:

Mr.M.D. Overbeek, P.O. Box 212, Edenvale, 1610. Tel: (011) 453 6918

EXPLANATION OF THE COLUMN HEADINGS IN THE TABLES:

SEQ : Sequential number in the year. The same number is attached to the corresponding track on the map.

NZC NO Zodiacal Catalogue number of the star.

MAG Magnitude of the star.

MON, DAY, H, M, S Month, day, hour, minute and second in SAST for the west end of the track.

SUNLIT (%) Percentage of the Moon sunlit (a minus sign indicates a

waning Moon).
LIMIT Whether the t

: Whether the track is the north (N) or the south (S) limit of the occultation.

(A) denotes that the Moon is at a low altitude.

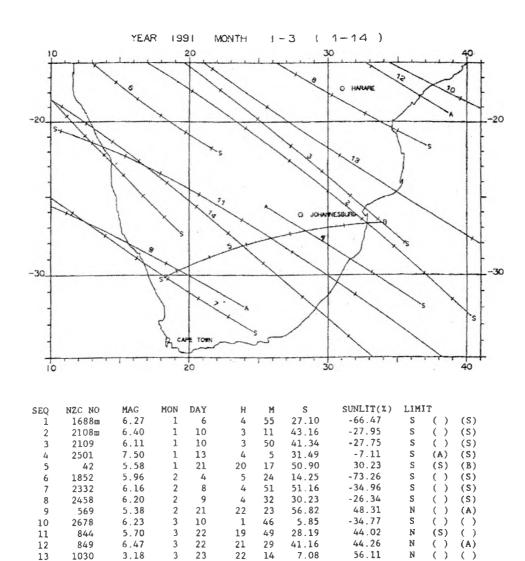
(B) denotes that the star is occultated at bright limb.

(S) denotes that the daylight interferes.

The map gives the graze tracks or the limits of occultations. Along each track on the map tick marks are given for the points corresponding to the multiples of five minutes of every hour, while the prediction for the west end of each track is shown in the table. e.g.if the time for the west end of a track is 5h 43m 2ls, the tick marks proceeding eastward correspond to 5h 45m 00s, 5h 50m 00s, 5h 55m 00s atc.

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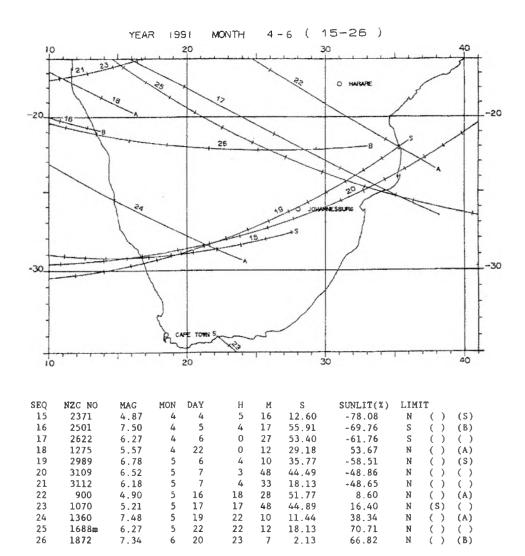
52

