# ASTRONOMICAL HANDBOOK FOR SOUTHERN AFRICA

published by
 the Astronomical Society
 of Southern Africa

1975

# **1975 CALENDAR**

#### JANUARY

S		5	120	19	26
м		6	13	200	270
т		- 7	14	21	28
w	1	8	15	22	29
т	2	9	16	23	30
F	3	10	17	24	31
S	40	11	18	25	

#### APRIL

S					
Μ					
Т	1	8	15	22	29
W	2	9	16	23	30
Т	3₽	10	17	24	
F	4	110	18	<b>25</b> O	
S	5	12	190	26	

#### JULY

S		6	13	20	27
Μ		7	14	21	28
Т	10	8	150	22	29
W	2	90	16	230	30
Т	3	10	17	24	31
F	4	11	18	25	
s	5	12	19	26	

#### OCTOBER

S		50	120	19	26
М		6	13	200	27
Т		7	14	21	<b>280</b>
W	1	8	15	22	29
Т	2	9	16	23	30
F	3	10	17	24	31
s	4	11	18	25	

#### FEBRUARY

## MARCH 2 9 16 23 30

6 13 200 270 7 14 21 28

5 12 19

1 8 15 22 29

4

JUNE 1 8 15 22 29 2 **9 16 2** 30 **30** 3 10 17 24

11 18

5 12 19 26 6 13 20 27 7 14 21 28

3 10 17 24 31 4011 18

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2	9	16	23
30	10	17	24
4	11•	18	25
5	12	190	26O
6	13	20	27
7	14	21	28
8	15	22	

#### MAY

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4	110	180	250
5	12	19	26
6	13	20	27
7	14	21	<b>2</b> 8
8	15	22	29
9	16	23	30
010	17	24	31

#### AUGUST

3	10	17	24	31
4	11	18	25	
5	12	19	26	
6	13	20	27	
70	140	210	28	
8	15	22	29	
9	16	23	30 O	

#### NOVEMBER

2	9	16	23	3
3 ●	100	17	24	
4	11	18	25	
5	12	190	26 🜒	
6	13	20	27	
7	14	21	28	
8	15	22	29	

- New Moon
- 0 First Quarter
- 0 Full Moon
- Last Quarter 0

JULIAN DATE - See Inside Back Cover

Front Cover - Telescope built by Mr. Arnold Mills, Hillside, Bulawayo

#### SEPTEMBER

	7	14	21	28 <b>C</b>
1	8	15	22	29
2	9	16	23	30
3	10	17	24	
4	11	18	25	
5 •	120	19	26	
6	13	200	27	

#### DECEMBER

	7	14	21	28
1	8	15	22	29
2	9	16	23	30
3●	100	17	24	31
4	11	180	25 0	
5	12	19	26	
6	13	20	27	

# ASTRONOMICAL HANDBOOK FOR SOUTHERN AFRICA 1975

This booklet is intended both as an introduction to observational astronomy for the interested layman even if his interest is only a passing one - and as a handbook for the established amateur or professional astronomer.

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(Conventional calendar - inside front cover, Julian Date - inside back cover)

#### NOTE

Unless stated otherwise, all times are SOUTH AFRICAN STANDARD TIME. In order to avoid confusion between a.m. and p.m., the 24-hour clock is used (e.g. 1800 hrs. is 6 p.m. and 2100 hrs. is 9 p.m.). Emphasis is given to phenomena visible in the evening sky - between sunset and midnight.

> This Handbook is produced in the Department of Astronomy, University of Cape Town, for the Astronomical Society of Southern Africa. Most of the data it contains have been adapted for Southern Africa from the "Astronomical Ephemeris for 1974" issued jointly by the Nautical Almanac Offices of the Royal Greenwich Observatory, Herstmonceux and the U.S. Naval Observatory, Washington D.C. Additional information has also been supplied direct from the Herstmonceux Office.

> Thanks are due to all who have contributed to this handbook, especially: The staff of HMNAO, Herstmonceux; Messrs. G. and C. Larmuth (planet rising and setting calculations and diagram, phenomena of Jupiter's satellites and time of Sun's transit) Mr. J.C. Bennet (director of Comet and Meteor Section); Mr. M.D. Overbeek (director of Variable Star Section); Mr. A.G.F. Morrisby (director of Occultation Section); Mr. J. Hers (co-ordinator for Grazing Occultations); and Mrs. P. Dobbie who typed the script.

> All correspondence concerning this booklet should be addressed to the Handbook Editor, Astronomical Society of Southern Africa, Department of Astronomy, University of Cape Town, Rondebosch 7700, Cape. Further copies can also be obtained from the same address. Enclose a postal order or cheque (in favour of the Astronomical Society of Southern Africa) for 50 cents per copy.

Although every care has been taken in the compilation of the Handbook, it is distributed and sold on the explicit condition that neither the Astronomical Society of Southern Africa nor any of its members accepts any responsibility for errors.

Dit is met spyt dat ons u meedeel dat as gevolg van beperkte fondse en produksiefasiliteite dit nie moontlik was om hierdie handboek in Afrikaans te laat druk nie.

> A. P. FAIRALL EDITOR

# ASTRONOMY IN SOUTHERN AFRICA

As one of the few parts of the Earth having both access to the rich southern skies and a suitable climate, Southern Africa holds a favoured position in astronomy. Consequently it has seen the establishment of a number of professional observatories engaged in research while many individuals have become enthusiastic amateur astronomers. Planetaria and visiting nights at observatories convey to the general public much of what goes on in this field.

#### **OBSERVATORIES**

Boyden Observatory, situated at Mazelspoort, just outside Bloemfontein, is operated by an international consortium representing American, Belgium, Irish and South African interests. Thus many astronomers come from overseas to make use of its observing facilities which include the 1,5 m Rockefeller Reflector and the 0,9 m ADH Baker Schmidt. Its site offers good observing conditions, without being remote from a large centre.

On the other hand, the observatory sites in the hearts of Cape Town and Johannesburg have become unsuitable with the rapid expansion of those cities. The South African Astronomical Observatory a joint venture between the South African Council for Scientific and Industrial Research and the British Science Council-has merged their facilities and moved the larger instruments to a new site near Sutherland in the Karroo.

During 1974 the SAAO acquired the 1.9 m Radcliffe reflector located on the outskirts of Pretoria and the largest telescope in Southern Africa. As in Cape Town and Johannesburg observing conditon in Pretoria have deteriorated seriously in recent years and therefore, in order to reap the full benefits of the large telescope, plans are in hand to move the instrument and its turret to the Sutherland site. Dismantling of the equipment in Pretoria started in mid-October and the base of the new building should be ready to receive the telescope in February 1975. Re-erection and optical testing will then take at least another six months.

The SAAO in Johannesburg also has a small outstation near Hartebeespoort, sharing the site with the Leiden Observatory Southern Station which has a 0,9 m "light collector" reflector.

In the field of radio astronomy, the 25 m dish of the Deep Space Tracking Station near Krugersdorp is used for research work when not required for tracking spacecraft, while the Rhodes University Radio Observatory just outside Grahamstown, has a number of arrays for receiving radio emission from the planet Jupiter.

In addition to the professional observatories listed above, South Africa and Rhodesia have numerous private observatories, built and operated by amateur astronomers.

### **OBSERVATORIES OPEN TO THE PUBLIC**

Visiting nights at Boyden Observatory are held twice per month usually around the time of first quarter. Intending visitors should contact the Information Office in Hoffman Square, Bloemfontein for tickets (gratis). Numbers are restricted to twenty persons on each visiting night.

SAAO Headquarters, Observatory, Cape are open to visitors on the second Saturday each month at 8.00 p.m. No tickets are necessary but parties of more than ten persons should contact the observatory in advance.

#### PLANETARIA

The major planetarium in South Africa is that situated on the grounds of the University of the Witwatersrand (entrance in Yale Road - alongside M1). It is equipped with a highly complex Zeiss



projector and seats over 400 persons.

A smaller planetarium, with a Spitz projector and seating approximately 70, 1s located within the South African Museum, Cape Town. Shows are given each Saturday at 3.00 p.m., each Sunday at 3.30 p.m., and at 11.00 a.m. and 3.00 p.m. on public and school holidays. Further information can be obtained by phoning the museum at 41-2668.

## **TEACHING DEPARTMENTS**

Both the University of the Orange Free State and the University of Cape Town have departments of astronomy - concerned with optical astronomy - while the Physics department of Rhodes University specialises in radio astronomy. The UOFS department is incorporated into Boyden Observatory and is headed by the director, Prof. A.H. Jarrett. Prof. Brian Warner occupies the chair of Astronomy at UCT, his department makes use of the SAAO observing facilities at Sutherland. The Physics Department at Rhodes has its own radio observatory outside Grahamstown.

## THE ASTRONOMICAL SOCIETY OF SOUTHERN AFRICA

The Astronomical Society of Southern Africa is a body consisting of both amateur and professional astronomers. Membership is open to all interested persons, regardless of knowledge or experience. In addition to this handbook, the Society issues twelve numbers of "The Monthly Notes of the Astronomical Society of Southern Africa" (MNASSA) each year. Members also receive copies of "Sky and Telescope", an excellent and very popular monthly magazine published in the United States. It provides up to date information on both professional and amateur activities, together with news of space research and other related subjects. The society's annual subscription is R8.00 and there is an entrance fee of R2.50. Information can be obtained from the Honorary Secretary, Astronomical Society of Southern Africa, c/o The South African Astronomical Observatory, P.O. Box 9, Observatory 7935, Cape.

## LOCAL CENTRES OF THE SOCIETY

Autonomous local Centres of the Society hold regular meetings in Cape Town, Bloemfontein, Durban, Johannesburg, Pietermaritzburg and Pretoria. Visitors are very welcome at meetings and may, if they wish, join a centre, without becoming a full member (i.e. receiving publications for R8 subscription) of the society.

CAPE CENTRE (Cape Town) - Meetings on 2nd Wednesday of the month (except Jan. and Dec.) at the South African Astronomical Observatory at 8.00 p.m. The Centre possesses a small observatory housing the twelve inch Ron Atkins Telescope. There is also an active occultation section. Secretarial address: 4 Chalfont Road, Newlands 7700. Information on meetings also available from Department of Astronomy, U.C.T. Telephone (day time) 698531 ext. 256.

NATAL CENTRE (Durban) - Monthly meetings are held at the Teacher's Centre. Occultation and telescope making sessions. The Centre has a portable twelve inch reflecting telescope. Secretarial address: 3 Bruce Crescent, Westville 3630. Telephone (evenings) 858053.

NATAL MIDLAND CENTRE (Pietermaritzburg) - Meetings usually on the first Thursday of the month either in a private home or at the Royal Hotel in Burger St. Secretarial address: 17 Yalta Rd., Pietermaritzburg 3201.

FREE STATE CENTRE (Bloemfontein) - The Centre has its own observatory at Brandkop. Meetings on 2nd Thursday of the month at the homes of members. For further information, contact Mr. G.J. Muller, 35 Wilcocks Road, Bloemfontein 9301. Telephone (evenings) 7-3442 or Mr. J. Rhodes, Telephone 7-1981 (day time). Associate members should endeavour to have communications in the hands of the secretary a week before the monthly meetings.

PRETORIA CENTRE - The Centre has its own observatory containing a twelve inch reflecting telescope. For information contact Mr. K.J. Sterling, 5 Hekla Road, Valhalla 0137 - Phone 713272. The Centre has a small Observatory housing a 30 cm f9 telescope.

## **OBSERVING SECTIONS OF THE SOCIETY**

These sections exist to coordinate and encourage constructive observing programmes. Mention of the type of observations and equipment involved are made in the appropriate parts of this handbook together with the names and addresses of the directors.

see	page	30
see	page	48
see	page	40
see	page	43
see	page	41
	sce see see see	sce page sce page see page see page see page

# THE CHANGING NIGHT SKY 1975

The annual cycle of the starry skies visible in the evening forms the backdrop to the continually changing configurations of the Moon and planets. The maps in this section are highly simplified – they show only the visible planets and the most easily identified constellations – and are intended as an identification aid for the reader. The edge of each circular map approximately represents the horizon, the centre the point overhead. The maps are best used by holding the booklet up, rather than laying it flat, and rotating it to match the direction the reader is looking. If in doubt of direction – the best place to start is where the Sun went down – which obviously must be somewhere on the western horizon.

## **EXPLANATION OF THE APPARENT MOVEMENT OF THE SKY**

The most convenient representation of the sky is that of a sphere viewed from the inside by an observer at its centre. This is illustrated in the diagram below. If, as is usual, the observer stands on a flat surface (a local portion of the Earth's surface can be considered flat) - then his horizon divides the sphere into two halves - one which he can see, the other obscured by the ground beneath his feet.



The Earth, of course, rotates - but the observer is not aware of his rotation - and rather thinks that the whole sky is rotating instead! The sky therefore appears to rotate about the axis marked SN in the diagram - an axis parallel to the Earth's axis of rotation. Most of the sky seen from Southern Africa rises above the eastern horizon and sets below the western horizon, but a portion above the southern horizon never sets and a portion below the northern horizon never rises.



The two planets Venus and Jupiter make a spectacular pair above the western horizon in the early evening - they pass within 092 of each other on February 17. The most prominent constellation is Orion - easily identified by the three bright stars in a straight line which make Orion's belt. The nearby constellation of Taurus contains the best known star clusters, the Pleiades and the Hyades. In the south the kite shaped Southern Cross is easily recognised. More detailed maps of both the Orion and Southern Cross regions, together with lists of objects for observation by naked eye, binoculars and small telescope, are given on pages 38 and 39. This is the best time for observing Saturn which is at opposition on January 6 - detailed information on the planets is given on page 22.