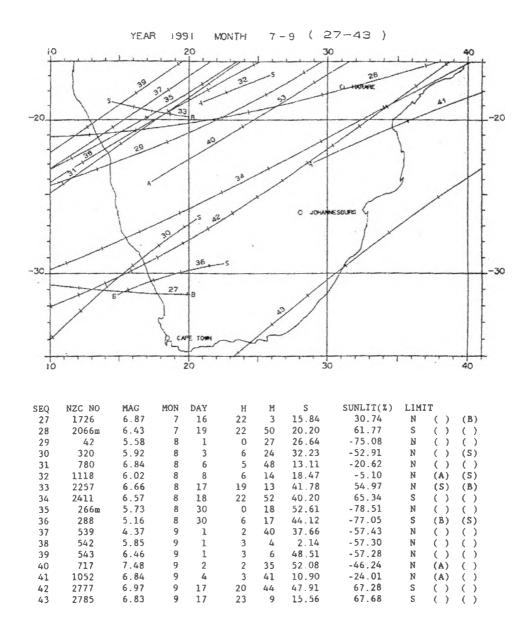
45.88

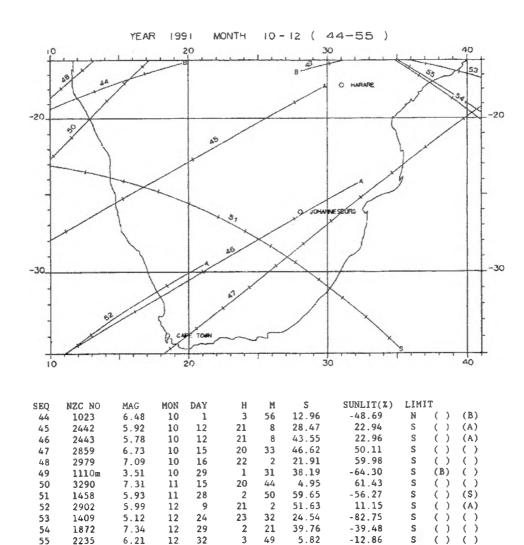
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MINOR PLANET OCCULTATIONS

A number of A.S.S.A. members and professional observatories form part of a worldwide network which observes the above events. The Southern Africa network comprises approximately 35 observers and more observers are badly needed. Very little experience is needed, apart from the ability to locate some of the fainter naked eye stars and familiarity with the user's telescope, which does not have to be a large equatorial. The only other equipment needed is a small FM radio and portable tape recorder.

Observations, especially when made by more than one observer, can be used to refine our knowledge of the size, shape and orbit of a minor planet, to greater accuracy than that obtainable with large Earth-based instruments.

Further information and detailed instructions on finding the occulted stars can be obtained from:

M.D. Overbeek, P O Box 212, Edenvale, 1610. Tel (011) 453-6918.

		0.0	CULTATAONS E	Y MINOR	PLANETS			
Date	SAST	MINOR PLANE	r Pmag	CAT	STAR	Smag	Mag	Duration
d	h m		_				drop	seconds
JAN 3	21 31	121 HERMION	E 12.3	FAC	176795	11.3	2.1	15.5
JAN 17	23 22	11 PARHTEN	OPE 10.0	LickV	1712	10.4	0.5	11.9
JAN 18	00 53	532 HERCULI	NA 9.4	PPM	96101	8.7	1.8	17.9
JAN 25	03 43	203 POMPEJA	14.9	SAO	183986	9.5	5.4	3.5
JAN 1	00 16	43 ARIADNE	11.8	SAO	157542	8.8	3.1	12.4
FEB 5	00 07	201 PENELOP	E 12.7	PPM	125971	10.5	3.2	9.7
FEB 12	19 10	886 WASHING	TINIA 13.9	PPM	92304	8.6	6.2	3.4
MAR 7	01 22	514 ARMIDA	13.9	PPM	156222	9.4	5.4	8.4
MAR 9	01 09	199 BYBLIS	13.6	PPM	1.01332	12.3	2.3	4.2
MAR 11	21 05	15 EUNOMIA	9.6	FAC	356408	10.7	0.7	66.9
MAR 17	04 45	334 CHICAGO	14.1	SAO	162723	7.2	6.9	7.0
MAR 17	21 20	788 HOHENST	EINA 12.2	PPM	178847	9.3	3.8	6.3
MAR 26	02 44	747 WINCHES	TER 12.8	PPM	128799	10.4	3.4	13.5
MAR 29	02 51	846 LIPPERT	A 16.0	SAO	158344	8.0	8.1	4.1
MAR 1	01 54	96 AEGLE	11.7	SAO	156875	9.2	2.6	10.0
APR 2	23 14	624 HEKTOR	14.5	SAO	181911	8.2	6.4	15.0
APR 15	05 17	177 IRMA	14.8	SAO	184383	7.3	7.6	12.0
APR 18	19 24	121 HERMION	E 13.9	FAC	211278	10.7	4.1	7.8
MAY 27	05 13	674 RACHELE	12.1	SAO	184425	7.3	4.8	6.2
JUN 15	06 11	356 LIGURIA	13.1	SAO	210543	7.9	5.3	11.4
JUN 18	02 42	198 AMPELLA	11.2	AC	15624	10.0	1.5	8.4
JUN 19	19 30	776 BERBERI	CIA 12.5	SAO	159636	9.2	3.3	13.9
JUN 24	19 39	103 HERA	12.8	PPM	159359	9.1	4.6	8.2
JUN 24	20 53	514 ARMIDA	15.3	PPM	156789	9.8	6.4	3.4
JUN 26	06 06	41 DAPHNE	13.4	AGK3	+06 0115	8.9	4.5	9.9
JUL 2	03 39	56 MELETE	11.9	PPM	173839	11.1	1.9	14.8
JUL 3	03 08	130 ELEKTRA	11.6	AGK3	-00 2485	8.4	3.2	11.6
JUL 4	22 34	53 KALYPSO	13.5	SAO	164223	9.1	4.4	12.5
JUL 7	21 16	899 JOKASTE	16.2	SAO	157866	9.3	6.9	5.5
AUG 8	22 41	432 PYTHIA	12.3	SAO	166014	6.2	6.0	6.7
AUG 11	00 05	842 KERSTIN	15.4	PPM	144729	12.4	4.0	3.5
AUG 23	23 38	404 ARSINOE	13.2	SAO	185353	9.0	4.3	10.6
SEP 28	19 49	379 HUENNA	14.5	AC	8235	10.9	3,6	2.8
OCT 1	23 22	920 ROGERIA	15.5	SAO	145812	9.1	6.4	2.5
OCT 16	21 13	91 AEGINA	11.7	PPM	117499	11.1	1.8	10.3
OCT 27	04 09	163 ERIGONE	13.2	LickV	27062	10.6	2.7	3.9
NOV 1	05 02	363 PADUA	12.4	PPM	118793	8.2	5.2	8.7
NOV 12	22 05	1723 KLEMOLA	15.0	SAO	129042	8.4	6.6	8.6
DEC 5	04 47	67 ASIA	13.5	PPM	156397	7.8	6.6	7.7
DEC 19	21 07	56 MELETE	13.0	AC	26247	10.5	2.6	6.4
DEC 30	23 11	287 NEPHTHYS	11.4	PPM	122416	8.0	4.3	5.8

TIME SYSTEMS

This section is intended to serve established amateurs and professional astronomers - i.e. those having some knowledge of time and coordinate systems. Space in this booklet does not permit full explanation, which in any case would appear complicated to the layman.

TIME SIGNALS

CSIR has recently developed a new time service available through the telephone line with an accuracy to within one millisecond. This service replaces the ZUO service which has been discontinued.

Prospective users need have access to a telephone, an IBM-compatible PC with a modem and a "pulse buffer unit" which will be needed to synchronise other timing equipment external to the PC.

Registered users will be supplied with an authorised access code and user manual on a floppy disc. On running the software supplied, the user's computer automatically dials the CSIR time service number and establishes a link with the time service computer. The user's PC is then set to within one electronic "clock tick" of CSIR's national time standard. At the same time a pulse is generated at a pin on the printer port of the user's computer which is accurate to within one millisecond of the national time standard. This pulse can be used to synchronise other timing equipment external to the PC.

Users must pay a registration fee, a monthly fee and a fee for each call made to the system to the CSIR. In addition the user incurs the normal Post Office telephone charges.

Enquiries to: CSIR Time Service, Rm 230, Division of Production Technology, CSIR, P 0 Box 395, Pretoria, 0001. Tel: (012) 841-2036/841-4623. Telefax: (012) 841-2131.

SOUTH AFRICAN STANDARD TIME

South African Standard Time (as in everyday use) is mean solar time on the 30 East meridian (which runs east of Johannesburg and just west of Durban) and is exactly 2 hours ahead of Universal Time.