Venus is prominent above the western horizon in the early evening and moves eastward amongst the stars of Taurus. Above Taurus, Orion is conspicuous (Map on page 38), although he will soon draw closer to the horizon and disappear from the evening sky. Saturn, which appears to move considerably slower than Venus, remains in Gemini. Leo is now high in the sky and Spica, the brightest star in Virgo is above the eastern horizon. Close to Spica is the planet Uranus (at opposition on Apr 21) which can be found with the aid of a small telescope and the map on page 29. The Southern Cross (map on page 39) is now climbing up towards its highest point.



Venus continues to grow brighter in its now familiar position above the western horizon. As it moves eastward against the starry background it passes 3° north of Saturn on May 24 and 4° south of Pollux on May 31. The angle between Venus and the Sun reaches maximum on Jun 18 and thereafter decreases. In the south the Southern Cross (map on page 39) reaches its highest point. Although Orion and neighbouring regions have now disappeared, the dominant winter constellation, Scorpius, has risen in the east. A more detailed map of the Scorpius region is given on page 40. The planet Neptune (at opposition on June 1) lies in the Scorpius region but is too faint to be seen with the naked eye. It can however be found with a small telescope and the map on page 29.



Although attaining greatest brilliancy on Jul 22, Venus now appears to slide back towards the setting Sun. It passes 0,4 south of Regulus on Jul 9 and disappears from the evening sky in late August. The Scorpius Region (map on page 40 ) is high overhead. In the south the Southern Cross (map on page 39) is somewhat passed its highest point.



During September, Mercury can be seen low over the western horizon - see map on page 23. Scorpius (map on page 40) is now progressing towards the western horizon, while Jupiter makes an appearance above the eastern horizon. This planet makes a fine object for a small telescope (see page 27) - four of its moons are easily visible and their configurations change from night to night (see page 32). Jupiter is at opposition on Oct 13. Neptune in the Scorpius region is occulted by the Moon on Sept 11 - see page 29.



After a six month absence Orion (map on page 38) makes a reappearance in the evening sky. He is seen low above the eastern horizon, with the star clusters of neighbouring Taurus somewhat higher. Jupiter (see pages 27 and 32) is overhead but its brightness is nearly matched by Mars - also above the eastern horizon - which is at opposition on December 15. An eclipse of the Moon takes place late in the evening of November 18 and finishes in the carly hours of the morning.

# **THE SUN 1975**

## **BASIC DATA**

Diameter: 1 392 000 km (109 times Earth diameter) Mass: 1,99 x  $10^{30}$  kg (330 000 times Earth Mass) Surface Temperature: Approx.  $6000^{\circ}C$ Temperature at centre: Approx. 10 million<sup>o</sup>C

The Sun is our nearest star. It is composed chiefly of hydrogen and is in a gaseous state throughout. So hot and dense is its interior that nuclear reactions occur there - thus producing the energy that is eventually radiated from its surface. At times its surface is disturbed by sunspots (which may persist for some weeks) and flares (short lived).

The Earth's orbit around the Sun is not quite circular. In 1975 we will be closest to the Sun on January 3 (perihelion - approx. distance 147 million km.) and furthest from the Sun on July 6 (aphelion - approx. 152 million km.) During the year, the Sun <u>appears to us</u> to make a complete circuit of the sky (i.e. relative to the starry background) as indicated in the diagram.



Permanent damage to the eye can be caused by looking directly at the Sun. The diagram below shows how a small telescope (or half a binocular) may be used to project an image of the solar disk onto a piece of white card. It may also be advisable to stop down the telescope aperture so that the eyepiece is not damaged by the intense light passing through it. Tiny black sunspots are generally visible on the otherwise white solar disk - if monitored over a period of a week or so, the rotation of the Sun should be apparent.



## The Sun 1975

			CAPE	TOW	/N		DUR	BAN		B	LOEMF	ONT	EIN <sup>–</sup>	J	HANN	ESBU	RG	;	SALISE	URY	
		SUI	NRISE	SUI	SET	SUI	VR ISE	SUI	NSET	SUI	NRISE	SUI	SET	SUI	RISE	SUL	SET	SUN	RISE	SUN	SET
Jan	1	05 <sup>h</sup>	38 <sup>m</sup>	20 <sup>h</sup>	01 <sup>m</sup>	04 <sup>h</sup>	58 <sup>m</sup>	19 <sup>h</sup>	01 <sup>m</sup>	05 <sup>h</sup>	21 <sup>m</sup>	19 <sup>h</sup>	18 <sup>m</sup>	05 <sup>h</sup>	18 <sup>m</sup>	19 <sup>h</sup>	04 <sup>m</sup>	05 <sup>h</sup>	24 <sup>m</sup>	18 <sup>h</sup>	35 <sup>m</sup>
	11	05	46	20	02	05	06	19	02	05	29	19	18	05	25	19	05	05	29	18	37
	21	05	55	19	59	05	14	19	00	05	37	19	17	05	33	19	04	05	37	18	38
Feb	1	06	07	19	52	05	24	18	55	05	46	19	13	05	42	19	00	05	42	18	36
	11	06	17	19	44	05	32	18	48	05	54	19	06	05	49	18	55	05	47	18	32
	21	06	26	19	33	05	41	18	39	06	02	18	57	05	56	18	47	.05	52	18	27
Mar	1	06	33	19	23	05	46	18	30	06	08	18	48	06	00	18	39	05	55	18	21
	11	06	41	19	11	05	53	18	19	06	13	18	38	06	06	18	29	05	57	18	15
	21	06	49	18	58	05	59	18	08	06	18	18	27	06	11	18	19	06	00	18	06
Apr	1	06	58	18	41	06	06	17	53	06	25	18	13	06	17	18	06	06	02	17	57
	11	07	04	18	30	06	11	17	43	06	30	18	03	06	21	17	56	06	04	17	50
	21	07	13	18	17	06	17	17	31	06	35	17	52	06	25	17	47	06	07	17	43
May	1	07	20	18	05	06	24	17	22	06	42	17	44	06	31	17	38	06	10	17	37
	11	07	28	17	57	06	31	17	14	06	49	17	36	06	37	17	31	06	13	17	32
	21	07	34	17	50	06	36	17	08	06	54	17	30	06	41	17	26	06	16	17	29
Jun	1	07	43	17	45	06	43	17	04	07	01	17	27	06	47	17	23	06	20	17	28
	11	07	48	17	44	06	48	17	03	07	05	17	26	06	52	17	22	06	23	17	27
	21	07	51	17	44	06	51	17	04	07	08	17	27	06	55	17	24	06	26	17	29
Jul	1	07	53	17	48	06	53	17	07	07	10	17	30	06	57	17	27	06	27	17	32
	11	07	51	17	52	06	51	17	11	07	08	17	34	06	55	17	30	06	27	17	35
	21	07	47	17	58	06	48	17	16	07	05	17	39	06	53	17	35	06	26	17	40
Aug	1	07	39	18	06	06	42	17	22	07	00	17	45	06	48	17	41	06	23	17	42
	11	07	30	18	13	06	34	17	29	06	53	17	51	06	41	17	46	06	18	17	46
	21	07	19	18	20	06	24	17	35	06	42	17	55	06	32	17	50	06	11	17	48
Sep	1	07	06	18	27	06	12	17	40	06	31	18	01	06	21	17	54	06	04	17	49
	11	06	52	18	34	06	00	17	46	06	19	18	06	06	11	17	59	05	55	17	51
	21	06	38	18	41	05	48	17	51	06	07	18	10	05	59	18	03	05	46	17	52
Oct	1	06	25	18	48	05	37	17	57	05	57	18	16	05	50	18	08	05	39	17	54
	11	06	12	18	55	05	25	18	03	05	45	18	22	05	39	18	12	05	30	17	57
	21	05	58	19	04	05	12	18	09	05	33	18	27	05	27	18	17	05	23	17	59
Nov	1	05	46	19	13	05	02	18	17	05	24	18	35	05	19	18	24	05	16	18	03
	11	05	38	19	23	04	55	18	26	05	17	18	44	05	13	18	32	05	14	18	08
	21	05	31	19	33	04	49	18	34	05	12	18	52	05	08	18	39	05	11	18	13
Dec	1	05	29	19	43	04	48	18	42	05	11	19	00	05	07	18	46	05	12	18	19
	11	05	28	19	50	04	48	18	50	05	11	19	07	05	08	18	53	05	14	18	25
	21	05	32	19	57	04	52	18	57	05	15	19	14	05	12	19	00	05	18	18	31

## TIMES OF SUNRISE AND SUNSET

#### SOLAR ECLIPSES

During 1975 there are two partial solar eclipses but neither is visible from Southern Africa. The first, on May 11, is visible in the artic and Europe; the second, on November 3, is visible in the antarctic and southern South America.

# THE MOON 1975

## **BASIC DATA**

Diameter: 3 480 km (0,27 of Earth) Mass: 7,35 x  $10^{22}$  kg (1/81 of Earth) Surface Gravity: 0,16 of Earth Average distance from Earth: 384 000 km

## PHASES AND VISIBILITY



### The Moon 1975

### THE MOON'S ORBIT



Dates of Apogee, when the Moon is furthest from the Earth (approx. 407 000 km) and of Perigee, when the Moon is closest to the Earth (approx. 357 000 km) are given below.

		III OGLL						1 1/1001	, L		
Jan	15	May	5	Sep	20	Jan	28	Мау	20	Oct	4
Feb	12	Jun	2	Oct	17	Feb	26	Jun	15	Nov	2
Mar	11	Ju <b>n</b>	30	Nov	14	Mar	26	Jul	11	Nov	30
Apr	7	Jul	27	Dec	11	Apr	23	Aug	8	Dec	26
		Aug	24					Sep	6		

As a result of its motion around the Earth, the Moon appears to make a complete circuit of the heavens in just under a month. It occasionally passes in front of bright stars (details given in Occultation section - page 43) and close to visible planets (details given in Changing Night Sky section - page 4)

### TIMES OF MOONRISE AND MOONSET

ADOGET

Times for Cape Town, Johannesburg and Durban are given on pages 16 to 21. Times for other places in Southern Africa can be roughly estimated from the predictions for these three stations.

#### LUNAR ECLIPSES

During 1975 there will be two total lunar eclipses but the first, on May 25, will not be visible from Southern Africa. The second, on November 18/19, will be visible and details are as follows:

Moon enters umbra	Nov	180	<sup>1</sup> 22 <sup>1</sup>	38,6	(at	58 <sup>0</sup>	East	of	North	point)
Total eclipse begins		19	00	02,6						
Middle of eclipse		19	00	23,4						
Total eclipse ends		19	00	44,1						
Moon leaves umbra		19	02	08,2	(at	730	West	of	North	point)

DERIGEE

## THE SURFACE OF THE MOON

In common with the inner planets of our solar system, the Moon's surface suffered bombardment by numerous minor bodies during the period 4,5 to 3,0 billion years ago. This has produced the heavily cratered topography now visible. Some particularly large impact caused large circular depressions, which were flooded by molten lava from the Moon's interior. These are the maria basins which appear smoother and darker then the rest of the surface (the latin words mare and maria come from older times when they were mistaken for seas). The maria surfaces, being younger have fewer large craters, but the entire surface is peppered with tiny craters produced by tiny bodies which have also served to plough up the ground thus forming the regolith - a layer of loose material a metre or so deep.

The Moon 1975





Jan 16 Feb 12 Mar 12 Apr 8 May 5 Jun 1/28 Jul 26 Aug 22 Sep 18 Oct 15 Nov 11 Dec 8

R)A

Jan 2/30 Feb 26 Mar 25 Apr 22 May 19 Jun 15 Jul 12 Aug 8 Sep 5 Oct 2/29 Nov 25 Dec 23 Dates of Maximum Exposure of Indicated Limbs Jan 23 Feb 20 Mar 20 Apr 17 May 13 Jun 9 Jul 6 Aug 3/31 Sep 28 Oct 26 Nov 22 Dec 19

Jan 7 Feb 4 Mar 4 Apr 1/29 May 27 Jun 23 Jul 19 Aug 15 Sep 12 Oct 11 Nov 8 Dec 6



MOONSET
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## The Moon 1975

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Predictions supplied by H. M. Nautical Almanac Office (British Science Research Council)

# THE PLANETS 1975

#### BASIC DATA

	Dist from Sun	Period of Revolution	Mass	Diameter	Rot <b>ation</b> Period	Inclination of Equator
	$10^{6}$ km	years	(Earth = 1)	10 <sup>3</sup> km		to Orbit
Mercury	58	0,24	0,056	4,98	59d	?
Venus	108	0,62	0,817	12, 4	<sup>244</sup>	?
Earth	150	1,00	1,000	12,8	23"56"	23 27'
Mars	228	1.88	0,108	6,76	24 37	23 59
Jupiter	778	11,9	318,0	142,7	09 51	03 04
Saturn	1426	29,5	95,2	120,8	10 14	26 44
Uranus	2868	84,0	14,6	47,1	10 49	97 53
Neptune	4494	164,8	17,3	44,6	14 ?	28 48
Pluto	5896	247,6	0,9?	2	6d?	?

#### GENERAL

Apart from Uranus, Neptune and Pluto, the planets of our solar system are amongst the brightest objects in the night sky. Unlike the distant stars, their relative positions do not remain fixed, but continually change as, like the Earth, they orbit around the Sun. Their apparent movements against the starry background are complicated as they result from a combination of their own motion and the Earth's motion. Their brightnesses also vary considerably, as both their distances from the Earth and the visible portions of their sunlit hemispheres change. Since the period of a planet increases with increasing distance from the Sun, so we find that the inner planets - Mercury and Venus - appear to "overtake" the Earth in their orbits, while the Earth in turn "overtakes" the outer planets - Mars, Jupitor and Saturn. The terms given in astronomy to the various Sun-Earth-Planet configurations are illustrated in the accompanying diagram. Dates of such configurations occurring in 1974 are listed chronologically in the Changing Night Sky Section (page 4) and are also mentioned in the text below.



## VISIBILITY AND IDENTIFICATION

To determine whether a particular planet is visible, the diagram for times of rising and setting (pages 24 and 25) should be consulted. Quite often the time of rising or setting is sufficient to identify a planet, but the sky maps (pages 5 to 10) can also be used. Where the information given in those maps is still considered insufficient, further finding aids are provided in the information for individual planets below. Planets can often be distinguished from stars as, due to their greater angular diameters, they do not twinkle.

## **OBSERVING THE PLANETS**

To the naked eye, planets appear as virtually pinpoint sources of light. However, their disks can be readily resolved with the aid of a small telescope. Even so, their angular diameters are of the order of 10 seconds of arc - roughly 1/200 of the Moon's angular diameter - so it is not always possible to distinguish details on their disks. The disks of Mcrcury and Venus are only seen fully illuminated when they are furthest from us - as they draw closer, their disks grow larger but the phase changes to a crescent as we see more of their dark hemispheres. In contrast, the disks of the outer planets are always seen fully or near fully illuminated.

#### MERCURY

The innermost planet, Mercury, revolves around the Sun faster than any other planets. To the Earth it appears to make slightly more than three complete circuits in 1975:

Greatest Eastern Elongation	Jan 23 (19 <sup>0</sup> )	May 17 (22 <sup>0</sup> )	Sept 14 (27 <sup>0</sup> )
Inferior Conjunction	Peb 3	Jun 10	Oot 9
Greatest Western Elongation	Mar 6 (37 <sup>0</sup> )	Jul 4 (22 <sup>0</sup> )	Oot 25 (18 <sup>0</sup> )
Superior Conjunction	Apr 18	Aug 1	Nov 38

As the diagram for rising and setting times (pages 24 and 25) shows, Mercury is difficult to observe - due to its close proximity to the Sun. The best early evening period is near the September Eastern Elongation and a finding diagram is provided here.

Mercury appears to pass close to other planets and bright stars as follows:

Near	Jupiter:	Apr	6	$(1,0S)^*$
Near	Aldebaran:	May	11	(8 <sup>0</sup> N)
Near	Regulus:	Aug	11	(1, 2N)
Near	Venus:	Aug	16	(9 <sup>0</sup> N)
Near	Spica:	Sep	24	(1,88)
	•	Sep	29	(1,8S)

(\* indicates visible only in morning sky)

The angular diameter of Mercury's disk rarely exceeds 10 seconds of arc so it is difficult to make out any features on the disk. Close up photography by the Mariner 10 spacecraft has shown that the planets surface is incredibly similar to that of the Moon.



(will be higher than shown here as both time of evening and date progress)

## TIMES OF RISING AND SETTING

To find the times of rising and setting of the planets for any particular night of the year, place a ruler (or suitable edge) vertically on the diagram according to the date given by the horizontal scale. The intersection of the ruler with the lines of rising and setting then give the appropriate times which can be read off against the vertical scale (24 hour clock on the left hand side, or conventional 12 hour clock on the right hand side).



## The Planets 1975

The times of rising and setting given by this diagram are accurate for position  $30^{\circ}$  East,  $30^{\circ}$  South, and approximately correct for other places in Southern Africa. Strictly speaking, corrections for latitude and longitude should be applied, but the latitude correction is, in general, sufficiently small to be ignored and in no case will exceed 15 minutes. Longitude corrections are given on page 55.



### The Planets 1975

#### VENUS

Over the period January to August, Venus is the most prominent object (aside from the Moon) in the early evening sky. During January to mid March, it is accompanied by Jupiter and the two make a spectacular pair as they pass within 0,2 on the evening of February 17. Venus reaches greatest elongation on June 18 and greatest brilliancy (magnitude -4, 2) on July 22. With the aid of a small telescope, its apparent diameter is seen to grow and its phase to change as its distance to the Earth diminishes - see diagram below. The opaque clouds of Venus' atmosphere cause the sunlit portion of the disk to appear brilliant white. Ultra violet photography, particularly that from the Mariner 10 spacecraft, have revealed cloud structure, but the actual surface is still to be seen.



On August 27, Venus passes between the Earth and the Sun (Inferior Conjunction) and thereafter appears in morning sky (Greatest brilliancy on Oct 3, greatest elongation on Nov 7).

Venus appears to pass close to other planets and bright stars as follows:

Near Jupiter:	Feb	17 (0, 2S)		
Near Aldebaran:	Apr	22 (7 N)		
Near Saturn:	May	24 (3 <sup>0</sup> N)		
Near Pollux:	May	31 (4 <sup>o</sup> S)		
Near Regulus:	Jul	9 (0,4S)	Oct	6(4 S)
Near Mercury:	Aug	15 (9 <sup>0</sup> S)		
Near Spica:	Nov	29 (5 <sup>0</sup> N)*		
(* indicates visible only in	n the me	orning sky)		

#### MARS

Mars is "hidden" away in the morning sky for most of the year. Only in December does it rise early enough to be seen in the evening sky. It becomes a very conspicuous object in the night sky as the Earth draws past it. The two planets are closest on December 9 and opposition is on December 15. Its brightness (magnitude -1, 5) and its characteristic reddish colour should make identification easy.



(Mar's period of rotation is only slightly greater than the Earths - the hemisphere visible each evening changes slowly as indicated)

It is worthwhile waiting to near opposition before attempting telescope observations. Even then, the angular diameter is only of the order of 15 seconds of arc (as Mars is a small planet and this opposition is not favourable). With a good instrument and good seeing conditions it may be possible to glimpse the features indicated in the diagram. Spacecraft observations reveal that much of Mars is heavily cratered - like the Moon and Mercury - while lava from conical volcanos has also covered a portion of the surface. Much loose sand is blown by the thin but ferocious Martian winds. In common with the Earth's desert scenery, it has sand dunes and dry river beds. A canyon on Mars completely dwarfs Earth's Grand Canyon. No water is visible at present and may lie under the surface - the polar caps are believed to be thin layers of solid carbon dioxide.

Mars appears to pass close to other planets and bright stars as follows:

Near Jupiter:	Jun	16 (0, 5S)	both only visible in meaning she
Near Aldebaran:	Sep	1 (4 <sup>0</sup> N)	both only visible in morning sky

#### JUPITER

In January and February Jupiter will accompany Venus in the early evening sky. The two planets will make a spectacular pair as they pass closest to one another on the evening of Feb. 17. Thereafter Jupiter will draw too near to the Sun for observation. After it passes round the far side of the Sun (relative to the Earth) on Mar. 22 it reappears in the morning sky. As the year progresses, it rises earlier and earlier to become a conspicuous evening object in the latter part of the year. It is at its brightest (magnitude -2, 5) at opposition on October 13.

Other planets pass close	to Jupiter as follows:
Venus:	Feb 17 (0 <sup>0</sup> <sub>1</sub> 2S)
Mercury:	Apr 6 $(1, 0S)^*$
Mars:	Jun 16 (0,5S)*
(* indicates visible only i	in morning sky)



Jupiter makes an excellent object for a small telescope. It is often possible to see features on the disk: dark and light cloud bands, running parallel to the equator, and spots, in particular the famous Great Red Spot. These are indicated in the diagram above. The Great Red Spot is not always visible because of the rotation of the planet.

Jupiter is composed chiefly of hydrogen and helium and it doubtful that the planet possesses any solid surface. Had it had more mass, nuclear energy generation would have started in the core and it would have become a small star - as it is, it is found to radiate more energy than it receives from the Sun so the interior is probably hot and slow convective motions in the outer liquid or near liquid layers probably account for the observed changes seen on the disk.

Jupiter possesses thirteen moons (the thirteenth was discovered in 1974), four of which are large and are clearly seen accompanying the planet. A section dealing with their observation is given on page 32.

#### **SATURN**

For the first half of the year, Saturn is visible in the evening sky (opposition on January 6); for the second half it is visible in the morning sky. As the map shows, it is to be found in the sky in the neighbourhood of Castor and Pollux, the twins of Gemini.



Venus passes 3<sup>0</sup> north of Saturn on May 24.

The famous rings of Saturn can be distinguished with the aid of a small telescope - their angular diameter is about 45 seconds of arc. As the diagram shows, they are presently tilted so that the north pole of Saturn is just hidden. Saturn possesses ten moons; details for observing some of them are given on page 36.



### URANUS

Uranus, at opposition on April 21, lies in the constellation of Virgo, near the bright star Spica (shown on the night sky maps, pages 6 to 8). It is on the borderline of naked eye visibility (magnitude 5,7) and may be found with the aid of binoculars and reference to the accompanying finding chart, which shows all stars in the region down to the same faintness. With a small telescope, its disk (angular diameter 4 seconds of arc) may just be distinguished.



#### **NEPTUNE**

Neptune lies in the constellation of Ophiuchus very close to the conspicuous constellation of Scorpius. It is far too faint to be seen with the naked eye - magnitude 7,7 at opposition on June 1 - but may be located using the finding chart below (which shows all stars down to magnitude 7,7) and a small telescope. Its angular diameter is only 2,5 seconds of arc, but its non-stellar appearance should aid in identification. Neptune is occulted by the Moon on September 11 (see page 47).



### **PLUTO**

Pluto lies in the Virgo/Coma region (at approximately R.A.  $= 12^{h}55^{m}$ , Decl.  $= +13^{0}$ ). Since it is very faint (14th magnitude), it can only be found using a large telescope and specially prepared finding charts.

# **COMETS AND METEORS**

### COMETS

Comets are celestial bodies moving around the sun, mostly in very elongated orbits. The typical comet consists of a nucleus surrounded by a hazy aura of gas and dust called the coma, and in many cases there is a tail stretching away from the sun. Faint comets, several of which are discovered each year, usually appear only as fuzzy patches without nucleus or tail.

Comets are the most mysterious and capricious of solar system objects and the nature of the physical changes which they exhibit is still not fully understood.

Observers with quite modest equipment (say a refracting telescope of not less than 7.5 cm) can do useful work by following known comets and reporting on their appearance. The ability to make accurate brightness estimates is especially useful and well worth cultivating.

Sweeping the sky for new comets, though requiring considerable patience and tenacity, is also well within the scope of the equipment mentioned.

Interested persons should contact the Director of the Comet and Meteor Section at 90 Malan Street, Riviera, Pretoria 0002.

#### METEORS

Meteors, or shooting stars, (resulting from small bodies entering the Earth's upper atmosphere) are generally seen in greater abundance after midnight (due to the direction of the Earth's motion) than in the early evening. There are two categories of meteors - the sporadic ones, and the showers. A meteor shower comes from a certain direction in space (the Radiant) and is thought to be associated with the remains of a comet. When the Earth passes close to the comet's original orbit, a shower can be expected. A list of these predicted showers is given in the table opposite.

Editor's note: Mr. J.C. Bennett, the Director of the Comet and Meteor Section, who contributed this section, is to be congratulated on discovering his second comet just as this handbook goes to press!

SHOWE RS	
METEOR	
PREDICTED	

	D	ate			Shower		Radi	ant	Ma	ximı	Ħ	TE	ansit	Recom	mended	Conditions at
				_		B	. A.	Dec.	Dat	eH	ourl	y (ap)	prox)	Time o	I Watch	Maximum
Mar	14 -	- Mai	r 18		Corona Australids	16 <sup>h</sup>	20 <sup>m</sup>	-480	Mar	16	Kate 5	4 <sup>4</sup>	15 <sup>m</sup>	02h30m	- dawn	Favourable
Mar	12 -	- Apt	r 25		Hydraids	12	16	-27	Mar	25	e.,	00 ]	10	22h	- 02h	Unfavourable
Apr	19 -	- Apr	24		April Lyrids	18	80	+32	Apr	22	15	04 1	5	03h	- dawn	Favourable
May	1 -	- Ma	y 80		Eta Aquarids	22	24	00	May	10	18	10	30	03h	- dawn	Unfavourable
Apr	20 -	Inf -	y 30		Sco-Sgr System	18	00	-30	Jup	14	¢.	00	30	23h	- 02h	Favourable
Jun	- 01	aut -	21		June Lyrids	18	32	+35	Jun	16	œ	01 (	00	00h	- 03h	Favourable
Jun	17 -	- Jun	26		Ophiuchids	17	20	-20	Jun	20	œ	23	30	21h	-02h	Unfavourable
Jul	10 -	- Aug	ιΩ		Capricornids	21	00	-15	Jul	25	œ	00	20	23h	-02h	Unfavourable
Jul	15 -	- Aug	5 15		Delta Aquarids	22	36	-17	Jul	27	35	02 ]	10	00h	- 04h	Unfavourable
Jul	15 -	- Aug	3 20		Pisces Australids	22	40	-30	Jul	30	IJ	02 ]	10	00h	-04h	Uniavou rable
Jul	15 -	- Aug	25		Alpha Capricornids	20	36	-10	Aug	63	10	00	00	20h	-24h	Favourable
lul	15 -	Aug -	24		lota Aquarids	22	32	-15	Aug	9	12	10 2	000	00h	-03h	Favourable
Oct	16 -	- Oct	27		Orionids	90	24	+15	Oct	21	35	04	30	02h	-dawn	Unfavourable
Oct	10 -	- Dec	<b>2</b>		Taurids	03	44	+14 +21	Nov	aD	16	00	20	400	-03h	Favourable
Nov	14	Nov -	, 20		Leonids	10	80	+22	Nov	11	10	90	30	02h	-dawn	Unfavourable
Dec	ŝ				Pinenicids	01	00	-55	Dec	Ω.	e.	20 1	10	21h	-23h	Favourable
Dec		- Dec	15		Geminids	20	28	+32	Dec	14	55	02 (	00	02h	-dawn	Favourable
Dec	- C	- Jan	2		Velaids	60	56	-51	Dec	29	e-	03	30	23h	-02h	Favourable

Comets and Meteors 1975

# THE MOONS OF JUPITER AND SATURN 1975

### **JUPITER'S MOONS**

One of the most popular sights for an observer with a small telescope is Jupiter and its moons. Four of Jupiter's thirteen moons are large enough to be seen very easily - they would even be visible to the naked eye were it not for the glare of the mother planet. In order of increasing distance from Jupiter, the four moons are Io, which orbits once around Jupiter in less than 2 days; Europa,  $3\frac{1}{2}$  days; Ganymede, 7 days; and Callisto which takes 17 days for a full circuit. All the orbits lie in Jupiter's equatorial plane and the system is seen almost edge on. As the moons circle Jupiter, they appear to oscillate from side to side alternatively passing in front of and behind the planet. Their configurations change from night to night and are shown in the diagrams on pages 34 and 35.