TIME OF SUN'S TRANSIT OVER 30° MERIDIAN

The table below gives the SAST when the Sun transists the $30\,^\circ$ meridian - and a sundial on that meridian reads noon.

		h	Œ	s			h	B	s			h	П	s
Jan	1	12	3	24	May	11	11	56	20	Sep	18	11	54	18
	11	12	7	47		21	11	56	31		28	11	50	48
	21	12	11	13		31	11	57	33	0ct	8	11	47	41
	31	12	13	24	Jun	10	11	59	18		18	11	45	16
Feb	10	12	14	14		20	12	1	25		28	11	43	50
	20	12	13	48		30	12	3	31	Nov	7	11	43	41
Mar	2	12	12	14	Ju1	10	12	5	16		17	11	44	54
	12	12	9	53		20	12	6	18		27	11	47	29
	22	12	7	2		30	12	6	25	Dec	7	11	51	17
Apr	1	12	4	0	Aug	9	12	5	32		17	11	55	55
•	11	12	1	9		19	12	3	42		27	12	0	52
	21	11	58	47		29	12	1	3		31	12	2	49
May	1	11	57	7	Sep	8	11	57	49					

CORRECTION FOR PLACES NOT ON THE 30° MERIDIAN

Approximate longitude corrections from the 30° East Meridian are provided below. To find the time of Sun's transit over the local meridian, apply the longitude corrections to the data in the table.

Bloemfontein	+15	East London	+8 th	Port Elizabeth	+18**
Bulawayo	+6 ^m	Grahamstown	+14**	Pretoria	+7 ^m
Cape Town	+46 ^m	Johannesburg	+8 ***	Harare	-4 ^m
Durban	- 4 ^m	Kimberlev	+21 ^m	Windhoek	+52 ^m

SIDEREAL TIME ON THE 30° MERIDIAN

			Αt	1	٩٤				Αt		At			· ·	Λt		At
		0 1	hrs	21	hrs	;		0	hrs	21	hrs			0 1	nrs	21	hrs
		h	m	h	m			h	m	h	[30]			h		h	100
Jan	1	6	40	3	44	May	11	15	13	12	16	Sep	18	23	45	20	49
	11	7	20	4	23		21	15	52	12	56		28	0	25	21	28
	21	7	59	5	3		31	16	32	13	35	0ct	8	1	4	22	8
	31	8	39	5	42	Jun	10	17	11	14	15		18	1	44	22	47
Feb	10	9	18	6	21		20	17	51	14	54		28	2	23	23	26
	20	9	57	7	1		30	18	30	15	33	Nov	7	3	2	0	6
Mar	2	10	37	7	40	Jul	10	19	9	16	13		17	3	42	0	45
	12	11	16	8	20		20	19	49	16	52		27	4	21	1	25
	22	11	56	8	59		30	20	28	17	32	Dec	7	5	1	2	4
Apr	1	12	35	9	39	Aug	9	21	8	18	1.1		17	5	40	2	44
	11	13	15	10	18		19	21	47	18	51		27	6	20	3	23
	21	13	54	10	57		29	22	26	19	30		31	6	35	3	39
May	1	14	33	11	37	Sep	8	23	6	20	9						

CORRECTION FOR PLACES NOT ON THE 30° MERIDIAN

Approximate longitude corrections from the 30° East Meridian are provided below. To find the sidereal times at SAST 0 hrs and SAST 21 hrs apply the following corrections to the data in the table.

Bloemfontein	-15 ^m	East London	-8 ^m	Port Elizabeth	-18 ^m
Bul awayo	-6 ^m	Grahamstown	-14 ^m	Pretoria	-7°
Cape Town	-46 ^m	Johannesburg	- 8 ^m	Harare	+4**
Durban	+4**	Kimberley	-21 ^m	Windhoek	-52 [®]

TELESCOPE SETTING

When a telescope equipped with setting circles is aimed on the meridian, its R.A. circle should read the sidereal time. Thus one can calculate the sidereal time and then set the circle, but it is usually simpler to aim the telescope at one of the well known stars given below and then to adjust the R.A. circle.