## EVENTS RELATED TO THE MOONS PASSING IN FRONT OF AND BEHIND JUPITER

The table below lists all events occurring between the end of twilight and just after midnight when the planet is above the horizon in Southern Africa.

Explanation of table:

Date and predicted times are given - these are for mid-phenomena and are not instantaneous.

The	moons	concerned	are	Ι	-	Io	ш	-	Ganymede
				п	_	Europa	IV	-	Callisto

Phenomena - the abbreviations used are:

Ec.	-	Eclipse: the	satellite passes th	rough the sha	dow of Jupiter	D -	Disa	ppearance
Oc.	-	Occultation:	the satellite is obs	scured by the	disk of Jupiter	R -	Reap	pearance
Tr.	-	Transit: the	satellite crosses t	he disk of Jup	biter	Ι -	Ingre	88
Sh.	-	Shadow transit	t: the shadow of	the satellite to	ansits the disk	E -	- Egre	88
		h m		h m		h	m	
Jan	1	21 04 I	Oc D July 21	24 23 II	Ec R   Aug 15	23 4	17 II	Oc R
	2	20 28 T	ThE	94 90 TT	0 n D 10	22 8	5 111	ጥ ተ ገ

	2	20	38	Ι	Tr E	i i		24	30	II	Oc D		19	23	55	п	Tr I
		21	48	I	Sh E		27	24	12	I	Ec D			24	23	I	Ec D
	9	21	29	I	Sh I	ļ	28	23	36	Ι	Sh E	1	20	22	45	I	Tr I
	10	21	14	II	Oc D			24	17	II	Ec D			23	45	I	Sh E
	12	20	36	11	Sh E			24	55	I	Tr E			24	53	I	Tr E
	16	21	22	III	Sh E		30	24	34	II	Tr E		21	22	14	I	Oc R
	19	20	28	п	Sh I	Aug	1	23	38	п	Ec R		26	22	55	пі	Sh I
		21	16	п	Tr E		4	23	19	1	Sh I	1	27	23	28	I	Sh I
	24	20	58	IV	Sh I			24	38	I	Tr I			24	32	I	Tr I
	25	21	10	I	Tr E		5	24	07	I	Oc R		28	24	01	1	Oc R
	26	21	16	II	Trl		6	24	29	п	Tr I		29	23	53	II	Ec D
	28	20	56	II	Ec R			24	29	11	Sh E		31	21	38	II	Sh E
Feb	1	20	56	I	Trl		8	24	56	III	Ec D			23	35	II	Tr E
	9	20	07	I	Oc D	1	13	23	04	I	Tr E	Sept	4	22	41	I	Ec D
								24	23	п	Sh I	_					

## The Moons of Jupiter and Saturn 1975

		h	m					ł	n m					1	1 m			
Sep	5	22 0	01 I	Sh	E	Oct	12	23	51	I	Sh I	Nov	20	21	30	I	Τr	I
		22 5	54 I	Tr	E			23	53	I	Tr I			22	01	ш	Τr	E
	6	22 8	5 <b>9</b> I.	ll Oc	R		13	21	12	I	Ec D			22	25	I	Sh	I
	7	21 3	3 <b>3 I</b>	í Sh	I			23	23	I	Ec R			23	15	III	Sh .	I
		23	21 I.	l Tr	· I		14	20	28	I	Tr E			23	41	I	$\mathbf{T}\mathbf{r}$	E
		<b>24</b> 1	14 I	I Sh	E			20	31	I	Sh E			24	36	I	Sh	E
	11	24 3	35 I	Ec	D		16	23	47	II	Tr E		21	21	56	I	Ec	R
	12	22 3	31 1	Tr	I			23	55	и	Sh I		24	24	20	11	Tr	I
		23 3	56 I	Sh	E		18	20	<b>4</b> 4	II	Ec R		26	23	11	H	Ec	R
		24 4	40 I	Tr	E		19	19	43	III	Ec R	ĺ	27	23	06	п	Τr	I
	13	21 (	02 I	ll Ec	D		20	22	56	I	Oc D	1		23	19	1	Tr	I
		22 (	1 00	Oc	R		21	20	02	I	Tr I	Í		24	20	I	Sh	1
		23 3	39 T.	II Ec	R			20	14	I	Sh 1		28	20	38	I	Oc	D
		<b>24</b> 1	19 I	II Oc	R			22	12	I	Tr E			23	51	1	Ec	R
	14	24	10 I	I Sh	I			22	26	I	Sh E		29	21	00	I	Sh J	E
	16	22	19 I	I Oc	R		22	19	47	I	Ec R	Dec	3	20	52	п	Oc	D
	19	23 3	39 I	Sh	I		25	20	04	II	Oc D	[	5	20	47	П	Sh )	Ē
		24	16 I	Τr	I			23	21	11	Ec R	1		22	27	I	Oc	D
	20	20	59 1	Ec	D		26	19	56	ш	Oc D		6	20	45	I	Sh J	ſ
		23 4	45 I	Oc	R			23	44	ш	Ec R			21	47	I	Tr	Е
	21	20	19 I	Sh	E		27	24	40	I	Oc D			22	56	I	Sh 1	E
		20 \$	51 1	Тr	E		28	21	46	I	Tr I		8	21	26	п	Ec	D
	23	20	57 I	I Ec	D			22	10	I	Sh I			23	53	111	Ec	R
		24	35 I	I Oc	R			23	56	I	Tr E	ł	10	23	20	11	Oc	D
	25	19 3	38 I	I Tr	·E			24	21	I	Sh E		12	20	49	II	Sh ]	I
	27	22	53 I	Ec	D		29	21	42	I	Ec R			21	00	Π	Τг	E
	28	20	02 I	Sh	I	Nov	1	22	19	п	Oc D			23	23	II	Sh )	E
		20	26 I	Тг	·I		2	23	13	Ш	Oc D			24	18	I	Oc	D
		22	13 I	Sh	E		3	19	59	II	Tr E		13	21	27	I	Tr	I
		22	35 I	Т	E			21	05	п	Sh E			22	41	1	Sh )	I
	29	19	55 I	00	R		4	23	31	I	Tr I			23	38	I	Τr	Ε
	30	23	33 I	I Ec	8 R			24	05	I	Sh I			24	52	I	Sh J	E
Oct	1	20 3	31 I	II Tı	·I		5	20	50	I	Oc D		14	<b>22</b>	11	I	Ec	R
		21	36 I	II Sh	E			23	37	I	Ec D		19	20	53	11	Tr	I
		22	37 I	п тı	• E		6	20	08	I	Tr E			23	26	II	Sh J	I
	2	19	19 I	I Tı	·I			20	45	I	Sh E			23	29	п	Τr	Ε
		21	22 I	I Sh	E		8	24	36	II	Oc D		20	23	20	I	Tr	I
		21	53 I	I Tı	·E		10	21	04	IJ	ShI			24	37	1	Sh 1	1
	4	24	48 I	Ec	D			22	16	п	Tr E		21	20	38	I	Oc	D
	5	21	56 I	Sh	I			23	41	Π	Sh E			24	06	I	Ec	R
		22	09 I	Τı	·I		12	22	36	I	Oc D		22	21	17	I	Sh 1	E
		24	08 I	Sh	E		13	20	29	I	ShI			24	06	ш	Oc	D
		24	19 I	Т	E			21	42	Ш	Sh E		26	21	52	HI	Sh 1	E
	6	21	39 I	00	2 R			21	54	I	Tr E			23	24	н	Τr	I
	8	23	02 I	11 Sh	1			22	41	I	Sh E		28	<b>22</b>	31	I	Oc :	D
		23	45 1	П Т1	• I		17	21	5 <b>9</b>	11	Tr I			23	06	11	Ec	R
	9	21	18 I	I Sh	I			23	41	11	ShI		29	21	02	I	Sh J	ĺ
		21	3 <b>3</b> I	I TI	: I			24	35	п	TrE			21	54	I	Tr	Е
		23	58 I	I Sh	Е		19	20	33	п	Ec R			23	13	1	Sh H	Ē
		24	07 I	ΙТι	۰E	i i		<b>24</b>	23	I	Oc D							

East DECEMBER ., Callisto 0 CHANGING CONFIGURATIONS OF JUPITER'S MOONS NOVEMBER ъ C Europa Ganymede OCTOBER Jupiter North South 0 0 0 0 SEPTEMBER 0 0 JANUARY 0 0 0 o West I 1 16 2 10 21 13 15 ŝ 14 3 ŝ Ξ

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They show the configurations at 9 p.m. on all Io is about to pass behind the planet, Europa in front. Ganymede is seen slightly to the Callisto is well to the east nearing its greatest apparent separation from the The The diagram at the top shows the orbits of the four bright moons of Jupiter - (as seen from the Earth) and their positions at evenings when Jupiter is clearly visible. The columns cover the months given, the horizon lines give the day of the month. wavy lines interpolate the apparent movement of the moons between successive diagrams. planet. The 150 or so small diagrams are miniature versions of the main diagram. east (but will also pass in front several hours later). 9 p.m. on the evening of January 1, 1975.

## SATURN'S MOONS



Saturn's moons are considerably fainter than the four Galilean moons of Jupiter. The diagram shows the orbits of four of Saturn's ten moons. The easiest to find is Titan (magnitude 8,5), according to the diagram and the information in the table below (which covers the period when Saturn is clearly visible in the evening sky).

Eastern Ele	ongation	Inferior Co	njunction	Western Eld	ongation	Superior C	onjunction
				Jan	4	Jan	7
Jan	12	Jan	16	Jan	19	Jan	23
Jan	28	Feb	1	Feb	4	Feb	8
Feb	12	Feb	16	Feb	20	Feb	24
Feb	28	Mar	4	Mar	8	Mar	12
Mar	16	Mar	20	Mar	24	Mar	28
Apr	1	Apr	5	Apr	9	Apr	13
Apr	17	Apr	21	Apr	25	Apr	29
Nov	27	Dec	1	Dec	5	Dec	9
Dec	13	Dec	17	Dec	21	Dec	25
Dec	29						

т	IT	A	N
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## THE STARS

#### CONSTELLATIONS

Apart from our Sun all the stars that we see are so incredibly distant that, despite their high speed velocities, their apparent positions change by only minute amounts each year. Consequently the patterns that they form appear unchanged. The Greeks and other ancient civilisations identified these patterns, or constellations, with various mythological characters and creatures, and most of the names they gave are still used today.

In all there are 88 constellations, roughly one half of which would be above the horizon at any one time. Some contain distinctive patterns of bright stars and are relatively easy to find; others are difficult to locate, even with suitable maps. The Changing Night Sky Section (page 4) gives information for finding some of the best known constellations - those easy to identify. Of those, the Southern Cross and Centaurus, Orion and Taurus, Scorpius and Sagittarius, are featured later in this section. Detailed information on other constellations is beyond the scope of this handbook and interested observers are advised to obtain a suitable star atlas.

#### STAR NAMES

Within each constellation, the brightest star is generally labelled  $\alpha$  (Alpha), the next  $\beta$  (Beta) and so on through the Greek alphabet. Most of the brightest stars also have their own names - usually of arabic origin. For example  $\alpha$  Canis Majoris, otherwise known as Sirius, is the brightest star in the constellation Canis Major.

### STELLAR MAGNITUDES

The apparent brightness of a star - which depends both on its true luminosity and its distance - is indicated by its magnitude. Equal intervals of magnitude represent equal ratios in light intensity. A star of magnitude 1,0 (typical of the brightest stars in the night sky) would be exactly one hundred times more luminous than a star of magnitude 6,0 (about the limit of visibility to the naked cye). The maps in this section show stars down to magnitude 4,5.

#### STELLAR DISTANCES

Distances are often expressed in units of light years - the distance light would travel in a year (equal to  $9.5 \times 10^{12}$  km).