A LIST OF BRIGHT STARS FOR CHECKING TELESCOPE CIRCLES (1991.5)

Star	R.A. h m	Dec.	Mag.	Sp.	Star	R.A. h m	Dec.	Mag.	Sp.
ACHERNAR	1 37.4	-57 16	0.6	B5	PROCYON	7 38.9	5 15	0.5	F5
ALDEBARAN	4 35.4	16 30	1.1	K5	REGULUS	10 7.9	12 1	1.3	B8
RIGEL	5 14.1	-8 13	0.3	В8	SPICA	13 24.8	-11 7	1.2	B2
BETELGEUSE	5 54.7	7 24	0.4	MO	ARCTURUS	14 15.3	19 14	0.2	KO
CANOPUS	6 23.7	-52 41	-0.9	FO	ANTARES	16 28.9	-26 25	1.2	M1
SIRIUS	6 44.8	-16 42	-1.6	AO	ALTAIR	19 50.4	8 51	0.9	A5

				JULIAN	DATE AT	1400 HOURS	- SAST	1991				
	JAN.	FEB.	MAR.	APR.	XAY	JUN.	JUL.	AUG.	SEP.	OCT.	NOV. 2448	DEC. 2448
	2448	2448	2448	2448	2448	2448	2448	2448	2448	2448		
1	258	289	317	348	378 379	409 410	439 440	470 471	501 502	531 532	562 563	592 593
2	259 260	290 291	318 319	349 350	380	411	441	472	502	533	564	594
4	261	292	320	351	381	412	442	473	504	534	565	595
5	262	293	321	352	382	413	443	474	505	535	566	596
6	263	294	322	353	383	414	444	475	506	536	567	597
7	264	295	323	354	384	415	445	476	507	537	568	598
8	265	296	324	355	385	416	446	477	508	538	569	599
9	266	297	325	356	386	417	447	478	509	539	570	600
10	267	298	326	357	387	418	448	479	510	540	571	601
11	268	299	327	358	388	419	449	480	511	541	572	602
12	269	300	328	359	389	420	450	481	512	542	573	603
13	270	301	329	360	390	421	451	482	513	543	574	604
14	271	302	330	361	391	422	452	483	514	544 545	575 576	605 606
15	272	303	331	362	392	423	453	484	515			
16	273	304	332	363	393	424	454	485	516	546	577	607
17	274	305	333	364	394	425	455	486	517	547 548	578 579	608 609
18	275	306	334	365 366	395 396	426 427	456 457	487 488	518 519	549	580	610
19 20	276 277	307 308	335 336	367	397	428	458	489	520	550	581	611
						429	459	490		551	582	612
21 22	278 279	309 310	337 338	368 369	398 399	430	460	490 491	521 522	552	583	613
23	280	311	339	370	400	431	461	492	523	553	584	614
24	281	312	340	371	401	432	462	493	524	554	585	615
25	282	313	341	372	402	433	463	494	525	555	586	616
26	283	314	342	373	403	434	464	495	526	556	587	617
27	284	315	343	374	404	435	465	496	527	557	588	618
28	285	316	344	375	405	436	466	497	528	558	589	619
29	286		345	376	406	437	467	498	529	559	590	620
30	287		346	377	407	438	468	499	530	560	591	621 622
31	288		347		408		469	500		561		022
	JANUARY			FEBRU	JART			MARCE			APRII	
Su No	Tu We Th	Fr Sa	5	Su No Tu We	e Th Pr Sa		Su No Tu	We Th Fr	Sa	Su	Mo Tu Vie 1	
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	7 8 9			2 3 4 5				10 11 12			5 6 7	
	14 15 16			9 10 11 12				17 18 19			12 13 14 1	
	21 22 23			6 17 18 19			28 29 30	24 25 26	21		19 20 21 2 26 27 28 2	
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	10 11 12			_	10 11 12		3 4 5	6 7 8		8	9 10 11 1	2 13 14
	17 18 19				17 18 19			13 14 15			16 17 18 1	
	24 25 26	27 28			3 24 25 26			20 21 22			23 24 25 2	26 27 28
29 30			2	28 29 30	31		24 25 26	27 28 29	30	29	30 31	

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