#### DOUBLE STARS

It now appears that single stars such as our Sun are the exception, the majority of stars being double or multiple - two or more suns in orbit around one another.

#### STAR CLUSTERS

These are of two completely different sorts. Galactic clusters, having of the order of 100 stars, are found close to the plane of the Milky Way. The ones we can see are relatively nearby. Globular clusters are much larger and far more distant. They contain of the order of 100 000 stars each and are seen above and below the Milky Way on that side of the sky towards the centre of our galaxy. So great is their distance that small telescopes fail to resolve individual stars - instead they appear as fuzzy balls.

#### NEBULAE

Possibly one third of the matter in our region of the Galaxy is in the form of gas and dust (the remainder being contained in stars, plus a tiny amount in planets). Condensations of this material are called nebulae, some of which are illuminated by nearby stars while others are dark. They are usually referred to by their numbers in Messier's catalogue (M) or the New General Catalogue (NGC).

### THREE POPULAR REGIONS

The dominating constellation of the summer skies is Orion, that of the winter skies is Scorpius, while the Southern Cross is conspicuous for most of the year. They can be found with the aid of the maps in the Changing Night Sky section (pages 4 to 8). The regions around these constellations are also rich in interesting objects - visible either to the naked eye, or with the aid of binoculars or a small telescope - and are featured in the maps and text below. <u>It may be necessary to rotate the maps</u> to match the orientation of the constellations in the sky.

The Stars



(1) The constellation of Orion. The figure of the legendary hunter of Greek mythology is unfortunately upside down when seen from Southern Africa. The faint stars by  $\lambda$  represent the head,  $\alpha$  and  $\Upsilon$  the shoulders,  $\delta - \epsilon - \eta$  the belt, and  $\beta$  and  $\kappa$  the legs. Orion forms part of the "great hunting scene" in which he faces the onslaught of (2) Taurus, the bull. Only the forepart of the bull is depicted and, like Orion, it is upside down.  $\alpha$  and  $\epsilon$  are the eyes,  $\Upsilon$  the nose. Orion is accompanied by (3) Canis major, the large dog, and the small dog (off map) while Lepus, the hare, crouches at his feet.

(4) A section of the Ecliptic - a line encircling the entire sky and representing the plane of the Earth's orbit. As the Earth revolves around the Sun, the Sun <u>appears</u> to move along the ecliptic through the constellations of the Zodiac, of which Taurus is one.

(5) A portion of the Milky Way (looking out towards the edge of our Galaxy).

(6) Sirius - the brightest star in the night sky. It is somewhat brighter than our Sun and relatively close by at a distance of 9 light years. It is a double star but the companion is a white dwarf (only slightly larger than the Earth, and with a mass comparable to our Sun ) and is only visible through a large telescope.

(7) Betelgeuse - most famous of the red giant stars. Its diameter is of the order of the size of the Earth's orbit and its luminosity is nearly 10 000 times that of our Sun. Its red colour should be obvious to the eye. It is 520 light years distant.

(8) Rigel, despite being physically smaller than Betelgeuse, is more luminous (higher surface temperature bluish colour) and more distant.

(9) The stars in Orion's belt are distant hot blue stars.

(10) The Pleiades or Seven Sisters form the best known nearby star cluster. Six or seven stars are visible to the naked eye, binoculars or a small telescope show more.

(11) The Hyades is another nearby galactic cluster, but Aldebaran is not a member (it lies closer to us).

(2) The Great Nebula in Orion, just visible to the naked eye, shows up as a fan shaped mass of luminous gas through binoculars or a telescope. A telescope will also show a tiny "Trapezium" of four stars in the centre.

13 The Crab Nebula, the remnant of a supernova recorded by the Chinese in 1054, requires a moderate sized telescope for observation. In its heart is located the extraodinary pulsar which emits a double flash of light 30 times every second. The current belief is that it is a rapidly rotating neutron star - a star with the mass of our sum but with a diameter of only 10 km.

The Stars



 $\begin{pmatrix} 1 \end{pmatrix}$  Crux, the Southern Cross, is one of the most compact patterns of bright stars to be found in the sky. It lies on the border of that region of the sky which never sets as seen from Southern Africa.

 $\binom{2}{2}$  The two "Pointer" stars lie close to the Cross. (A similar pattern to the Southern Cross - called the False Cross - lies just outside and above the map, but has no accompanying pointer stars).

(3) The South Celestial Pole: This is one of two opposite points in space towards which the Earth's axis of rotation is directed. As the Earth rotates so the sky appears to pivot about this point. It always lies above the south point on the horizon, elevated by an angle equal to the observer's southern latitude. (The north celestial pole lies below the northern horizon and can never be seen from the Earth's southern hemisphere)

(4) The intersection of a line extended through the major axis of the Cross and the perpendicular bisector to the Pointers indicates the approximate position of the South Celestial Pole.

(5) a Centauri has the distinction of being the closest star to our solar system - at a distance of approximately 40 million million km or 4,3 light years. A small telescope readily shows that it is a double star - the two components take 80 years to revolve about one another. A much fainter third star also belongs to the system.

(6)  $\alpha$  Crucis can also be resolved as a double star by a small telescope (separation 5 sec of arc).

(7) The region indicated is one of the brightest sections of the entire Milky Way.

(8) The Large and Small Magellanic Clouds are the nearest of the oxternal galaxies (see also next section). They can be seen with the naked eye provided the sky is reasonably dark.

(9) The Great Looped Nebula - possibly the remnant of a supernova explosion - in the Large Magellanic Cloud. (Naked eye or binoculars).

(10) 'The "Coal Sack" - a dark mass of gas and dust obscuring a part of the Milky Way. (Naked eye or binoculars).

(1) Herschel's "Jewel Box" - a galactic cluster containing stars of different colours. (Small tolescope or binoculars).

(12)  $\omega$  Centauri and (13) 47 Tucanae arc perhaps the best known globular clusters. Binoculars will show their fuzzy appearance. (14) NGC 362 and (15) NGC 2808 are fainter globular clusters.

(16) NGC 3760 - a fine galactic cluster. (Binoculars or small telescope).

(17) The n Carinae nebula - site of a slow supernova that brightened to magnitude -0.8 in 1843 and is now of magnitude 6.4.

#### THE SCORPIUS REGION



(1) The constellation of Scorpius. The creature is depicted with  $\alpha$  in the centre of the body and  $\beta$  and  $\pi$  the claws. The distinctive tail  $\epsilon - \zeta - \theta$  curls round to the sting  $\lambda$ .

(2) Sagittarius - the figure of the centaur archer is very difficult to make out.

(3) A section of the Ecliptic. Like Taurus, Scorpius and Sagittarius are constellations of the Zodiac.

(4) The direction of the centre of our Galaxy - the Milky Way is that part of our Galaxy visible to us. Unfortunately the central nucleus is obscured by foreground gaseous and dusty matter - both dark and luminous hence the irregular shape of the Milky Way in this region. Luminous nebulae include (5) the Lagoon nebula and (6) the Omega nebula. These are best seen with the aid of binoculars.

( $\eta$ ) Antares - a distant red giant, several hundred times the diameter of our Sun - is so named because its red colour rivals that of the planet Mars.

8 β Scorpii can be resolved as a double star (separation 16 sec of arc) with a small telescope. In fact the brighter component is in itself a triple star, and the fainter component a double star!

This region includes a number of galactic clusters including (9) M7, (10) M8, (11) M4 and (12) NGC 6067. (Use binoculars or a small telescope).

Further from the plane of the Milky Way are some globular clusters: [13] M80 [14] M19 and [15] M22.

## NOVA SEARCHING

On rare occasions a star may undergo a nova outburst - its brightness increasing tremendously. Even observers having no equipment can perform a useful task in keeping a watch for such novae in an allocated area of the sky. Interested persons should contact the director of the recently revived Nova Search Section, Mr. J.C. Bennett, 90 Malan Street, Riviera, Pretoria 0002.

## The Stars

## VARIABLE STAR OBSERVING

The General Catalogue of Variable Stars lists some 20 000 stars. Professional observatories cannot possibly monitor all of these and the observation of variable stars is a field therefore in which amateurs can make real contributions to astronomical knowledge.

Of the 20 000 stars, at least 2 000 are suitable for monitoring by Southern Hemisphere observers but less than 200 are in fact observed from South Africa and a still smaller number receive adequate attention.

The Variable Star Section of the ASSA exists for the purpose of encouraging observers and of acting as a medium of communication. The Section disseminates incoming information amongst observers and will forward (on request) the observations of individuals to various variable star bodies. These include the American Association of Variable Star Observers and the Variable Star Sections of the British Astronomical Association and Royal Astronomical Society of New Zealand. These bodies combine the South African observers' light estimates with those from other parts of the world. The resulting light curves and tables are sent to a large number of professional observatories where astronomers are interested in investigating certain of the stars more fully.

In addition to the international work, the VSS of the ASSA supplies information direct to certain South African astronomers. It is in a position to warn observatories of sudden changes in certain "VIP" stars before the overseas bodies can do so.

These are variable stars to suit all levels of skill, types of equipment and tastes:

A few stars can be followed throughout their variation by the naked eye but the majority necessitate optical aid.

Some stars vary through a light range of more than 1000 to 1 (8 magnitudes) but others change less than a magnitude.

Flare stars should be observed several times a minute, certain other stars several times a night, some nightly and long period variables only need attention a few times a month.

Variable stars near the pole can be observed during convenient evening hours all the year round but those near the Ecliptic have to be observed at times near the horizon just after sunset or just before sunrise to give maximum coverage

Some stars can be located with great ease but others need a certain amount of skill to identify.

Certain stars are almost completely regular in their variations but others are quite unpredictable.

Amateurs who want to start observing stars in any of the above categories should contact the Director, Mr. M. D. Overbeek, 60 Edward Drive, Glendower, Edenvale, 1610. New observers will be given charts of a few easy objects and (if possible) a certain amount of instruction at the eyepiece. When writing, prospective observers should give brief details of their equipment.

# **OUR GALAXY AND OTHERS**

Our Sun is one amongst some 100 000 million other stars in our disk-shaped galaxy. As the diagram below shows, the Sun is situated towards the edge of the system and apparently orbits around the centre once in about 200 million years. We see the galaxy stretching round us as the Milky Way, but even with optical telescopes we can only see that portion closest to us as the re-mainder is obscured by dust and gas. However when we look above or below the plane of the Milky Way, we can see an indefinite distance into extragalactic space. Thus we have found the



universe to be populated by millions of other galaxies akin to ours. Our galaxy is a member of a rather poor cluster of galaxies which includes the Magellanic Clouds (our nearest neighbours) and the Great Galaxy in Andromeda (a large spiral that is very similar to our own galaxy). Spectral features in the light of distant galaxies are redshifted. This is usually interpreted as an overall expansion of the universe - the further one looks the faster the galaxies are receding.

# **ORDINARY OCCULTATIONS**

An occultation occurs when the moon passes in front of a star. The disappearance or reappearance of the star is an instantaneous event, or almost so, and is therefore easily timed. Time signals are broadcast continuously through the 24 hours from station ZUO (see page 54) and, radio propagation permitting, occultation timings can be made, either by eye and ear or by stop watch.

The essential record of a occultation observation consists of

- Date and observed time of the event (to a tenth of a second if possible). Universal Time is preferable.
- (ii) The Z.C. number of the star.
- (iii) Whether disappearance or reappearance.
- (iv) The observer's estimate of the accuracy of the observation.
- (v) Whether the time given in (i) has been corrected for personal error, and if so by how much.

Also required are the size and type of telescope used, the method of timing used, and the position of the telescope.

The telescope position must be to an accuracy of 1" or better in latitude and longitude and 30 metres or better in altitude. An observer can usually read his position off a 1/50 000 map published by the Director General of Trig. Survey, Mowbray, or his equivalent in other countries. Positions determined astronomically are NOT acceptable, not matter how carefully determined.

Timings of occultations are very valuable in determining the moon's shape and motion and consitute a field in which the amateur astronomer, often with only a small, unsophisticated telescope, can make an extremely useful contribution. Interested persons are urged to contact the Director of the Society's Occultation Section, Mr. A.G.F. Morrisby (c/o Dept. of Surveyor General, P. O. Box 8099, Causeway, Salisbury, Rhodesia).

Predictions of occultations of stars brighter than magnitude 7,5, supplied by H.M. Nautical Almanac Office, are reproduced on the next two pages. Explanation of Table:

Z.C. - the number of the star in the "Catalogue of 3539 Zodiacal Stars for the Equinox 1950.0" by James Robertson (U.S. Naval Observatory, 1939). A short index of the brighter stars is given on page 47.

Mag. is the visual magnitude of the star

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P Phase. 1 = Disappearance 2 = Reappearance
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El. of

- Moon the elongation of the Moon.  $0^{\circ}$  = New Moon,  $90^{\circ}$  = First Quarter,  $180^{\circ}$  = Full Moon,  $270^{\circ}$  = Last Quarter.
- U.T. The predicted time in Universal Time which is exactly two hours behind South African Standard Time. For example  $0^{h55,9}$  UT is  $2^{h55,9}$  a.m. SAST.
- a, b the approximate time of an occultation at a place  $\Delta\lambda$  degrees west and  $\Delta\phi$  degrees north of the city concerned is

Predicted time  $\div a.\Delta\lambda + b.\Delta\phi$ 

where a and b are in minutes of time.

- P the position angle on the Moon's limb, measured eastward from the north point.
- N no occultation A Moon at very low altitude
- S sunlight interferes G grazing occultation

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## **OCCULTATION PREDICTIONS**

					21.	0	APE 120	) MN	c		JOH/ o	UNESI	3URG	0		54 0	ALISB	URY	c
Dat	e	Z.C. No.	Мак.	Р.	of Moon	E. 13.	500, 8	3. 33.	900	E.	28.0	000,8	3, 26,	200	Ε.	31.	100, 5	3. 17	.700
						V.T.	a	b	P	U	.7.	a.	b	Ρ	U	.T.	8.	b	Р
Jan.	3 15 18 21 21	1623 3235 42 411 421	5.4 7.5 5.6 7.3 6.6	2 1 1 1	0 247 36 70 104 105	h 12 0 55.9 18 36.6 19 59.8 19 53.3 21 27.4	n -2.5 0.0	n -0.1 +2.3 +0.7	0 251 37 7 20 107	հ 1	m 12.5	та -2.1 А N II А	m -1.2	0 289	h 1	m 05.1	m -1.7 N N N A	_2.1	0 316
řeb.	23 24 31 1 4	709 861 1815 1949 2217	4.3 6.5 4.8 5.8 5.5	1 1 2 2 2	129 142 241 254 281	19 59.2 18 52.7 22 24.3 2 52.4	-2.4 -2.2 -0.6 M -0.1	+1.5 -1.0 -1.6 -4.4	51 108 295 345	19 22	10.4 15.1	N -2.6 -0.5 N N	-0.2 -2.2	92 323	19 21 22	21.3 52.2 50.4	N -3.0 +0.3 -1.6 N	+0.9 -3.8 +0.1	69 356 254
	7 16 17 18	2633 2638 244 371 485	4.0 5.4 6.9 6.4 6.9	2 2 1 1	316 316 61 73 84	2 08.5 2 55.9 18 16.7 19 36.7	-0.8 0.0 -1.2 -0.5 S	+0.8 -2.4 +0.2 -0.4	221 308 118 132	2 18 18	14.3 30.7 28.3	-0.6 N -0.8 A -2.2	-0.6 +0.8 +2.7	261 94 41	2 16	07.2 42.4	-0.3 Ν -0.δ Α Ν	-1.4 +1.4	293 68
	19 20 20 21 22	634 784 792 969 1109	5-3 6-2 5-1 7-1 7-3	1 1 1 1	97 109 110 123 136	20 20.9	N N N -2.2	-0.1	100	20 22 20	31.4 34.0 49.5	+0.3 ม ม -3.1	-2.3 +1.8	153 179 64	20 19 21 22	24.) 07.0 22.2 17.8	-0.7 -1.7 -0.1 -0.3 N	-0.3 -1.4 -1.3 -0.8	115 131 140 131
Mar.	27 28 28 1 3	1759 1884 1888 2017 2307	6.5 5.3 6.2 6.4 4.1	2 2 2 2 1	208 221 221 234 262	21 47.7 20 54.2 23 32.3	-1.9 N -0.2 N -0.2	-0.5 -2.2 -1.9	257 323 122	21 20 20 23	57.9 32.1 02.3 27.2	-1.7 N -0.6 -1.0	-1.4 +0.8 -0.8	293 2 231 88	21 20 20 23	48.0 04.7 03.8 33.0	-1.3 -1.2 N -0.4 -2.5	-2.2 +0.2 -0.5 +1.8	320 250 266 50
	4 4 5 19 22	2307 2310 2457 739 1198	4.1 4.6 6.3 7.4 6.2	2 2 1 1	262 262 274 73 118	0 37.4 0 45.4 2 39.2 17 51.1	-1.2 -1.9	-1.0	268 223 261 25	0 1 2 22	37.4 06.9 49.8 26.3	-1.1 -2.0 -2.2 N +0.1	-2.1 -0.7 -1.9	304 270 297 151	0 1 2 22	14.3 01.6 26.1 21.0	+0.2 -1.6 N -0.4	-4.6 -1.9 -0.3	344 303 342 116
Apr.	30 30 5 16 16	2233 2241 2969 843 851	5.5 5.0 3.2 7.2 6.3	2 2 1 1	228 229 290 60 60	21 31.8 23 20.3 3 38.1 17 52.7	-0.1 -1.4 -1.9 S -0.7	-2.4 -1.5 -1.8	321 287 104 136	23 3 16 18	16.0 52.7 15.4 01.2	N -1.0 -2.6 -2.2 -0.9	-3.1 0.0 +1.0 +0.5	326 79 82 100	16 18	38.7 12.9	N N S -1.4	+1.6	45 67
	18 20 23 23 27	1158 1410 1662 1759 2330	5.2 5.3 6.3 6.5 6.3	1 1 1 2	87 113 142 153 209	20 42.3 20 48.5 0 20.6 18 17.7	-0.4 -1.5 -0.6 -0.8 N	0.0 +0.3 -0.9 -2.3	126 105 144 145	21 0 18	16.6 23.5 15.3	A -0.5 -1.7 N	+0.1	54 111 112	0 18 23	30.1 15.3 25.9	A N -0.5 -2.8	+1.2	77 82 222
	29 29	2498 2498	4.5 4.5	1 2	224 224	2 49.3 4 13.2	-2.2	+0.9 +1.1	81 266	3	19,0	-1.6 s	+1.9	62	3	45.1	-0.6 S	+3.6	32

## Ordinary Occultations 1975

	30 30 30	2638 2639 2773	5.4 6.0 6.1	2 2 2	235 235 246	2	41.9	11 -5°µ 11	+0.6	260	2 3	32.3 03.6	-2.7 11	-0.1	<b>194</b> 279	3 3 22	09.7 08.3 12.8	-2.0	+2.4	230 313 195
fiay	30 1 3 14 17	2774 2787 3045 969 1364	6.3 6.4 6.0 7.1 5.5	2 2 2 1 1	246 247 270 52 81	1 3 17	03.9 17.1 00.3	"-2.0 -1.5 S	+0.9	235 183 73	22 1 16	17.1 29.2 09.1	5 17 17 -2.4	0.0	193 262 107	22 1 16	38.9 31.5 20.0	-1.6 -3.4 S II -3.6	+0.6 -2.0 +1.4	21-3 294 73
	18 20 20 20 21	1489 1717 1726 1727 1853	ย์.8 7.3 6.9 7.1 4.9	1 1 1 1	95 121 123 123 136	19 16 19 20	17.3 59.4 56.2 48.6	-1.5 -1.5 JI -2.0 -2.0	-0.3 -1.6 -0.4 -0.4	113 121 110 110	19 17 20 21	37.6 11.7 24.4 15.4	-2.0 -2.9 G -2.9 -2.6	+1-7 -0.1 +2-7 +2.2	75 84 64 60	20	04.6	11 11 -1.4 11	-2.0	144
June	28 30 1 7 12	2733 3104 3229 415 1198	6.4 6.5 5.6 6.2	2 2 2 2 1	216 249 261 329 37	14 23 14 16	04.1 05.7 36.7 15.7	-1.4 -0.9 11 ~1.0 -0.5	+1.4 +1.7 -1.1 -0.9	258 208 268 146	23 16	24.1	5 -1.5 ∦ ¦3 -1.0	+0.5	239 107	23 1 16	29.1 12.9 31.0	£ -1.9 -0.9 5 -1.5	-0.6 +4.4 +1.5	269 189 71
	12 13 15 16 17	1210 1341 1566 1633 1609	5.9 4.3 6.6 6.3 6.9	1 1 1 1	38 51 73 91 100	17 19	35.9 58.7	-0.3 11 13 5 -0.9	+0.9	96 165	16 16 19	56.3 11.6 58.4	n A -1-0 -2-3 -1-4	-2.2 -0.9 -0.7	155 110 125	18 16 16 20	24.1 47.7 20.0 02.4	k -0.1 -1.9 -4.1 -1.7	-1.0 -1.0 +1.1 +0.5	135 122 73 57
June July	17 18 19 21 25 25 25 13	1815 1949 2063 2217 2826 2826 3045 1662 1901	4.8 5.7 5.5 4.0 4.0 5.3 6 5.3 6 5.3 6 5.3 6	1 1 1 1 2 2 1 1	106 119 131 146 196 316 316 61 88	20 16 1 3	53.7 58.4 52.7 09.0	21 -0.8 -1.9 -0.0 A 11 22	-2,6 +1,1 +2,1	40 145 33 234	21 16 2 3 19 18 17	58.5 56.2 21.7 18.3 <b>30.3</b> 12.7 48.9 42.2	11 -1.3 -1.9 -1.4 -1.4 -0.5 -0.4 -0.5 -1.7	-3.6 -1.3 -1.6 +1.5 +1.7 +0.2 -1.0 -1.7	167 107 143 75 240 239 142 138	21 17 2 3 19 15 17	46.0 00.0 19.0 39.3 43.5 11.7 46.1 40.3	1) -1.2 -3.4 -0.0 -0.9 -0.6 -0.4 -0.4 -5.0 -2.3	-0.0 +0.7 -0.2 +1.0 +1.0 -0.7 +0.1 -0.3	1.00 73 110 259 270 105
	15 15 15 17 18	1907 1913 1914 2181 2307	G.7 7.1 6.8 7.1 4.1	1 1 1 1	80 39 39 115 127	18 20 20 21	57.9 19.2 38.9 58.6	-1.4 -1.0 -0.9	-1.7 -1.0 +2.1	146 141 40 66	19 20 22 19	05.6 25.2 23.6 45.1	-1.5 -0.6 N 0.0 -2.6	-0.1 +0.1 +3.3 -3.1	112 112 41 143	19 20 19	14.3 31.7 40.0	-1.4 -0.6 N E -2.8	+1 +1.0	79 61 107
Aug.	19 19 26 28 1	2338 2445 3371 3501 432	6.6 7.4 6.4 5.3 5.9	1 1 2 2 2	129 138 220 232 277	1 0 1	10.1 22.8 16.3	и ц -1.0	-0.8	17 179 260	21 0 1	46.8 54.9 24.6	H N -0.5 -1.7	+3.3 -0.3	180 188 269	17 22 1 1	15.0 15.4 26.2 19.9	-1.9 -1.4 -1.6 -2.9	-2.3 +2.2 +2.9 -2.2	126 214 209 295
	1 12 13 13 14	433 2002 2119 2136 2267	5.6 6.8 6.7 5.8 5.1	2 1 1 1	277 71 83 35 97	19 20	02.0 40.4	N -1.1 N -0.4 N	+0.5 +2.9	104 50	1 19	03.3 17.8	-0,€ И И	<b>+</b> 1 <b>.</b> 3	175 79	1 19 17 20	28.3 37.6 40.8 22.3	-0.6 +0.2 -2.2 H -2.1	+2.0 +4:0 -3.3	211 38 152 148
	16 17	2556 2715	7.1 6.5	1 1	122 134	20 21	31.6 06.6	-2.8 -1.9	-1.9 +1.7	130 64	20 21	49.1 37.7	-2.3 -1.3	-0.3 +2.3	111 52	20 22	58.1 03.8	-1.8 -0.3	+0.8 +3.4	85 27

## Ordinary Occultations

## **OCCULTATION PREDICTIONS**

				-	El.	CAPE TOWN	JOHANNESBURG	SALISBURY O O
Date	9	No.	Mag.	P.	or Moon	E. 18.500, S. 33.900 U.T. a b P	E. 28.000, S. 26.200 U.T. a b P	E. 31.100, S. 17.700 U.T. a b P
	17 19 19	2718 2969 2969	6.7 3.2 3.2	1 1 2	0 134 156 156	h h h h h h h h 21 52.9 -2.0 +0.8 90 17 52.8 . 135 18 33.2 . 194	h m m m o 22 17.4 -1.4 +1.2 80 17 48.4 -1.9 -1.3 97 19 10.7 -2.2 +1.9 225	h m m m o 22 33.3 -0.9 +1.7 60 17 52.3 -2.2 +0.7 66 19 29.7 -3.0 +0.7 253
Sept.	23 30 30 8 8	3453 646 651 1925 1925	4.9 6.1 5.9 1.2 1.2	2 2 2 1 2	201 269 269 36 36	N A 0 41.2 -0.5 0.0 235 11 27.0 -1.9 -0.9 97 12 39.6 -1.5 -2.0 318	19 48.0 -0.6 +2.1 204 0 00.0 -0.3 +0.8 224 0 48.8 -1.1 +0.1 245 II N	20 03.7 -1.3 +1.1 232 0 06.0 -1.0 +0.2 247 0 52.0 -1.8 -0.4 266 11 11
	11 13 14 14 15	4008 2687 2826 2830 2936	7.7 6.5 4.0 6.9 6.8	1 1 1 1	79 104 115 116 126	18 21.2 -1.8 +0.5 99 22 46.9 0.0 +2.0 53 ii 22 32.5 -0.5 +2.2 51 18 18.3 -2.0 +1.7 47	18 42.9 -1.3 +1.2 81 23 00.5 +0.4 +2.0 43 N 22 51.3 0.0 +2.2 43 16 56.2 -1.7 +3.3 31	18 59.2 -0.7 +2.1 5% A 21 47.0 . 135 23 11.5 +0.7 +3.1 18 19 39.0 . 357
Oct.	16 27 27 8 8	3065 752 755 2302 2302	7.5 4.7 6.3 2.9 2.9	1 2 2 1 2	133 251 251 46 46	0 1 17.3 -2.4 -2.1 303 1 58.4 -1.6 +1.3 219 S S	19 23.1 115 1 25.6 319 2 28.9 -2.7 +1.3 235 9 17 14.2 -0.7 +1.7 247	19 28.7 -3.5 +0.3 87 II 2 46.0 -3.2 +0.5 259 16 15.4 -1.4 +0.8 85 17 25.4 -0.8 +0.5 273
	0 8 9 10 11	2303 2322 2465 2611 2764	5.1 4.3 7.4 6.8 6.3	1 1 1 1	46 47 60 72 84	3 19 06.6 0.0 +2.0 57 19 17.9 -1.1 -0.1 122 17 35.3 -1.5 +2.2 55 S	S A 19 26.7 -0.5 +0.3 109 18 05.3 -0.7 +2.8 41 17 07.9 -3.0 -0.4 105	16 15.7 -1.4 +0.8 84 II 19 31.8 -0.1 +0.7 86 18 38.1 . 4 17 18.9 -2.5 +0.9 80
	12 14 17 24 _ <sup>h</sup>	2913 3166 3501 847 847	5.0 4.8 5.3 3.0 3.0	1 1 1 2	97 120 152 231 231	21 15.6 -2.0 -0.6 126 23 47.5 -0.2 +2.0 53 18 25.0 -1.8 -0.3 75 A A	21 28.0 -1.1 0.0 115 A 18 45.7 -2.3 +0.6 66 A 21 37.9 -1.4 -1.6 289	21 33.3 -0.5 +0.6 92 A 19 03.2 -2.1 +1.9 46 20 45.3 +0.5 +2.5 26 21 24.6 -2.3 -3.4 319
Nov.	28 30 2 2 5	1271 1519 1925 1925 2394	5.9 6.5 1.2 1.2 6.5	2 2 1 2 1	269 297 342 342 26	N 2 50.7 9 31.9 -2.4 +0.1 10 43.3 -1.5 -1.7 321 N	12 53 11 12	0 14.8 . 214 N N N 17 07.1 -0.9 -1.3 135
	5 8 11 13 22	2401 2856 3235 3464 995	5.6 6.6 7.5 7.1 4.1	1 1 1 1	27 65 100 121 216	5 18 04.4 -0.6 +2.9 34 21 43.8 -0.7 +1.9 62 17 52.1 -1.6 +1.7 40 2 36.1 -1.9 0.0 108	17 31.1 +0.3 +1.8 53 18 30.1 0.0 +3.1 25 22 02.4 -0.2 +1.9 52 18 23.8 -1.7 +2.4 36 S	A II 18 54.0 -1.1 +3.7 17 N
Dec.	25 7 14 20	1364 3054 363 1198	6.5 6.4 7.3 6.2	2 1 1 2	252 56 136 209	0 50.0 -1.6 -0.9 266 S 22 54.0 -1.6 +2.2 50 23 52.0 -2.3 -0.4 270	0 59.4 -2.0 -1.3 289 S 23 32.1 15 24 09.7 -2.3 -1.0 297	0 51.6 -2.0 -2.1 314 16 56.8 -0.2 +3.5 17 N 24 02.3 -1.9 -2.4 326

## Ordinary Occultations

21	1210	5.9	2	209	1 41.6 -1.6 -1.0 313	1 34.3	G	1	
22	1341	4.3	1	223	14	13	2	2 49.7 .	. 190
:6	1925	1.2	1	288	21	24	2	3 44.9 +0.	1 -2.2 150
27	1925	1.2	2	233	31	1	(	33.7 -1.0	) -0.2 250
29	2192	6.2	2	315	71	G.		1 09.9 .	. 200

## INDEX TO OCCULTED STARS BRIGHTER THAN FIFTH MAGNITUDE

Z.C.	Name	Mag.	Sp	Z.C.	Name	Mag.	Sp
709	τ Taur	4,3	B5	2302*	$\beta$ Scor	2,9	B1
752	، Taur	4,7	A5	2307	w Scor	4,1	B2
847	ζ Taur	3,0	B3p	2310	ω <sup>1</sup> Scor	4,6	G0
995*	v Gemi	4,1	B5	2498	ξ Ophi	4,5	F5
1341*	<sup>a</sup> Canc	4,3	A3	2633	# Sgtr	4,0	B8p
1815	X Virg	4,8	КО	2826	ρ Sgtr	4,0	A5
1853	<b>∉</b> Virg	4,9	M3	2969*	β Capr	3,3	G0, A0
1925	Spica	1,2	B2	3453	« Pisc	4,9	A2p

#### NOTES

It is interesting to note that Spica is included in the list of occulted stars. Unfortunately of the three occultations of this star that occur in 1975, two are seen only from Cape Town in the daytime, the other only from Salisbury in the early hours of December 27.

"Z.C. 4008" is the planet Neplune which is occulted during the evening of September 11. It should be interesting to observe as, unlike a star, its disappearance does not seem instantaneous.

# **GRAZING OCCULTATIONS**

When a star moves tangentially to the limb of the Moon, and is occulted for a very short period only - a few minutes, or even seconds - a grazing occultation is said to occur. Because the limb, as seen from the Earth, is in fact the outline of numerous mountains and valleys, there may be several disappearances and reappearances, which are not only fascinating to observe, but which may be accurately timed to yield valuable data on the relative positions of star and Moon, in both right ascension and declination, as well as on the shape of the Moon. Some of these data cannot readily be obtained in any other way.

The maps on the following pages have been prepared by H.M. Nautical Almanac Office to show the tracks of stars brighter than 7.5 magnitude which will graze the limb of the Moon when it is at a favourable elongation from the Sun and at least  $10^{\circ}$  above the observer's horizon ( $2^{\circ}$  in the case of bright stars). Each track starts in the West at some arbitrary time given in the key and ends beyong the area of interest, except where the letters "A", "B", or "S" are given. "A" denotes that the Moon is at a low altitude, "B" that the bright limb interferes, and "S" that sunlight interferes. The tick marks along the tracks denote 5 minute intervals of time which, when added to the time at the beginning of the track, give the approximate time of the graze at places along the tracks.

The tracks as shown on the maps are approximate only. Since the observer's location is very critical, successful observations call for very accurate predictions. With the aid of the IBM computer of the CSIR at Pretoria such predictions are at present prepared at 6-monthly intervals for a number of centres in South Africa, Rhodesia and Malawi. By plotting the predicted graze track on a reliable survey map (e.g. the South African 1:50 000 series) it is usually possible to select a convenient site from where the graze may be observed. Ideally a team of observers would be stationed at intervals along a line running at right angles to the graze track - say, along a main road - each with his own telescope and timing equipment. Each observer will see a different sequence of events, the combined results forming an accurate picture of the limb of the Moon.

The equipment needed is similar to that used for ordinary (or 'total') occultations, but must, of course, be portable. A 75 mm refractor is ideal for average events, but better instruments with a larger aperture have often shown their superiority under difficult conditions. Timing is best carried out with a portable tape recorder and radio receiver tuned to ZUO or other time signal station.

It will be seen from the maps that many grazing occultations occur in regions which are rather far removed from the main cities, and which cannot easily be reached by teams of observers from one of the ASSA centres. It is worth remembering, however, that a team of many observers, while ideal, is by no means essential; that a single good observer is worth more than many unsuccessful ones, and that one good observation is worth infinitely more than no observations at all.

Observers in other parts of southern Africa – especially the more distant regions – who may be interested, are therefore invited to contact the coordinator for grazing occultations: Mr. J. Hers, 48, Central Road, Linden Extension, Randburg 2001, so that they may be informed of all favourable grazes occurring within their neighbourhood.

## JANUARY 10 TO APRIL 16



No.	2.C.	Mag.	Da	le	Begin	ning	Sunli	t Lim	it	Z.Ç.	Mag.	Date	9	Begi	nning	SunIn	Limit
					L		%		/					1	ı m	%	
1	2529	6.6	Jan	10	4 <sup>n</sup> 3	35	5	S	11	485	6.9	Feb	18	20	35	45	N
2*	411	7.3	Jan	21	22 1	14	62	N	12	487	5.2	Feb	18	21	39	46	N
3	525	6.4	Jan	22	19 3	33	72	S	13	2171	6.8	Маг	3	3	23	67	S
4*	709	4.3	Jan	23	22 2	22	82	N	14	2327	6.7	Mar	4	6	6	55	S
5	1705	7.5	Jan	31	2	1	83	s	15	3002	6.2	Mar	9	5	43	11	N
6	1809	6.9	Jan	31	23	1	74	S	16	739	7.4	Маг	19	20	3	40	N
7*	1949	5.8	Feb	2	0 1	19	63	s	17	2230	6.8	Mar	30	22	50	83	S
8	2207	7.0	Feb	4	1 3	37	41	s	18	534	6.0	Apr	14	19	2	10	N
9*	2214	6.2	$\mathbf{Feb}$	4	4	7	40	S	19*	843	7.2	Apr	16	19	2	25	N
10	2794	6.7	Feb	8	5 2	22	7	Ν									

Explanation of map - see text on page 48 Explanation of column headings of table - see page 53 An asterisk denotes the star is double - see page 53

## APRIL 18 TO JUNE 12

30

1364

6.5

May 17

18 53

42

![](_page_54_Figure_2.jpeg)

Explanation of map - see text on page 48 Explanation of column headings of table - see page 53 An asterisk denotes the star is double - see page 53

N 41

1210

5.9

Jun

12

20 3

11

N

## JUNE 16 TO SEPTEMBER 1

![](_page_55_Figure_2.jpeg)

Explanation of map - see text on page 48 Explanation of column headings of table - see page 53 An asterisk denotes the star is double - see page 53

![](_page_56_Figure_1.jpeg)

![](_page_56_Figure_2.jpeg)

No.	Z.C.	Mag.	Dat	te	Begi	inning	g Sunlit	Li	mit	Z.C.	Mag.	Dat	e	Begi	nning	Sunli	Limit
					ł	n m	%		/					h	m	%	
64	1114	6.8	Sep	2	6	7	19	S	76	2394	6.5	Nov	5	19	18	5	$\mathbf{s}$
65*	1925	1.2	Sep	8	13	25	10	Ν	77	2972	6.7	Nov	9	18	51	38	s
66	2233	5.5	Sep	10	23	21	31	S	78	2975	7.0	Nov	9	19	16	38	S
67	2826	4.0	Sep	14	23	27	72	$\mathbf{S}$	79	3104	6.5	Nov	10	19	46	48	s
68	467	6.7	Sep	25	0	56	84	N	80	3111	6.8	Nov	10	22	7	49	S
69	752	4.7	Sep	27	2	37	66	Ν	81	29	7.2	Nov	14	19	24	84	s
70	1060	7.4	Sep	29	4	36	45	$\mathbf{S}$	82	2927	7.2	Dec	6	19	23	14	S
71*	2629	6.3	Oct	10	23	4	36	S	83*	230	7.4	Dec	13	22	1	78	S
72	2913	5.0	Oct	12	23	32	56	S	84*	1925	1.2	Dec	27	2	9	34	S
73*	847	3.0	Oct	24	22	57	81	Ν	85	2192	6.2	Dec	29	3	5	14	s
74	1025	7.4	Oct	26	3	5	71	$\mathbf{S}$	86	2353	4.6	Dec	30	4	53	7	S
75	1271	5.9	Oct	28	2	6	50	S									

Explanation of map - see text on page 48Explanation of column headings of table - see page 53 An asterisk denotes the star is double - see page 53

## EXPLANATION OF COLUMN HEADINGS IN TABLES

- No. the number of the track on the map. An asterisk denotes that the same is double notes are given below.
- Z.C. the number of the star in the Zodiacal Catalogue.
- Date

Beginning-an arbitrary time of the beginning of the track in the west.

- Sunlit the percentage of the Moon's disk lit by the Sun
- Limit N = northern limit (a complete occultation takes place south of the track)
  - S = southern limit (complete occultation north of the track)

## NOTES ON DOUBLE STARS

TRACK NO.	ZC	
2	411	is the brighter component of the double star Aitken 2122. The companion
		is 8.2 magnitude; separation 3."6 in p.a. 309 <sup>0</sup> .
4	709	is a spectroscopic binary star
7,45	1949	is the mean of the double star Aitken 8954. The components are 6.5 and
		6.9 magnitude; separation 0.6.
9	2214	is the brighter component of the double star Aitken 9681. The companion
		is 8.5 magnitude; separation 11.11 in p.a. 281°.
19	843	is the brighter component of the double star Aitken 4200. The companion
		is 7.8 magnitude; separation 3!'8 in p.a. 269 <sup>0</sup> .
21,49	1410	is the brighter componnet of the double star Aitken 7416. The companion
		is 8.4 magnitude; separation 37" in p.a. 75 <sup>0</sup> .
24	2773	is the brighter component of the double star Aitken 11972. The companion
		is 9.5 magnitude; separation 7."5 in p.a. 88 <sup>0</sup> .
29,63	969	is the brighter component of the multiple system Aitken 4962. The four
		companions, with wide separations, are 9th, 11th, 8th and 10th magnitudes
34	2871	is the brighter component of the double star Aitken 12728. The companion
		is 8.5 magnitude; separation 10" in p.a. 234 <sup>0</sup> .
35	2969	is a spectroscopic binary star.
42	1688	is the brighter component of the double star Aitken 8247. The companion
		is 10th magnitude; separation 5."2 in p. a. 279 <sup>0</sup> .
47	3524	is the mean of the double star Aitken 17111. The components are 7.5 and
		8.0 magnitude; separation 0."4 in p.a. 220 <sup>0</sup> .
59	2267	is a spectroscopic binary star.
65,84	1925	is a spectroscopic binary star.
71	2629	is the mean of the double star Aitken 11127. The components are 6.9 and
		7.3 magnitude; separation 1.10 in p.a. 2040.
73	847	is a spectroscopic binary star.
83	230	is the brightest component of the system Aitken 1238. The brightest
		companion is 8.4 magnitude; separation 16" in p.a. 79 <sup>0</sup> .

## **OCCULTATIONS DURING LUNAR ECLIPSE**

During the lunar eclipse on Nov. 18/19, it will be possible to observe grazes of two stars fainter than the normal limit. These are:

No.	Name	Mag.	Beginning	Limit
81a	SAO 093498	8,8	23 <sup>n</sup> 38 <sup>m</sup>	N
81b	SAO 093518	9,3	0 53	S

Tracks are shown on map 4.

# TIME SYSTEMS AND TELESCOPE SETTING

This section is intended to serve established amateurs and professional astronomers -i.e.those having some knowledge of time and coordinate systems. Space in the booklet does not permit full explanation, which in any case would appear complicated to the layman.

## TIME SIGNALS FROM RADIO STATION ZUO

Radio signals of mean solar time are generated by the Precise Physical Measurements Division of the National Physical Research Laboratory in Pretoria and broadcast by the Post Office transmitting station at Olifantsfontein

Carrier Frequency	Radiated Power	Time of Transmission
2,5 MHz	4 Kw	2000 - 0600 SAST
5 MHz	4 Kw	Continuous
100 MHz	80 w	Continuous

The signals consist of one pulse per second, each pulse consisting of 5 cycles of 1000 Hz tone. The first pulse in every minute is lengthened to 500 milliseconds. Morse code announcements are made during the minute preceding every fifth minute. They consist of the call sign ZUO (repeated 3 times) and the Universal Time (formally known as Greenwich Mean Time) at the next minute. (A special coding indicating UT1 minus UTC is also indicated in the first 15 seconds of the minute by slightly lengthened second pulses)

## SOUTH AFRICAN STANDARD TIME

South African Standard Time (as in everyday use) is mean solar time for the 30° East meridian (which runs east of Johannesburg and just west of Durban) and is exactly 2 hours ahead of Universal Time.

## TIME OF SUN'S TRANSIT OVER THE 30° MERIDIAN

The table below gives the SAST when the Sun transits the  $30^{\circ}$  meridian - and a sundial on that meridian reads noon.

Jan	1	12 <sup>h</sup>	03 <sup>m</sup>	22 <sup>8</sup>	May	11	11 <sup>h</sup>	56 <sup>m</sup>	21 <sup>8</sup>	Sep	18	11 <sup>h</sup>	54m	21 <sup>8</sup>
	11	12	07	46		21	11	56	29	· ·	28	11	50	52
	21	12	11	13		31	11	57	31	Oct	8	11	47	43
	31	12	13	27	<b>J</b> un	10	11	59	13	1	18	11	45	18
Feb	10	12	14	18	Í	20	12	01	19		28	11	43	52
	20	12	13	52		30	12	03	27	Nov	7	11	43	40
Mar	2	12	12	21	Jul	10	12	05	11		17	11	44	52
	12	12	10	00		20	12	06	14	1	27	11	47	27
	22	12	07	80		30	12	06	24	Dec	7	11	51	12
Apr	1	12	04	07	Aug	9	12	05	32		17	11	55	50
	11	12	01	15		19	12	03	43	1	27	12	00	48
	21	11	58	50	{	29	12	01	05					
May	1	11	57	10	Sep	8	11	67	52					

## RISING AND SETTING TIMES AT 30°S ON THE 30° MERIDIAN

The scale at the bottom of pages 24 and 25 is for finding rising or setting times of any object who right ascension and declination are known. Set dividers or a strip of paper from the index at the centre of the scale to the object's declination and in the direction desired for either rising or

### Time Systems

setting. Measure this same distance and direction along the midnight line, beginning at the object's right ascension indicated by the numerals. (Should this end point fall outside the chart, 12 hours should be added to or subtracted from the right ascension. Reset the dividers using the end of the scale instead of the centre index, and measuring in the opposite direction to that first used.) Through the point established draw a line parallel to the March Equinox transit line (indicated by the dashed line on the chart). This line will show the time of rising or setting of the object at position  $30^{\circ}$ E and  $30^{\circ}$ S.

### SIDEREAL TIME ON THE 30° MERIDIAN

Side real Time is given by the line of Right Ascension coinciding with the meridian.

		At I SAS	0 hrs T	At 2 SAS	1 hrs T			At C SAS	) hra T	At 2 SAS	1 hrs T			At ( SAS	) hrs T	At 2 SAS	21 hrs T
Jan	1	6h	40 <sup>m</sup>	3h	43 <sup>m</sup>	May	11	15 <sup>h</sup>	12 <sup>m</sup>	12 <sup>h</sup>	16 <sup>m</sup>	Sep	18	23 <sup>h</sup>	45 <sup>m</sup>	20h	48 <sup>m</sup>
	11	7	19	4	23		21	15	52	12	55	1	28	0	24	21	28
	21	7	59	5	02		31	16	31	13	35	Oct	8	1	04	22	07
	31	8	38	5	41	ป บบ	10	17	10	14	14		18	1	43	22	46
Feb	10	9	17	6	21	1	20	17	50	14	53		28	2	22	23	26
	20	9	57	7	00		30	18	29	15	33	Nov	7	3	02	0	05
Mar	2	10	36	7	40	Jul	10	19	09	16	12	1	17	3	41	0	45
	12	11	16	8	19		20	19	48	16	52	1	27	4	21	1	24
	22	11	55	8	59		30	20	28	17	31	Dec	7	5	00	2	04
Apr	1	12	34	9	38	Aug	9	21	07	18	10	1	17	5	40	2	43
	11	13	14	10	17		19	21	46	18	50		27	6	19	3	22
	21	13	53	10	57		29	22	26	19	29						
Mav	1	14	33	11	36	Sen	8	23	05	20	08	1					

## **CORRECTION FOR PLACES NOT ON THE 30º MERIDIAN**

Approximate longitude corrections from the 30° East Meridian are provided below.

To find time of Sun's transit over local meridian, apply the longitude corrections to the data in the table above.

To find the sidereal times at SAST 0 hrs and SAST 21 hrs apply the corrections with the sign reversed to the data in the table.

Bloemfontein	+15 <sup>m</sup>	East London	- 8 <sup>m</sup>	Port Elizabeth	+18 <sup>m</sup>
Bulawayo	+ 6	Grahamstown	+14 <sup>m</sup>	Pretoria	+ 7
Cape Town	+46 <sup>m</sup>	Johannesburg	+ 8 <sup>m</sup>	Salisbury	- 4 m
Durban	- 4 <sup>m</sup>	Kimberley	+21 <sup>m</sup>	Windhock	+52

#### **TELESCOPE SETTING**

When a telescope equipped with setting circles is aimed on the meridian, its R.A. circle should read the sidereal time. Thus one can calculate the sidereal time and then set the circle, but is is usually simpler to aim the telescope at one of the well known stars given below and then to adjust the R.A. circle.

### A LIST OF BRIGHT STARS FOR CHECKING TELESCOPE CIRCLES

Star	R.A.	Dec.	Mag.	Sp.	Star	R. A.	Dec.	Mag.	Sp.
Achernar	1 <sup>h</sup> 36,8	-57 <sup>0</sup> 21'	0,5	B5	Procyon	7h38,0	+15 <sup>0</sup> 17'	0,4	F5
Aldebaran	4 34,5	+16 28	0,9	K5	Regulus	10 07,1	+12 06	1,4	B7
Rigel	5 13,3	- 8 14	0,1	<b>B</b> 8	Spica	13 23,8	-11 02	0,9	<b>B1</b>
Betelgeuse	5 53,8	+ 7 24	0,4	M2	Arcturus	14 14,5	+19 19	-0,1	K2
Canopus	6 23,4	-52 41	0,7	$\mathbf{F0}$	Antares	16 27,9	-26 33	0,9	M1
Sirius	6 44,0	-16 41	-1,4	A1	Altair	19 49,5	+8 48	0,8	<b>A</b> 7

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1941-42	H. Knox Shaw	1967-68	J. Hers
1942-43	A.F.I. Forbes	1968-69	J.C. Bennett
1943-44	W.H. van den Bos	1969-70	J. Churms
1944-45	A.W.J. Cousins	1970-71	W.C. Bentley
1945-46	R.H. Stoy	1971-72	A.H. Jarrett
1946-47	W.P. Hirst	1972-73	K.J. Sterling
1947-48	J. Jackson	1973-74	G.A. Harding

## HONORARY SECRETARIES

1922	H.W. Schonegevel	1930	S. Skewes
1922	T. Mackenzie	1931	H. Horrocks
1923	C.L. O'Brien Dutten	1934	H.W. Schonegevel
1923	H.E. Houghton	1935	A. Menzies
	1965	T.W. Rus	SO

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## **GILL MEDALLISTS**

1956	H. Knox Shaw	1963	A.W.J. Cousins
1957	W.P. Hirst	1965	R.H. Stoy
1958	J. Jackson	1967	W.S. Finsen
1960	W.H. van den Bos	1970	J.C. Bennett

## JULIAN DATE AT 1400 HOURS

The Julian Calendar is generally used for recording the time of variable star observations. It numbers the days consecutively from the beginning of the Julian era in 4713 B.C. The Julian day begins at Greenwich mean noon, that is, at 1400 hrs. (2 p.m.) SAST. The Julian date on 1975 Jan. 1, at 1400 hrs. is 2442 414,0 - the first digits are not repeated for each entry in the table below.

	Jan	Feb	Mar	Apr	May	Jun	Jul	Aug	Sep	Oct	Nov	Dec
	2442	2442	2442	2442	2442	2442	2442	2442	2442	2442	2442	2442
1	414.0	445.0	473.0	504.0	534.0	565,0	595.0	626,0	657,0	687,0	718,0	748,0
2	415.0	446.0	474.0	505.0	535.0	566,0	596.0	627.0	658,0	688,0	719,0	749,0
3	416.0	447.0	475.0	506.0	536.0	567.0	597.0	628,0	659,0	689,0	720,0	750,0
4	417.0	448.0	476.0	507.0	537.0	568.0	598.0	629,0	660,0	690,0	721,0	751,0
5	418.0	449.0	477.0	508.0	538.0	569,0	599.0	630,0	661.0	691,0	722,0	752,0
6	419.0	450.0	478.0	509.0	539.0	570.0	600.0	631,0	662,0	692,0	723,0	753,0
7	420.0	451.0	479.0	510.0	540.0	571,0	601.0	632,0	663,0	693,0	724,0	754,0
8	421.0	452.0	480.0	511.0	541.0	572,0	602.0	633,0	664,0	694,0	725,0	755,0
9	422.0	453.0	481.0	512,0	542,0	573,0	603,0	634,0	665,0	695,0	726,0	756,0
10	423.0	454.0	482.0	513,0	543,0	574,0	604,0	635,0	666,0	696,0	727,0	757,0
11	424.0	455.0	483,0	514,0	544,0	575,0	605,0	636,0	667,0	697,0	728,0	758,0
12	425.0	456.0	484.0	515,0	545,0	576,0	606,0	637,0	668,0	698,0	729,0	759,0
13	426.0	457.0	485,0	516,0	546,0	577,0	607,0	638,0	669,0	699,0	730,0	760,0
14	427,0	458,0	486,0	517,0	547,0	578,0	608,0	639,0	670,0	700,0	73,0	761,0
15	428.0	459,0	487,0	518,0	548,0	579,0	609,0	640,0	671,0	701,0	732,0	762,0
16	429.0	460.0	488.0	519,0	549,0	580,0	610,0	641,0	672,0	702,0	733,0	763,0
17	430,0	461,0	489,0	520,0	550,0	581,0	611,0	642,0	673,0	703,0	734,0	764,0
18	431,0	462,0	490,0	521,0	551,0	582,0	612,0	643,0	674,0	704,0	735,0	765,0
19	432,0	463,0	491,0	522,0	552,0	583,0	613,0	644,0	675,0	705,0	736,0	766,0
20	433,0	464,0	492,0	523,0	553,0	584,0	614,0	645,0	676,0	706,0	737,0	767,0
21	434,0	465,0	493,0	524,0	554,0	585,0	615,0	646,0	677,0	707,0	738,0	768,0
22	435,0	466,0	494,0	525,0	555,0	586,0	616,0	647,0	678,0	708,0	739,0	769,0
23	436,0	467,0	495,0	526,0	556,0	587,0	617,0	648,0	679,0	709,0	740,0	770,0
24	437,0	468,0	496,0	527,0	557,0	588,0	618,0	649,0	680,0	710,0	741,0	771,
25	438,0	469,0	497,0	528,0	558,0	589,0	619,0	650,0	681,0	711,0	742,0	772,0
26	439,0	470,0	498,0	29,0	559,0	590,0	620,0	651,0	682,0	712,0	743,0	773,0
27	440,0	471,0	499,0	530,0	560,0	591,0	621,0	652,0	683,0	713,0	744,0	774,0
28	441,0	472,0	500,0	531,0	561,0	592,0	622,0	653,0	684,0	714,0	745,0	775,0
29	442,0		501,0	532,0	562,0	593,0	623,0	654,0	685,0	715,0	746,0	776,0
30	443,0		502,0	533,0	563,0	594,0	624,0	655,0	686,0	716,0	747,0	777,0
31	444,0		503,0		564,0		625,0	656,0		717,0		778,0

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